

CRESST

Cryogenic Rare Event Search with Superconducting Thermometers



UNIVERSITY OF OXFORD

Detectors, Shielding and (α, n) Background in the CRESST Dark Matter Search

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for the CRESST Collaboration

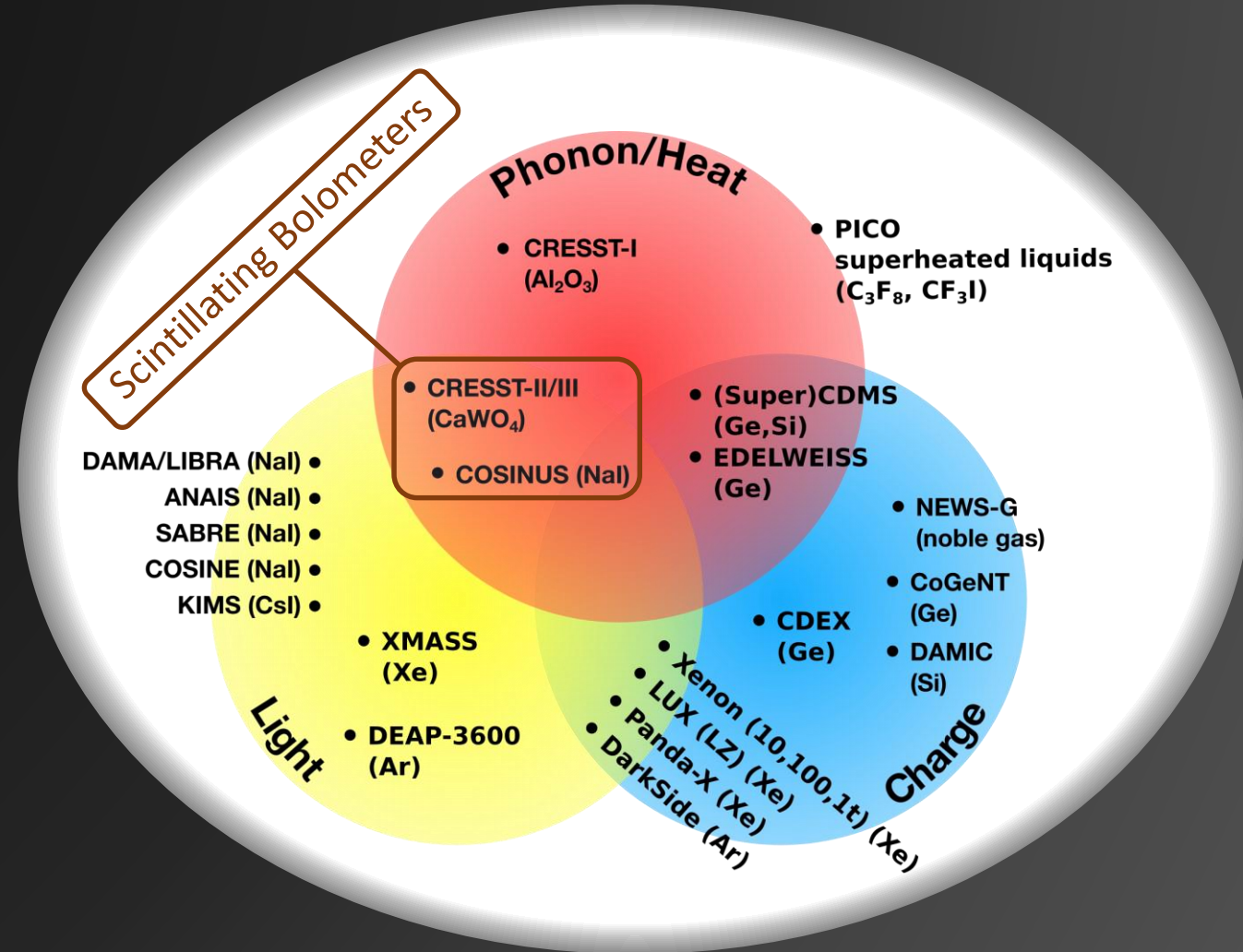
(α, n) yield in low background experiments
21-22 November 2019, CIEMAT, Madrid

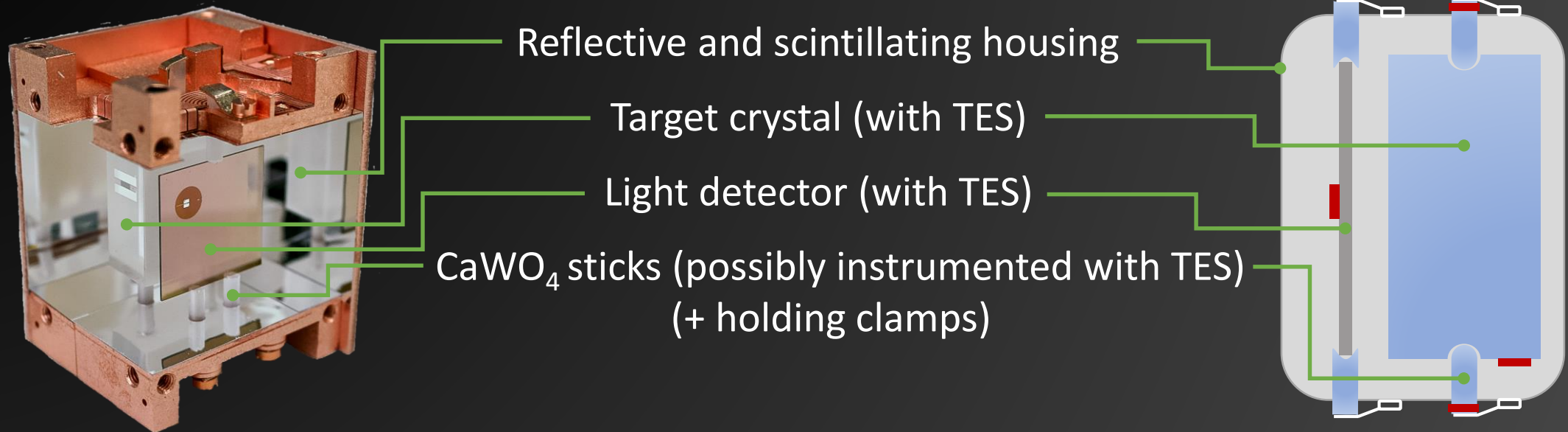
- CRESST Detectors
- Shielding used in CRESST
- Background Simulation with focus on possible (α, n) contribution

