

Workshop on  $(\alpha, n)$  yield in low background experiments  
CIEMAT, Madrid 21-22 November, 2019

$(\alpha, n)$  backgrounds in **nEXO** 

*Andrea Pocar*  
*University of Massachusetts, Amherst*



UMass  
Amherst



AMHERST CENTER FOR FUNDAMENTAL INTERACTIONS  
Physics at the interface: Energy, Intensity, and Cosmic frontiers  
University of Massachusetts Amherst

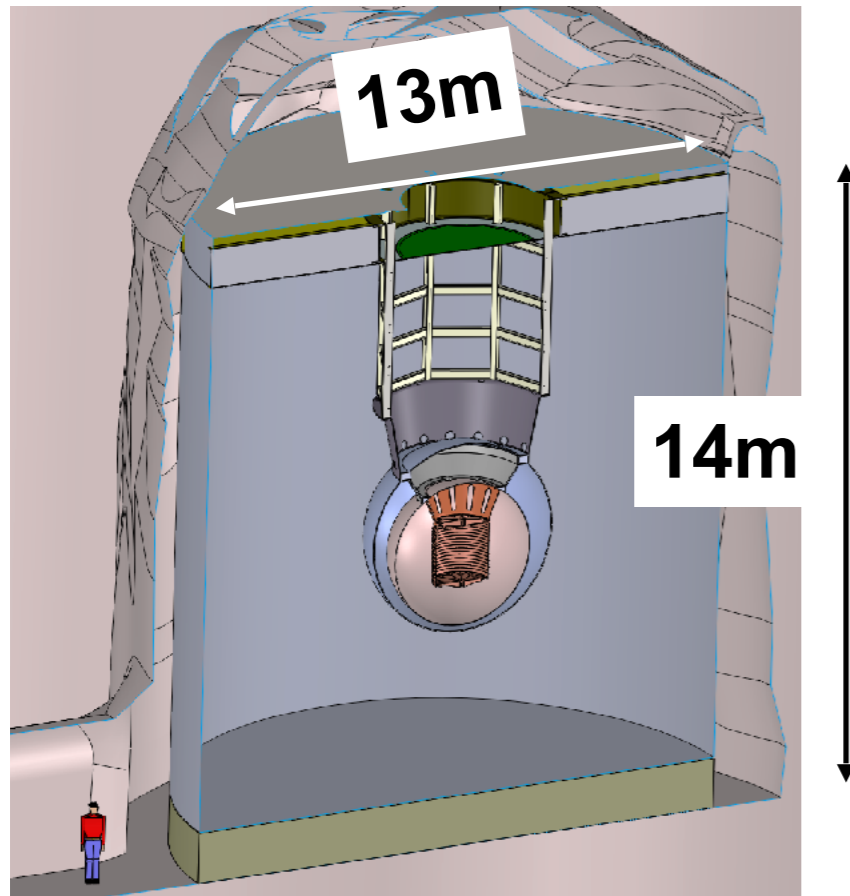
# Outline

---

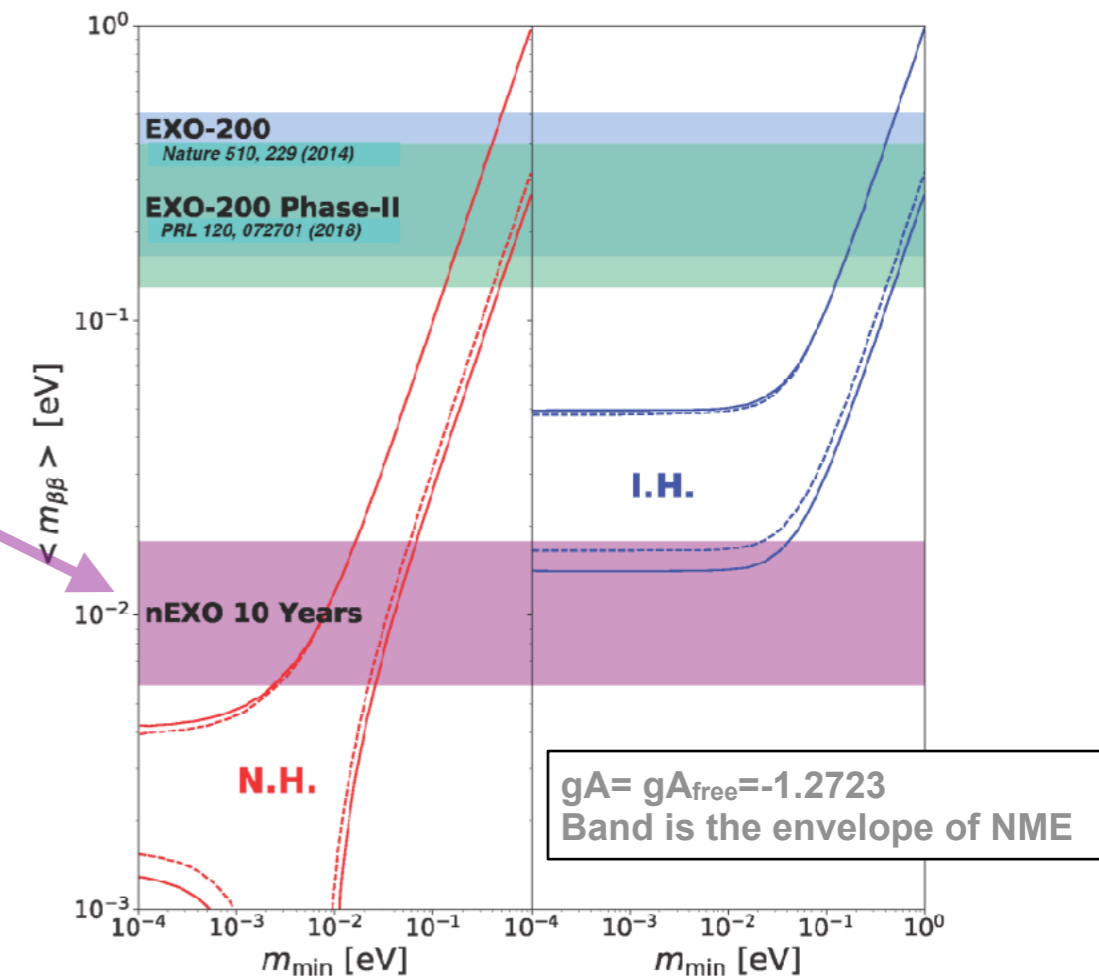
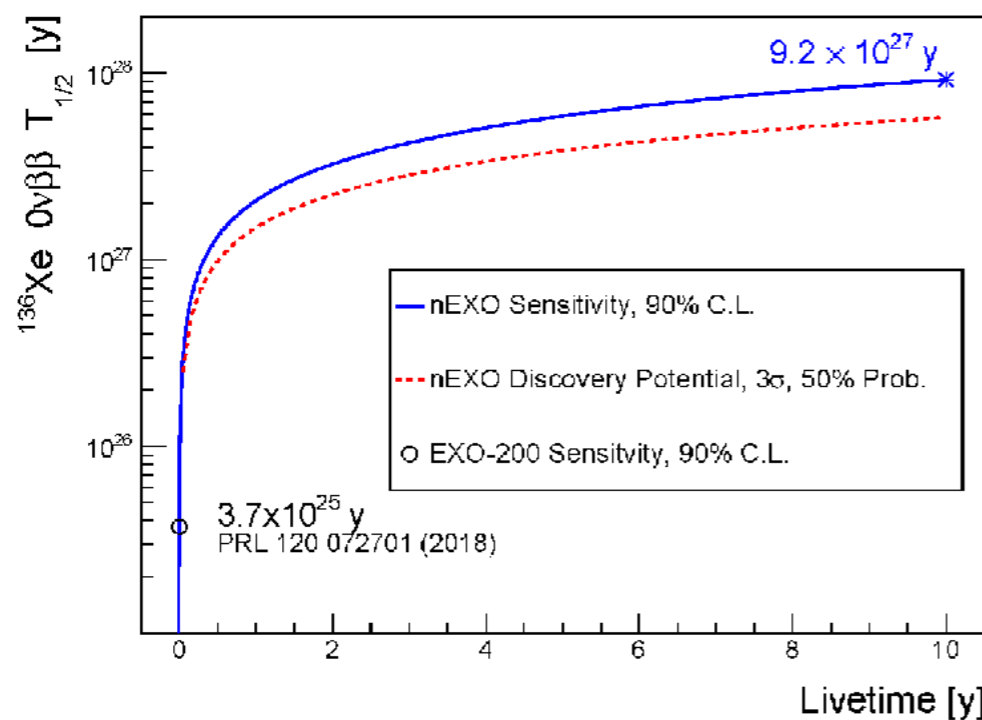
- *the nEXO detector*
  - *physics goals and backgrounds*
- *( $\alpha, n$ ) background estimations*
  - *background channels*
  - *treatment so far*
- *open issues and outlook*

