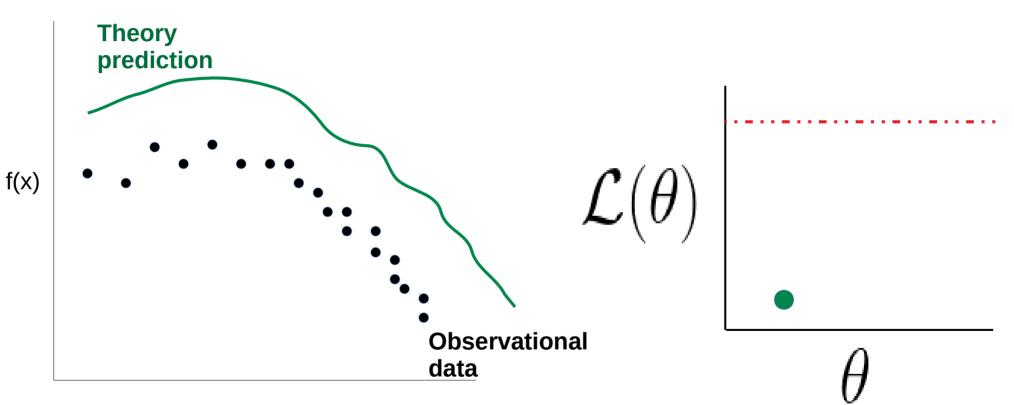
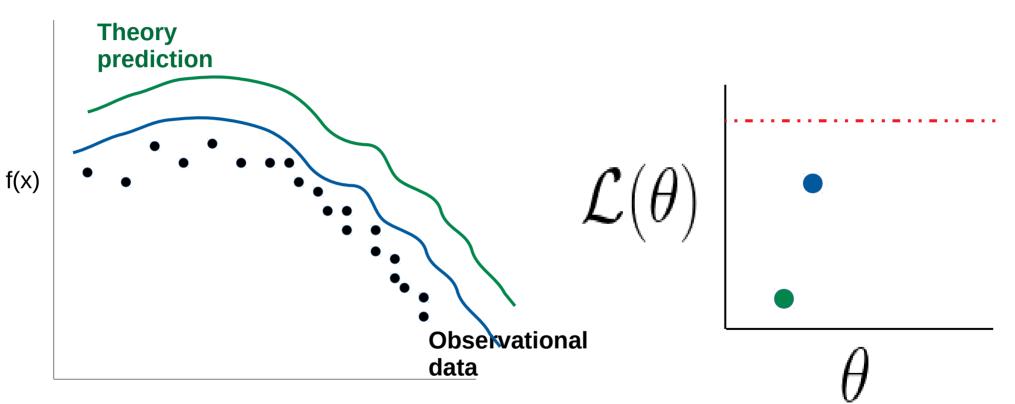
erc Uncertain theory predictions and Gaussian emulation of likelihoods

Marcos Pellejero Ibáñez, Giovanni Arico, Raul Angulo, Matteo Zennaro, Sergio Contreras... Donostia International Physics Center (San Sebastián)

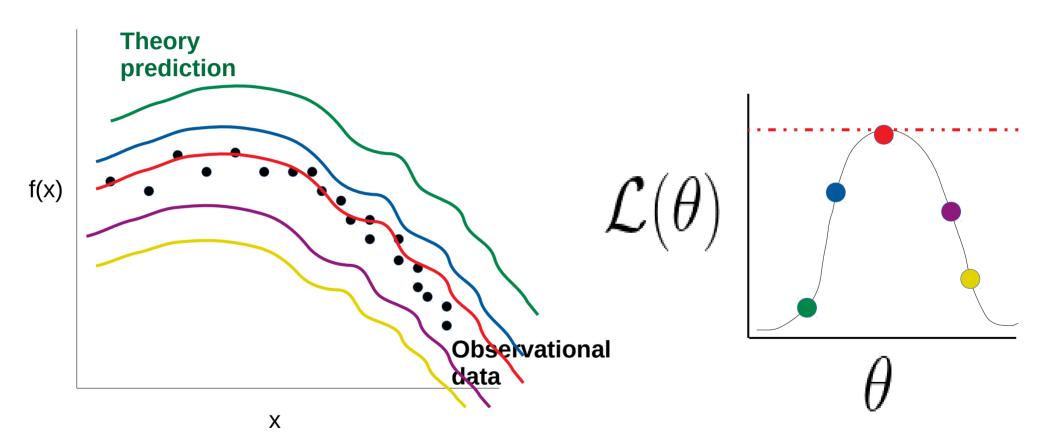
Big Picture • Go from: • To: Likelihood function: Theoretical model Credit Sloan Digital Sky Survey and its **Baryon Oscillation Spectroscopic Survey**