The CRESST Collaboration



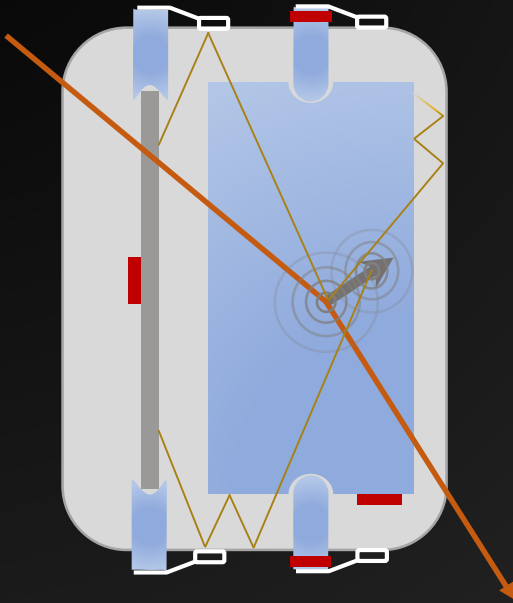
Searching for Dark Matter (DM)
via direct detection



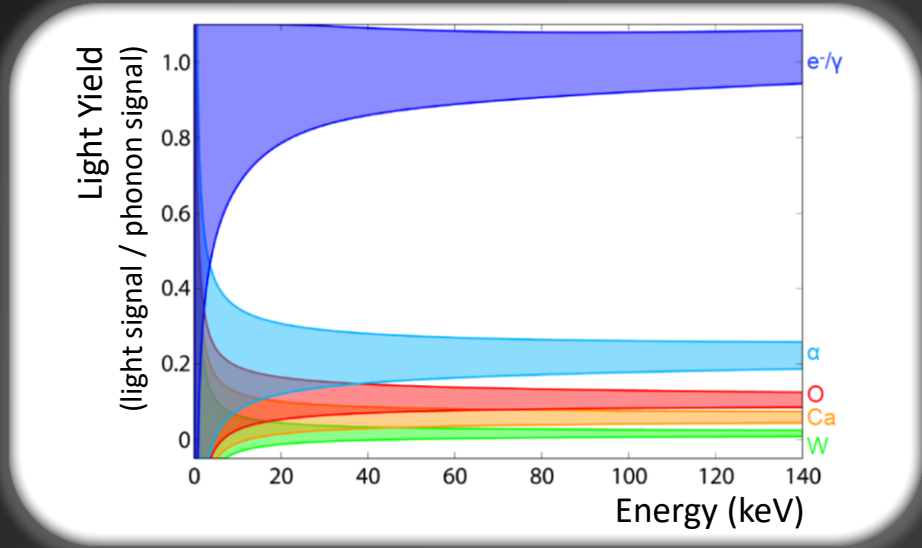
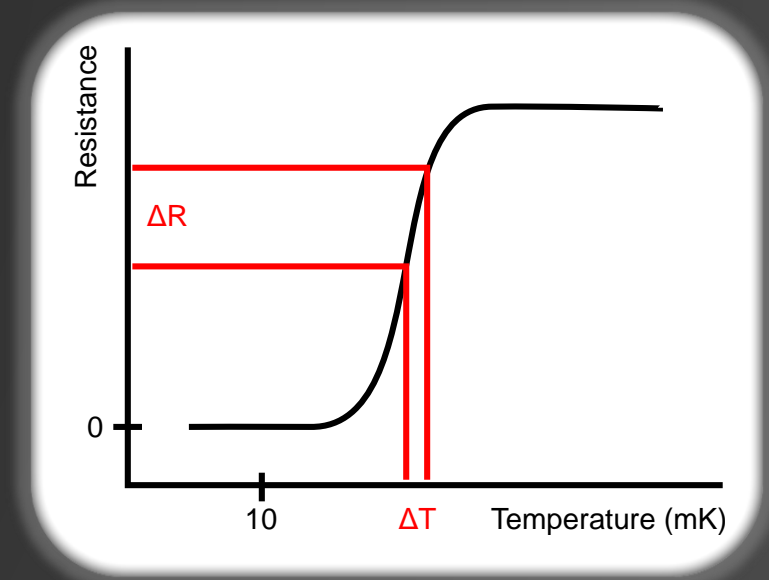


Detector design optimized for **low-mass dark matter search**:

- **Small cuboid crystals** (20x20x10) mm³ → mass: ~24 g
- Nuclear recoil energy **threshold < 100 eV**
- **Low background rate in region of interest** (ROI: threshold to ~16 keV) → ~4 – 6 dru
- **Veto against surface related background** (scintillating housing + instrumented sticks)



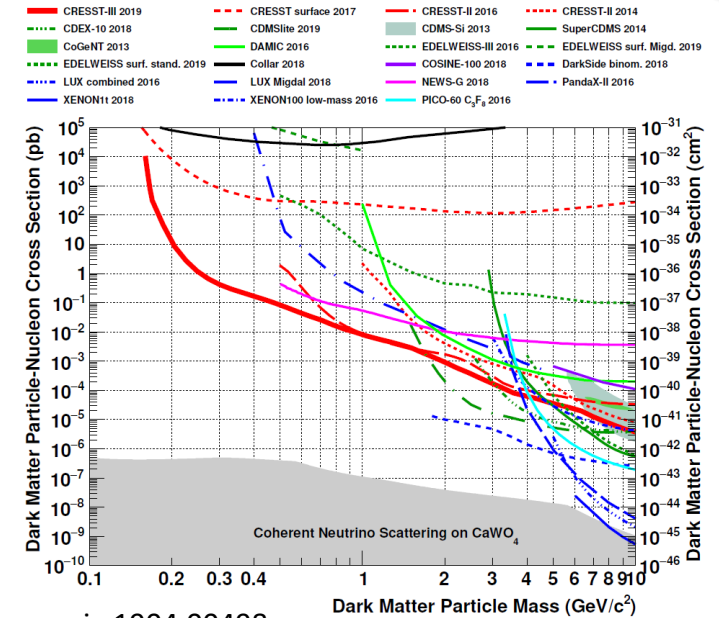
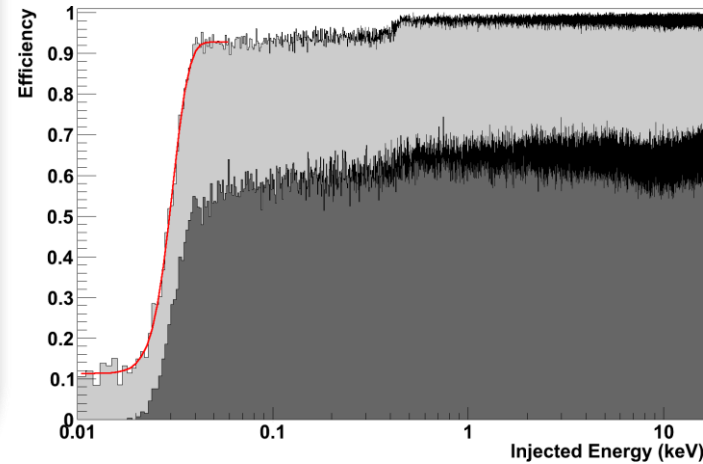
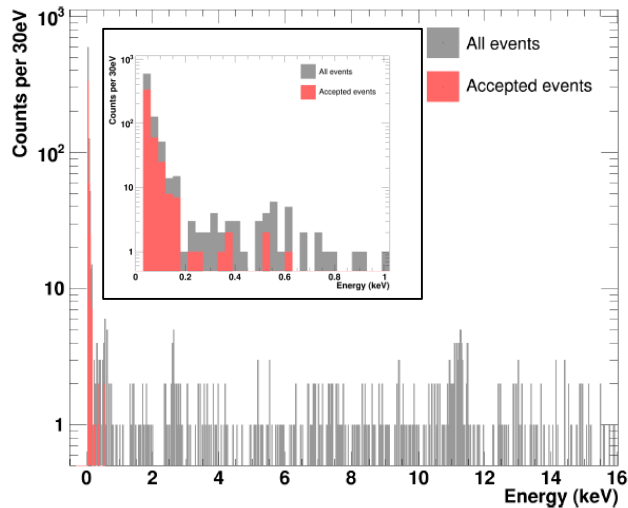
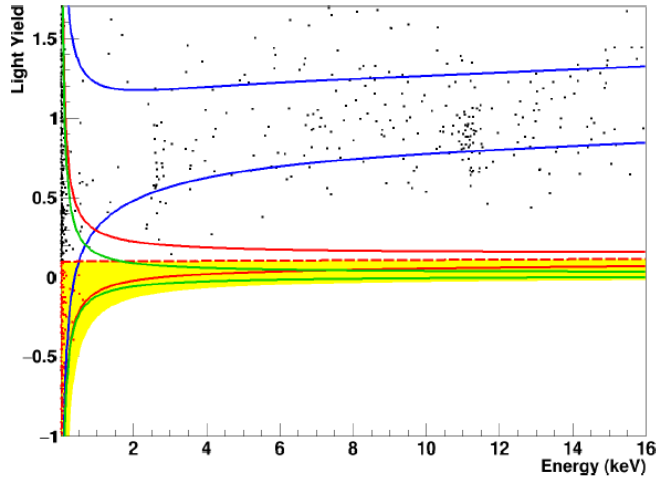
- Cryogenic operation at ~ 15 mK
- **Phonon + scintillation light signal** simultaneously read out with TESs
- **Particle discrimination** via ratio between light and phonon signal



First CRESST-III Results

Detector A (data taking: 10/2016 – 01/2018):

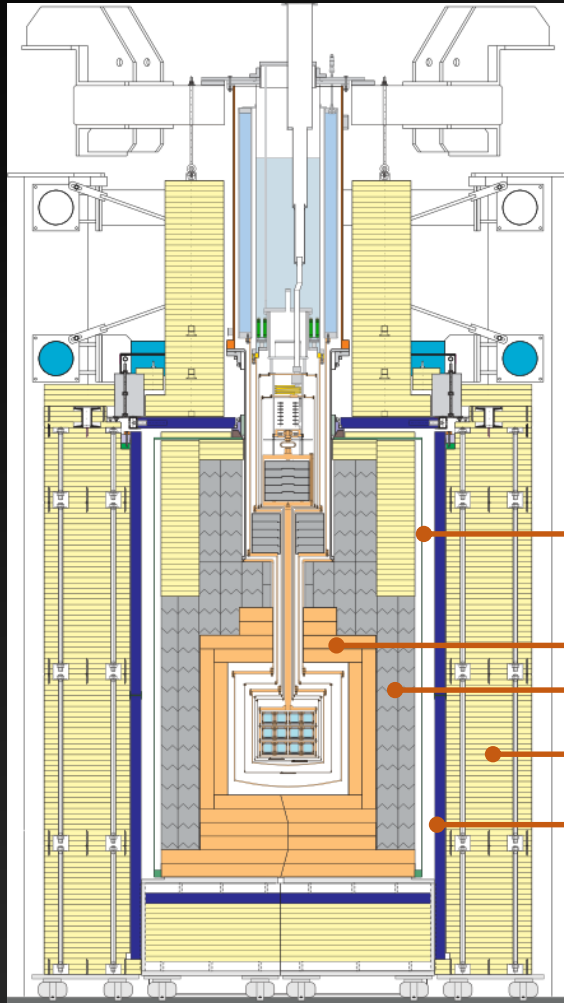
- Target crystal mass: 23.6 g
- Nuclear recoil threshold: 30.1 eV
- Gross exposure: 5.689 kg days