ln 10¹⁰A.





Х



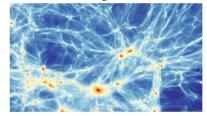
Problems

1) Expensive evaluations of the likelihood (advance modeling).

Complicated loop evaluations, high-n integrals in PT, modeling observational effects



Direct N-body simulations



2) Expensive posterior inferences. Very long chains to achieve convergence, high dimensionality.

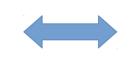
3) Noisy evaluations. Cosmic variance from simulations, uncertainties in the group finder or in the gravity solver.

Traditional MCMC not adequate

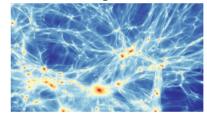
Problems

1) Expensive evaluations of the likelihood (advance modeling).

Complicated loop evaluations, high-n integrals in PT, modeling observational effects



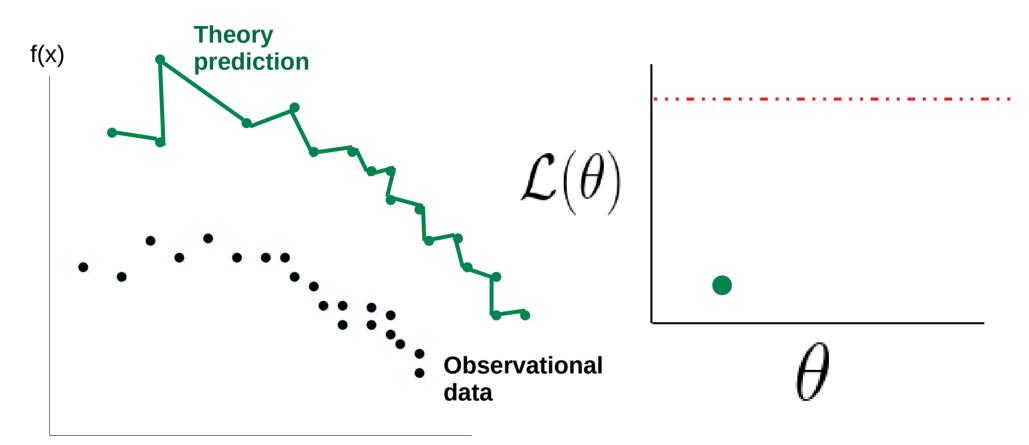
Direct N-body simulations

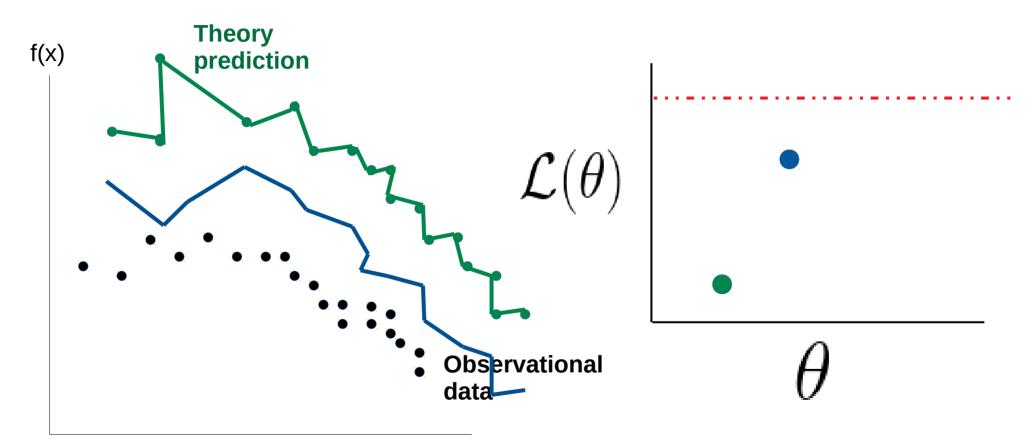


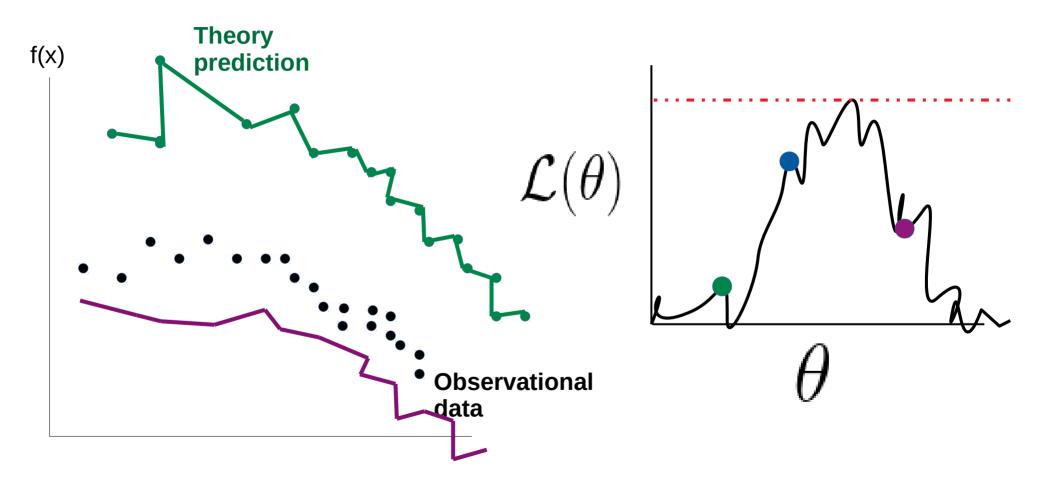
2) Expensive posterior inferences. Very long chains to achieve convergence, high dimensionality.

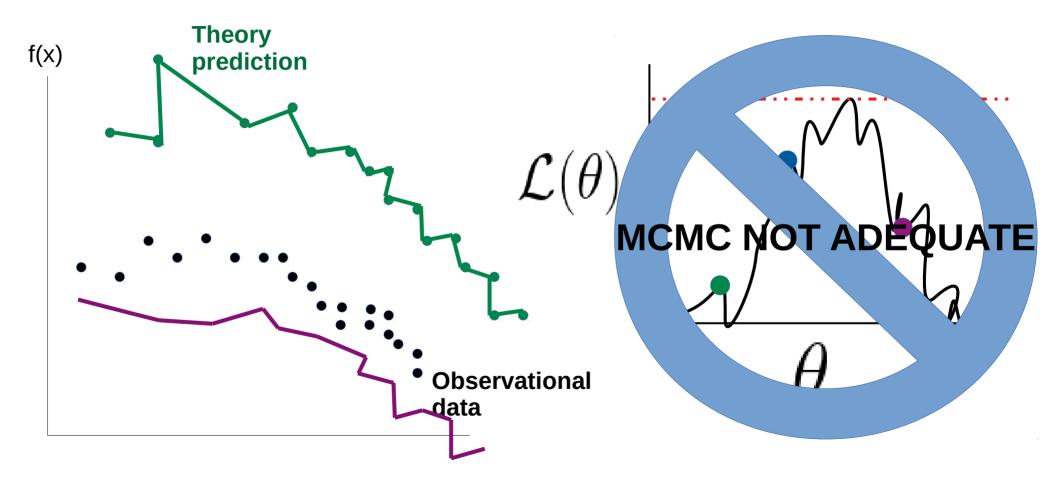
3) Noisy evaluations. Cosmic variance from simulations, uncertainties in the group finder or in the gravity solver.

Traditional MCMC not adequate

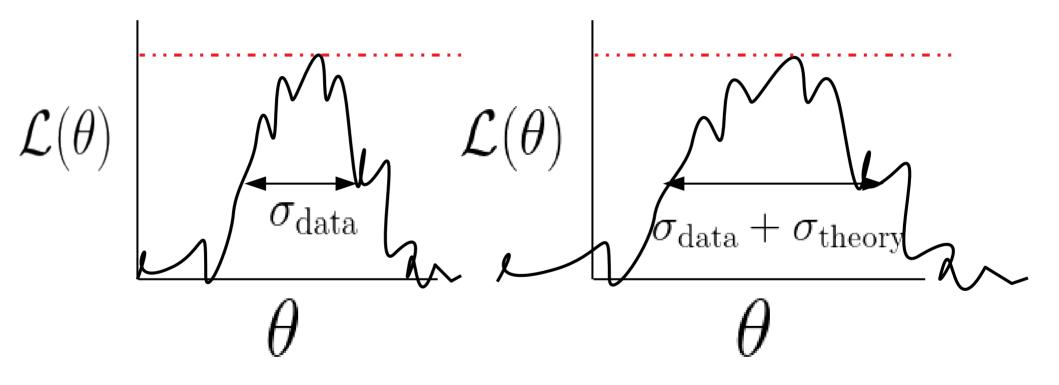


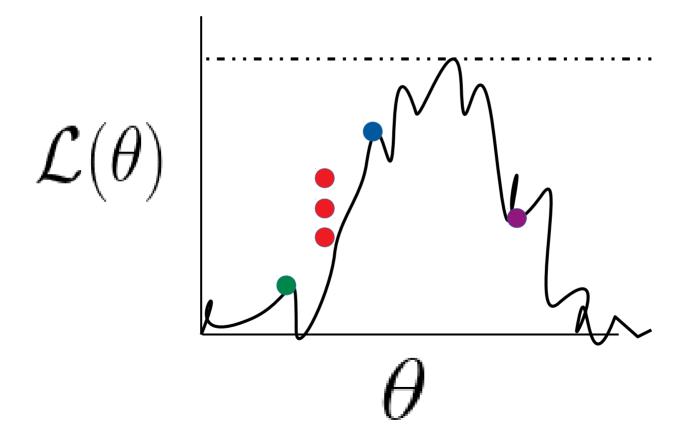


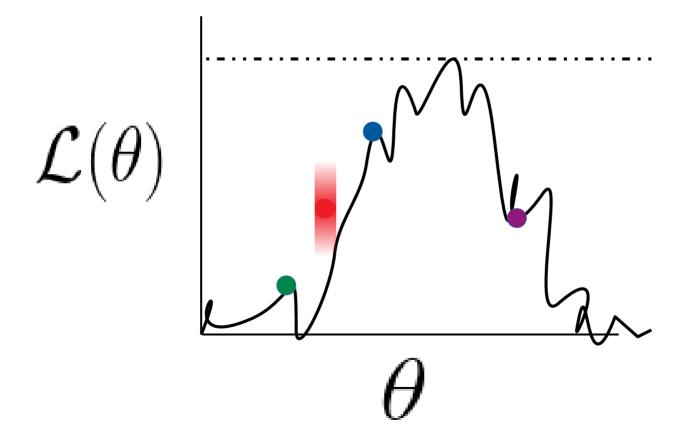


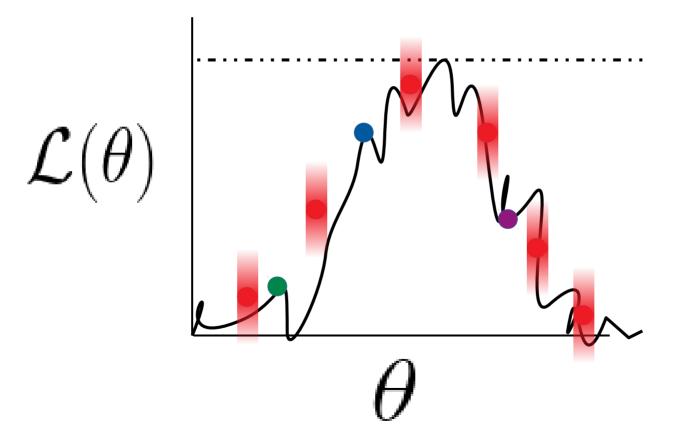


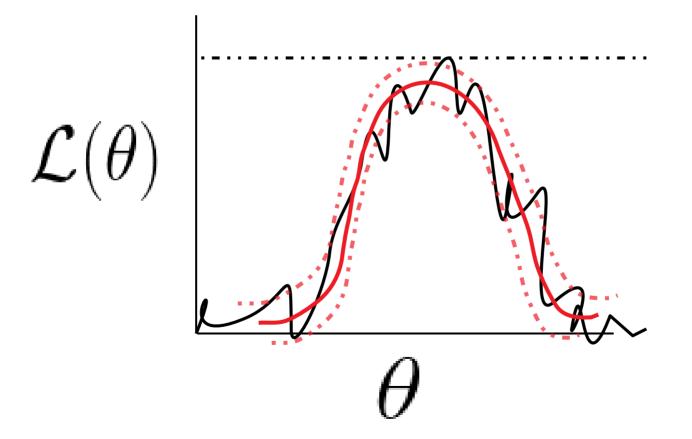
1) Include theory error