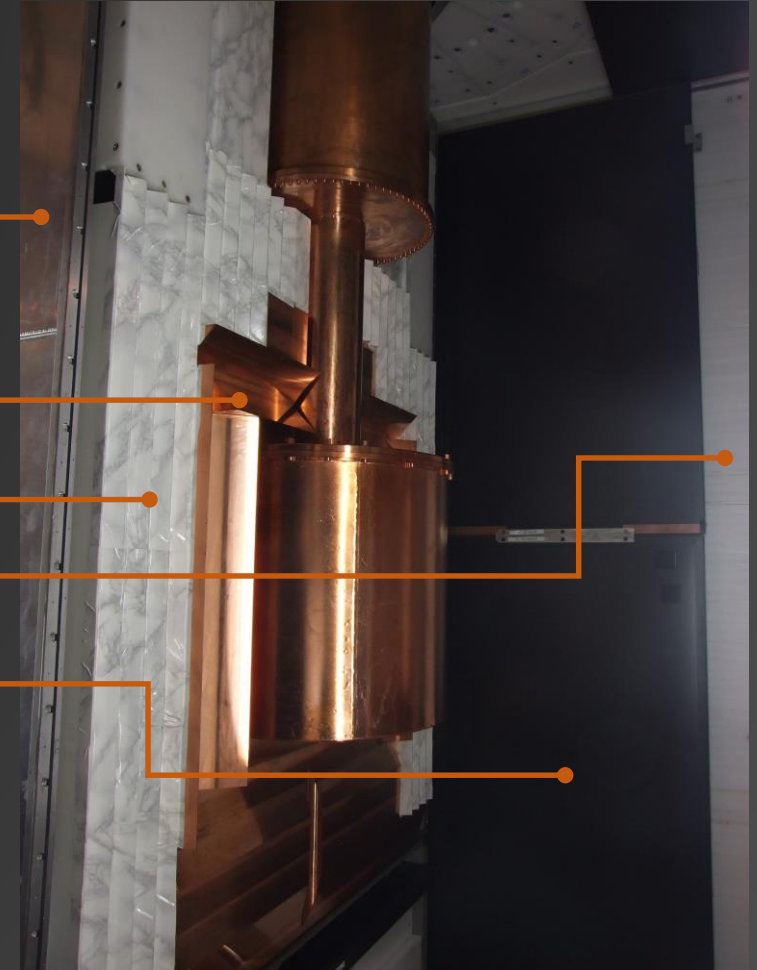
arxiv:1904.00498

CRESST Shielding Structure

Located at the LNGS Underground Laboratory
~3600 m.w.e. rock overburden



- Radon Box
- Min. 14 cm Cu
- Min. 20 cm Pb
- Min. 45 cm polyethylene (PE)
- Active muon veto



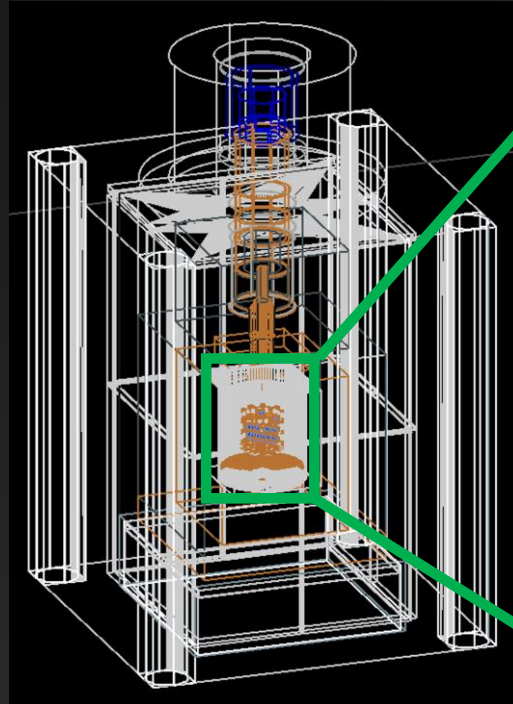
Additional inner PE shield directly outside and inside the OVC of the cryostat since 2012



- A dedicated simulation code based on Geant4 has been developed
- The shielding geometry has been implemented (up to necessary levels of detail)

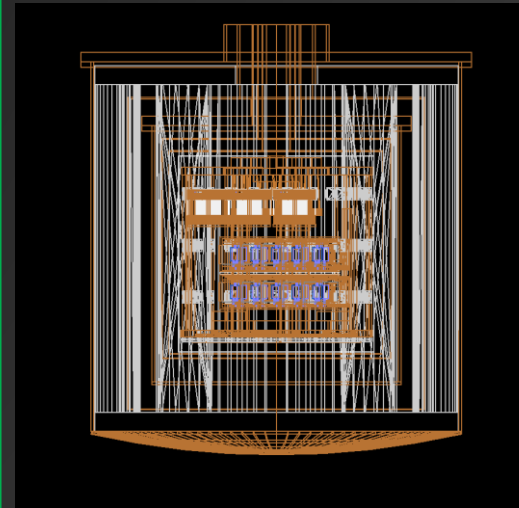
Simulated Geometry

Whole Setup

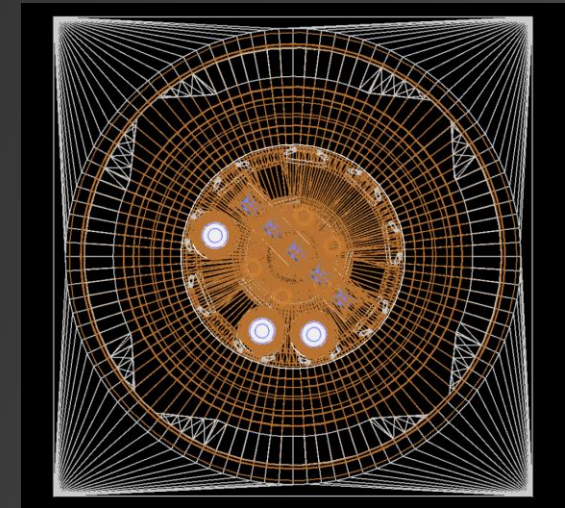


Cryostat Cavity (Cryoshields + Carousel)