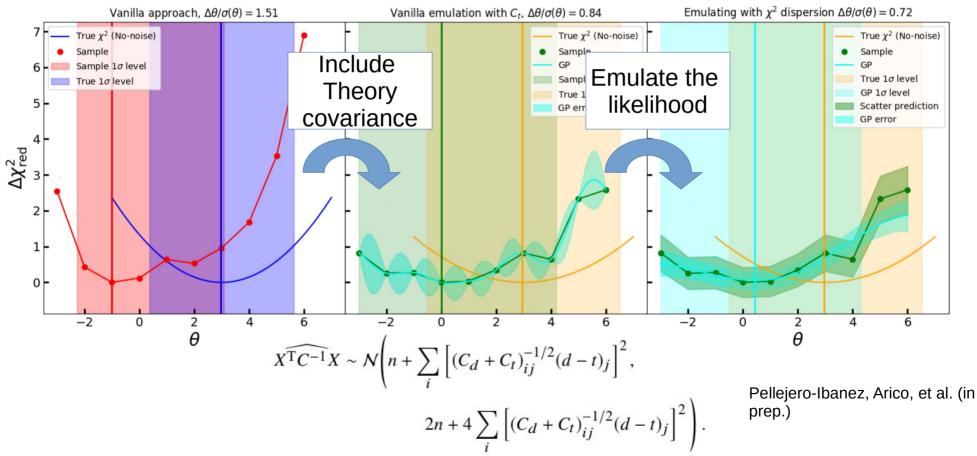








Noise on the likelihood function: toy model and Gaussian Process



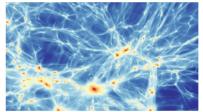
Problems

function evaluations to recover 1) Expensive evaluations of the likelihood (advance modeling).

Complicated loop evaluations, high-n integrals in PT, modeling observational effects



Direct N-body simulations



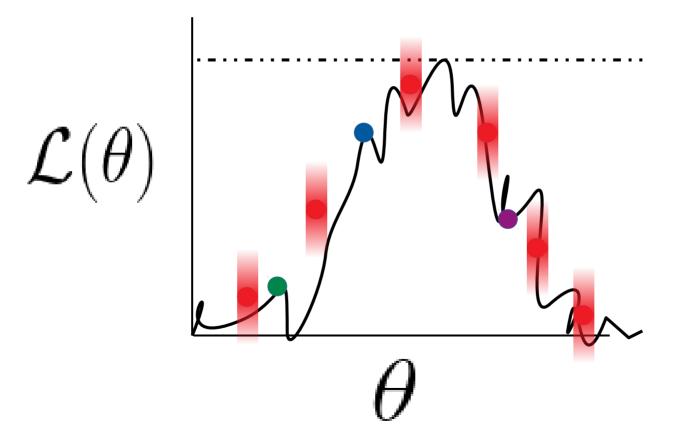
2) Expensive posterior inferences. Very long chains to achieve convergence, high dimensionality.

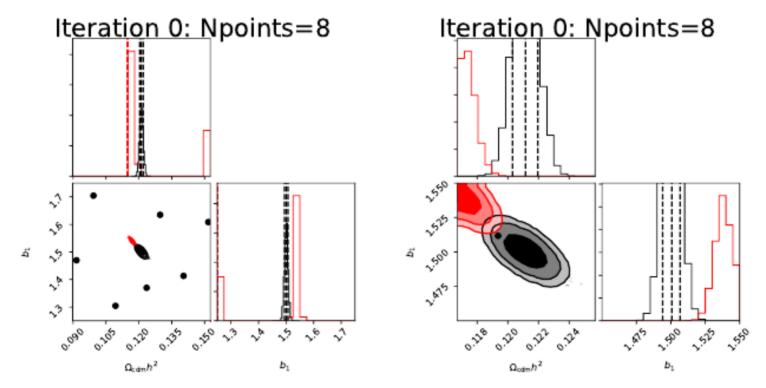
3) Noisy evaluations. Cosmic variance from simulations, uncertainties in the group finder or in the gravity solver.

Traditional MCMC not adequate

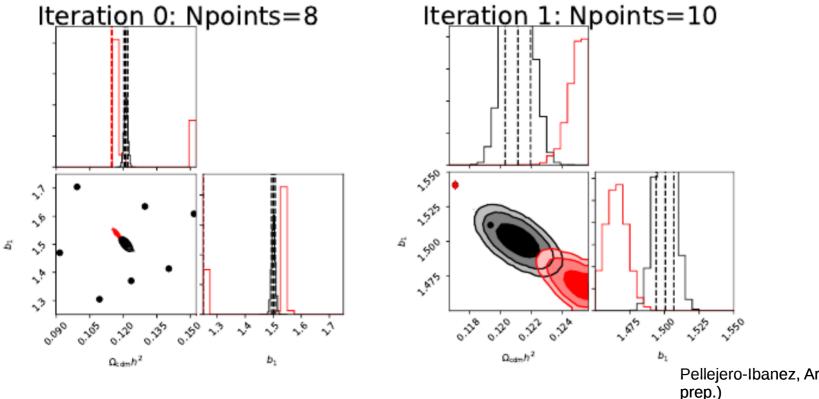
Minimal amount of likelihood

the true one?