Front View

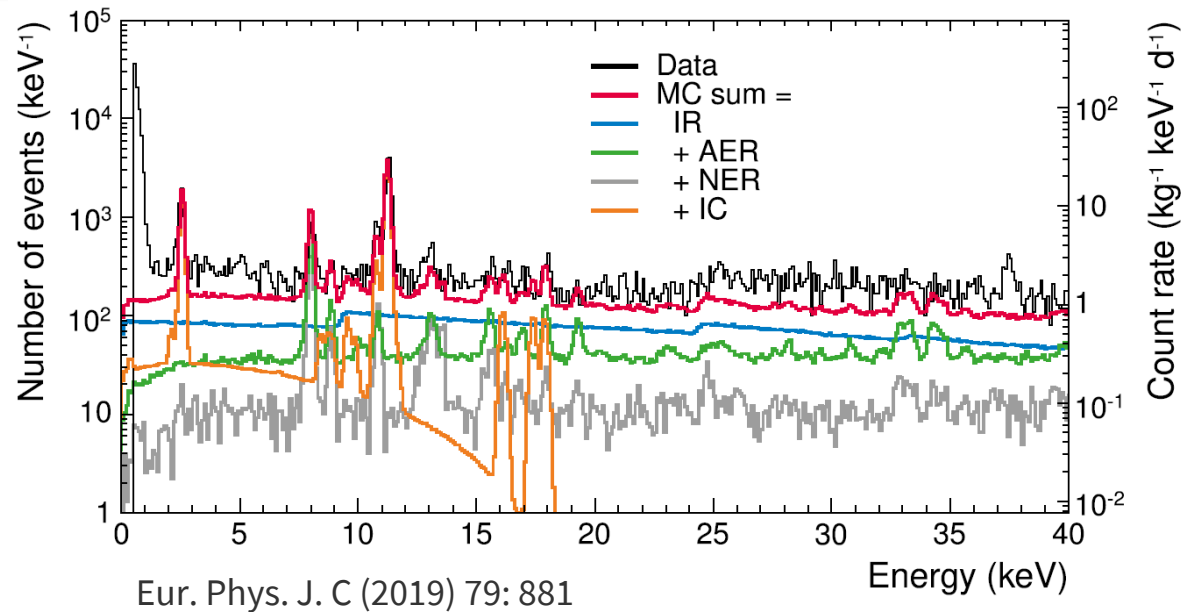
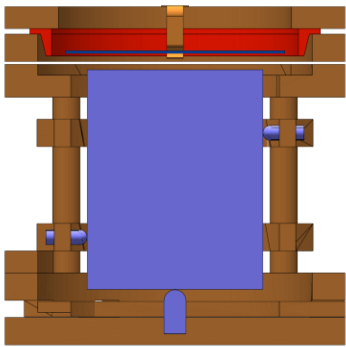


Top View



- An electromagnetic background model has been developed for the lowest-background module (TUM40) of CRESST-II

TUM40
schematic design

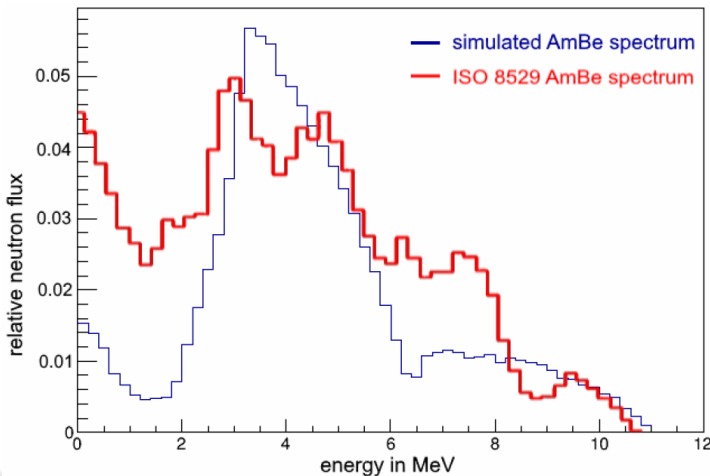


$$(\text{MC} / \text{Data})_{1-40 \text{ keV}} \leq (68 \pm 16) \%$$

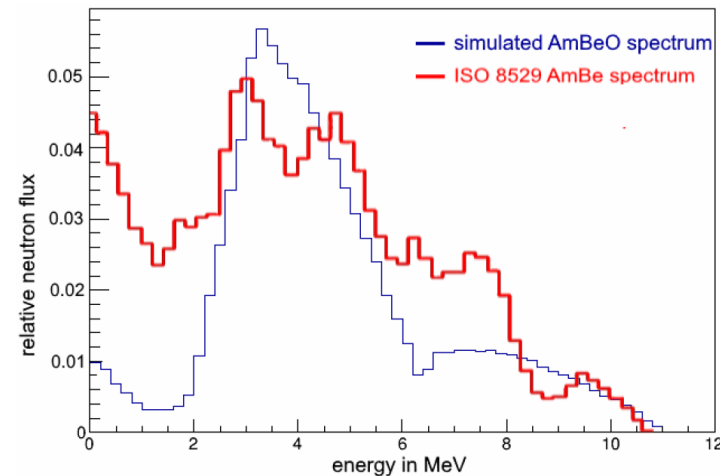
- Neutron background studies were done for past shielding configurations
- Neutron studies for current shielding and detector design are ongoing
- All sources are taken into account (ambient, radiogenic, μ -induced)
- For radiogenic neutrons, we use the SOURCES4C code
 - Radiogenic neutrons include those produced in (α,n) reactions as well as those produced through spontaneous fission reactions

- In 2016, we did a comparison of a real and a simulated (**Geant4.10.2.1**) AmBe neutron spectrum
- The **simulated spectra did not match** the measured spectrum for any simulated configuration:

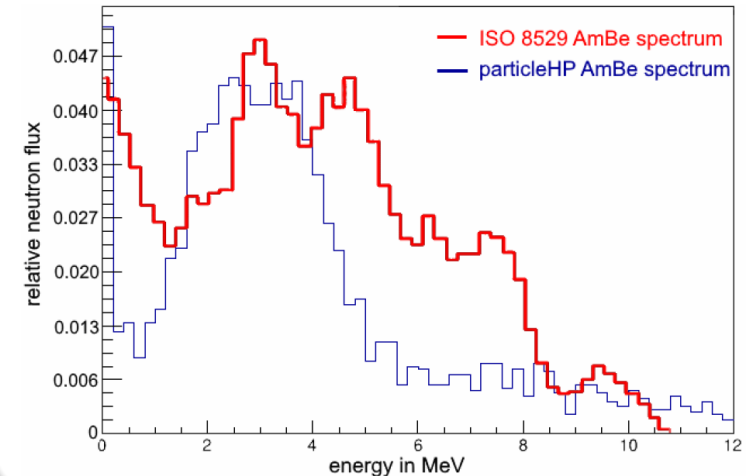
Using pure ^9Be base material



Using BeO base material

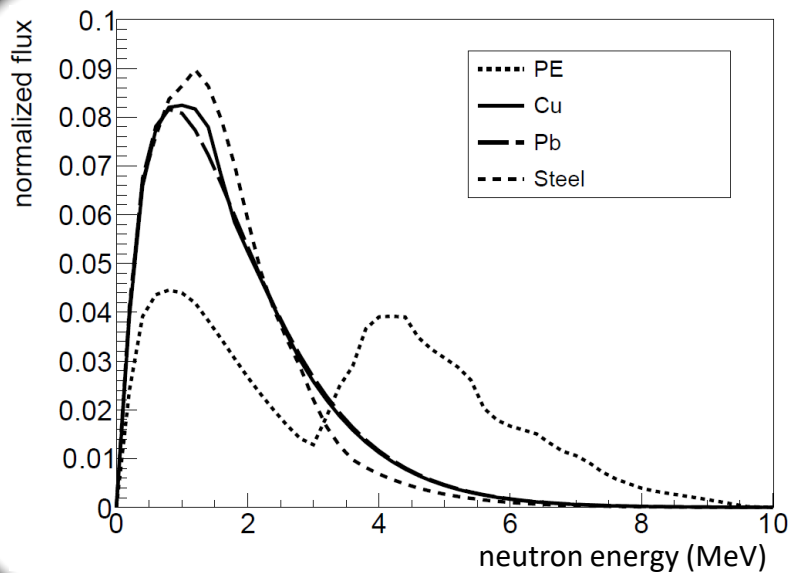


Using pure ^9Be base material
+ ParticleHP model



- Conclusion: **Geant4.10.2.1 does not simulate (α ,n) reactions reliably**
- We hence use **SOURCES4C** to attain (α ,n) neutrons (as well as s.f. neutrons)

Radiogenic neutron spectrum due to SOURCES4C (spontaneous fission + (α, n))



- PE: $9.368 \cdot 10^{-12} \text{ n / (cm}^3 \text{ s)}$
- Cu: $6.607 \cdot 10^{-13} \text{ n / (cm}^3 \text{ s)}$
- Pb: $1.249 \cdot 10^{-13} \text{ n / (cm}^3 \text{ s)}$
- Steel: $2.995 \cdot 10^{-12} \text{ n / (cm}^3 \text{ s)}$

- As a first approach, **contamination levels measured by other rare event search experiments** were considered

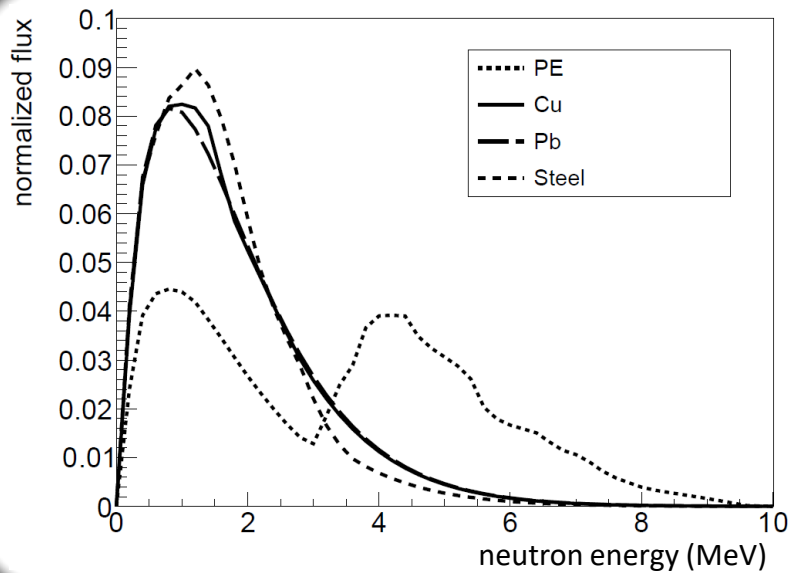
Reference values from CUORE and XENON:

C. Alduino et al., JINST 11 07 (2016), P07009
E. Aprile et al., Astroparticle Physics 35 2 (2011), p. 43
D. R. Artusa et al., Eur. Phys. J. C74 (2014), p. 3096

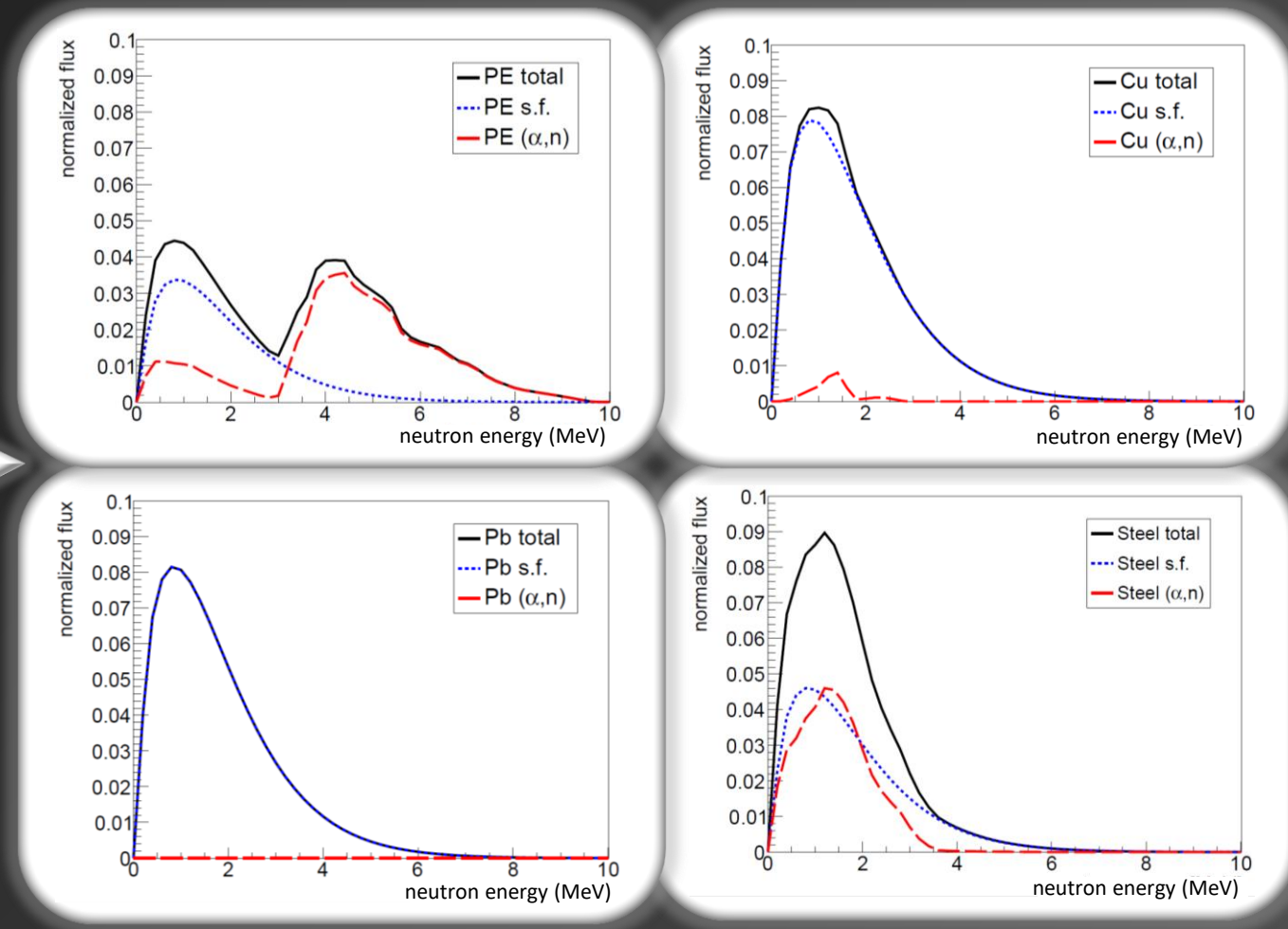
Material	^{238}U [mBq/kg]	^{235}U [mBq/kg]	^{232}Th [mBq/kg]
Cu	< 0.065	–	< 0.002
PE	< 3.8	< 0.37	< 0.14
Pb	< 0.01	–	< 0.07
Steel	< 0.2	–	< 0.1