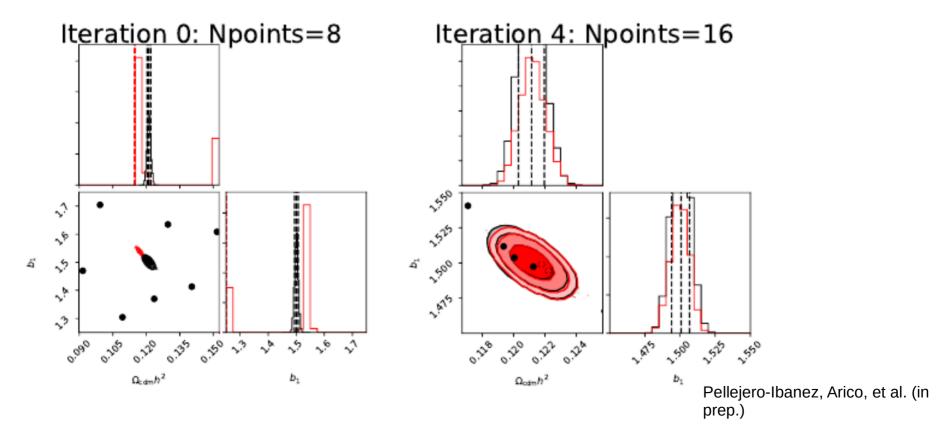


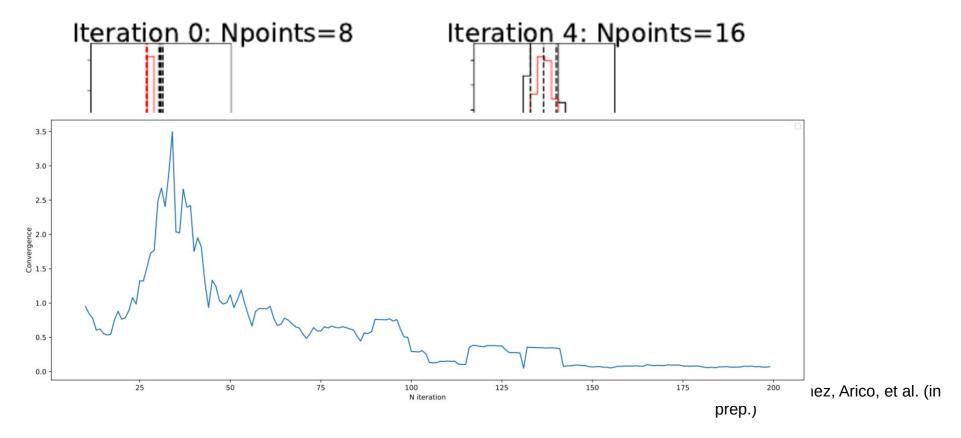


Pellejero-Ibanez, Arico, et al. (in prep.)

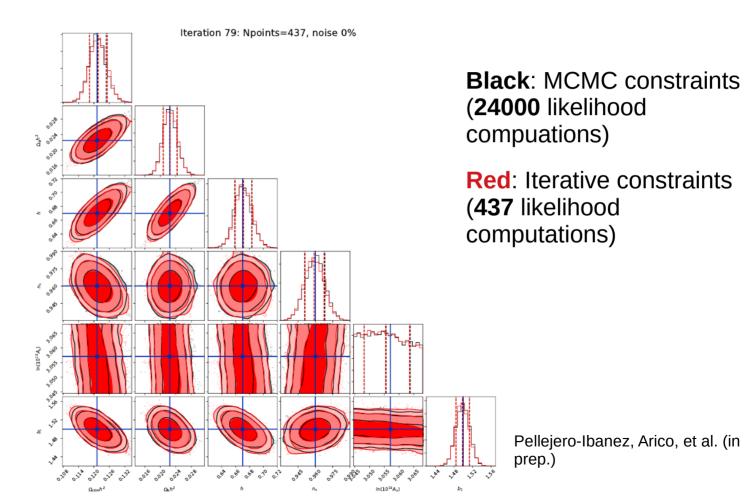


Pellejero-Ibanez, Arico, et al. (in





In higher dimensions: clustering

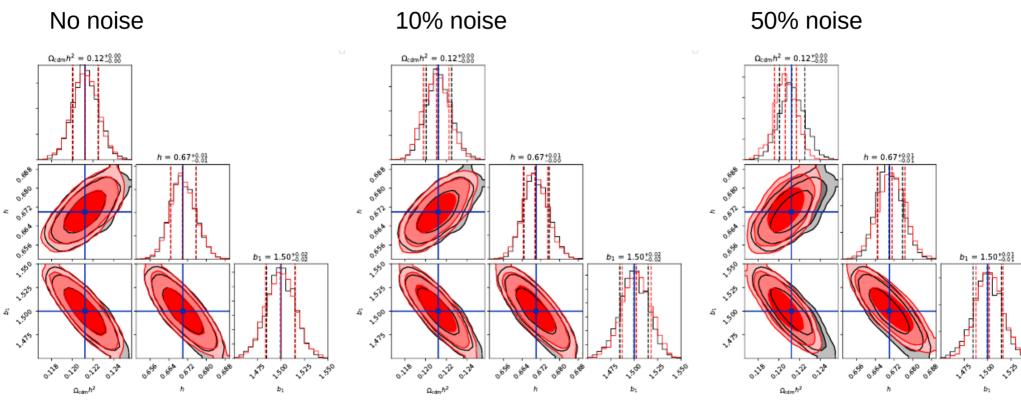


Summary

- Noisy likelihood: Include theory errors and use Gaussian Process to recover "true" results.
- Use iterative **Bayesian Optimization process** to get minimal computations of the likelihood function.

¡Muchas gracias!

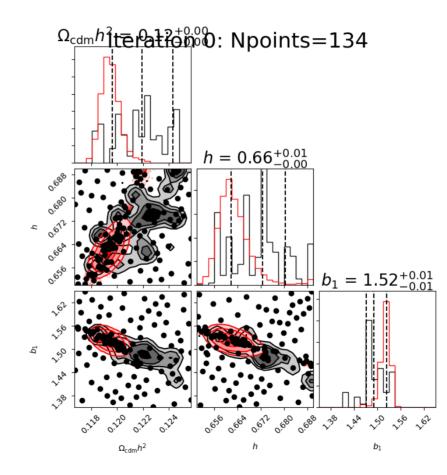
Appendix: Combining ideas

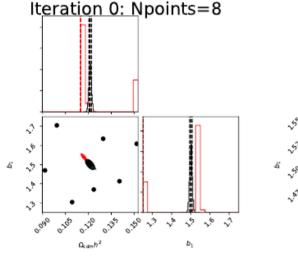


Pellejero-Ibanez, Arico, et al. (in prep.)

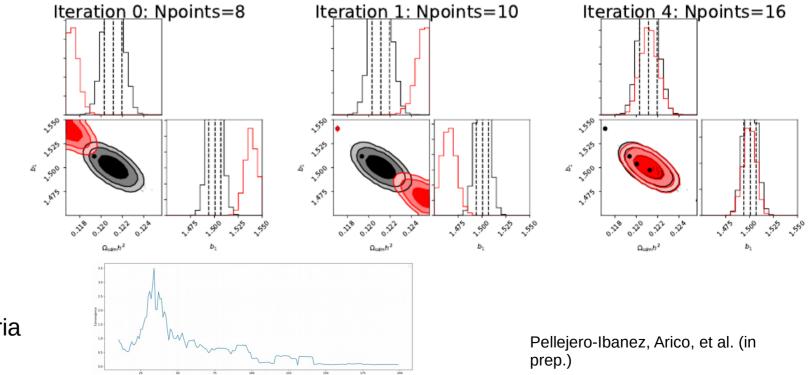
550

Appendix: MCMC performance

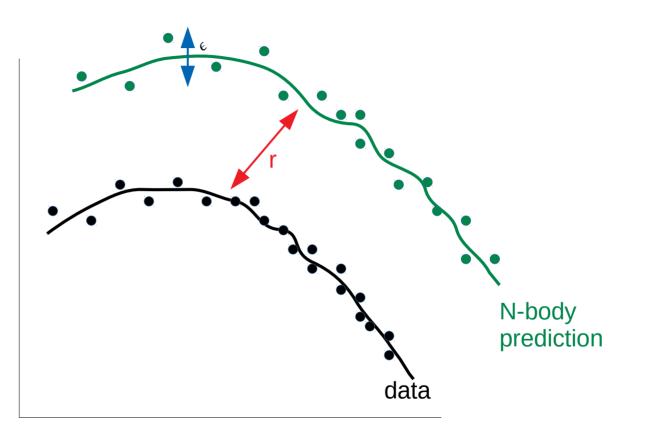




Convergence criteria



Appendix



$$\epsilon = \epsilon(\theta) !!$$

Translates into noise in the likelihood.