- Currently, **screening** of our materials is **ongoing** to **characterize new batches** and **improve sensitivity** of old measurements
- Next step: simulations using our screening results

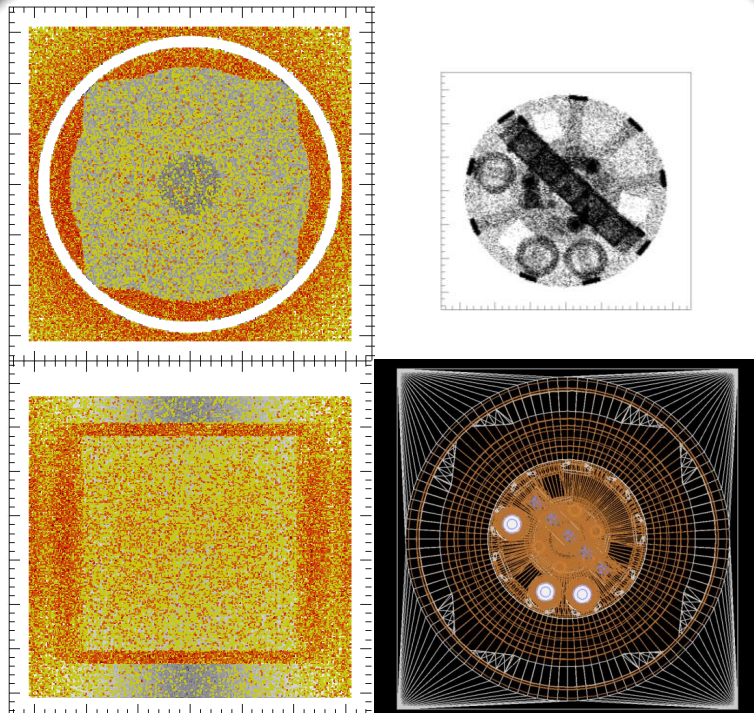
Radiogenic neutron spectrum due to SOURCES4C (spontaneous fission + (α,n))



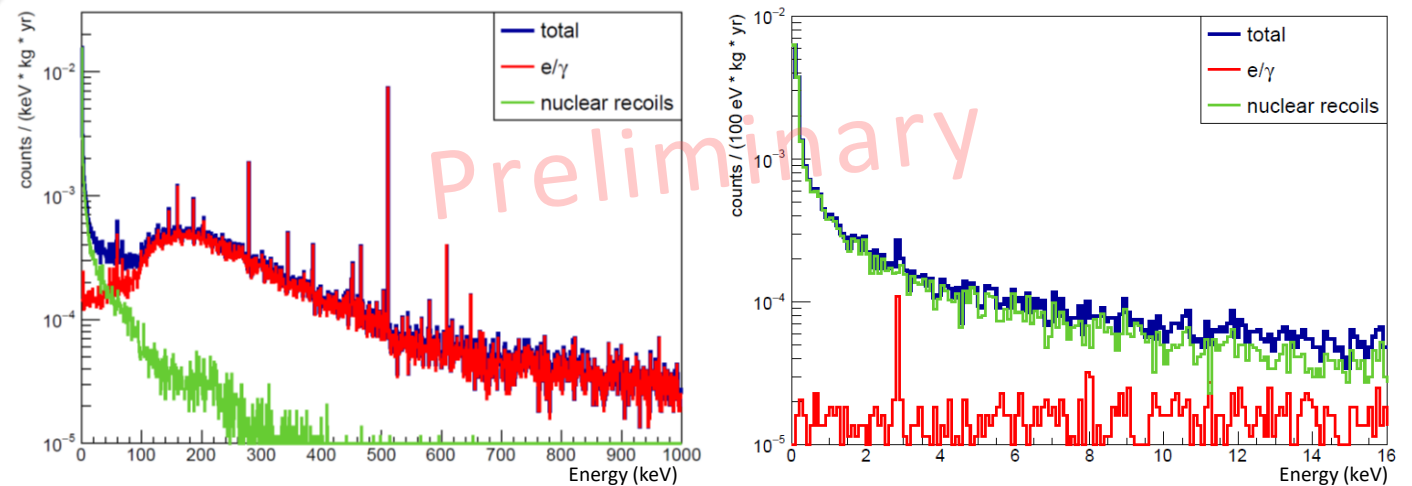
- PE: $9.368 \cdot 10^{-12} \text{ n / (cm}^3 \text{ s)}$
- Cu: $6.607 \cdot 10^{-13} \text{ n / (cm}^3 \text{ s)}$
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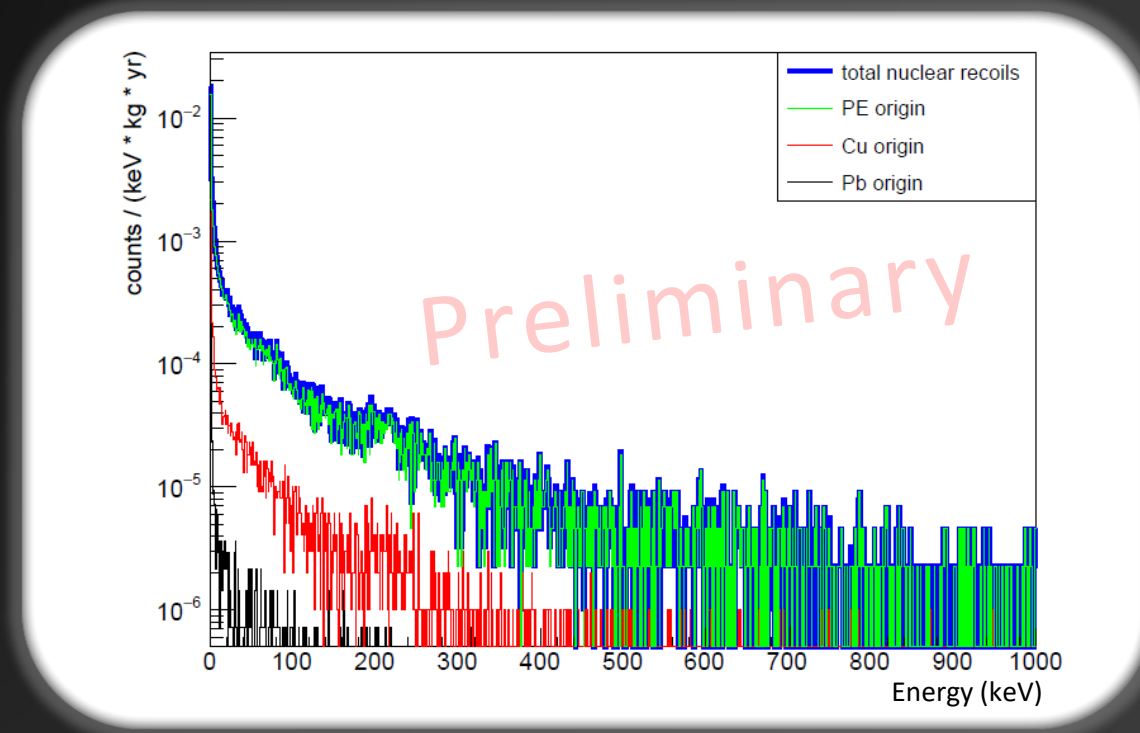
Simulation of homogeneous contamination in inner PE



Detected events originating from radiogenic neutrons produced in the inner PE shields



Total nuclear recoil background due to radiogenic neutrons



- With assumed contamination levels, the inner PE shields actually contribute most to the radiogenic neutron background
- But: expected radiogenic neutron background is very low $\rightarrow O(10^{-2} \text{ kg}^{-1}\text{yr}^{-1})$

- CRESST has a well-shielded setup and highly sensitive detectors for low-mass Dark Matter search
- Screening measurements and simulations for developing and improving background models are ongoing
- (α,n) reactions are not supposed to give a high contribution to our background, but a precise knowledge of the processes is necessary for a precise neutron background model



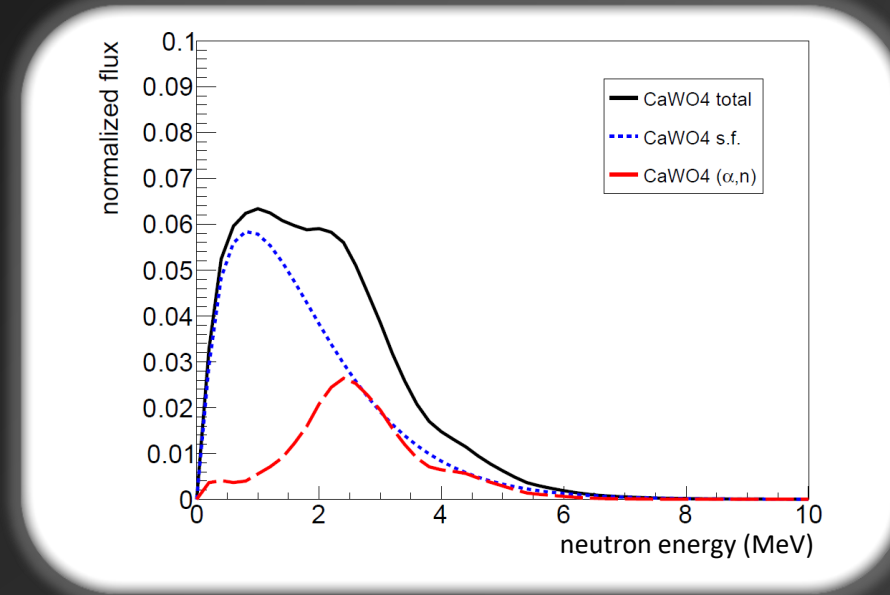
Thank you for your attention!

“You're entitled to say, if you're so smart, why don't you tell me what that dark matter is? And I'll have to confess I don't know” – Jim Peebles

Additional Material

Radiogenic neutrons due to intrinsic contamination in CaWO₄,
taking contamination levels of TUM40 as reference:

Material	²³⁸ U [mBq/kg]	²³⁵ U [mBq/kg]	²³² Th [mBq/kg]
CaWO ₄	1.073	0.045	0.011



$$1.003 \cdot 10^{-11} \text{ n / (cm}^3 \text{ s)}$$

CRESST Shielding Material Contamination Levels

Material	²³⁸ U [mBq/kg]	²³⁵ U [mBq/kg]	²³² Th [mBq/kg]
Cu (Cryostat + Detector Holders)	< 0.02	< 0.05	< 0.021
Outer PE	52.66	–	43.55
Inner PE	1.0 ± 0.1	< 0.28	0.3 ± 0.1
Pb	< 2.85	–	< 0.91

This is not meant to be a complete list of our screening measurements
(e.g. not listed are crystals and scintillating foils)

ICP-MS and NAA of new batch of Cu currently ongoing