$\beta$

$\beta$

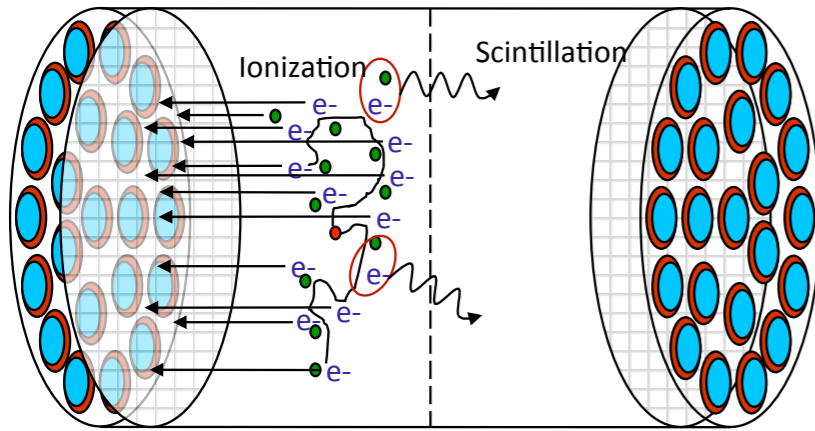
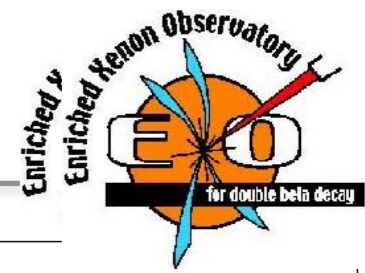


- Ultra-low background ‘core’
- Precisely measure background at the periphery
- Incorporate knowledge of background in sensitivity calculation
- ‘Background index’ is fiducial volume-dependent

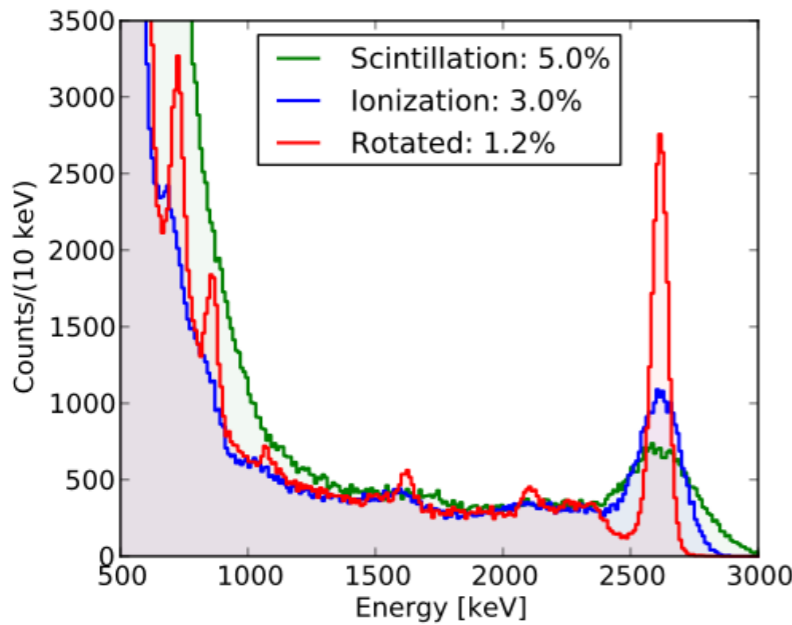
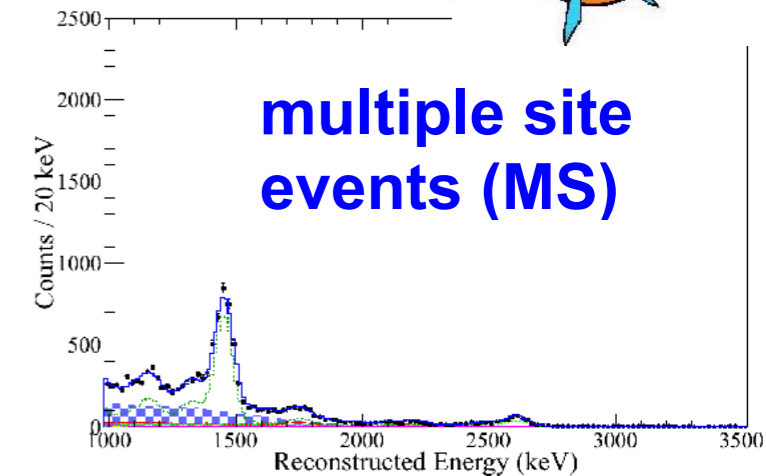
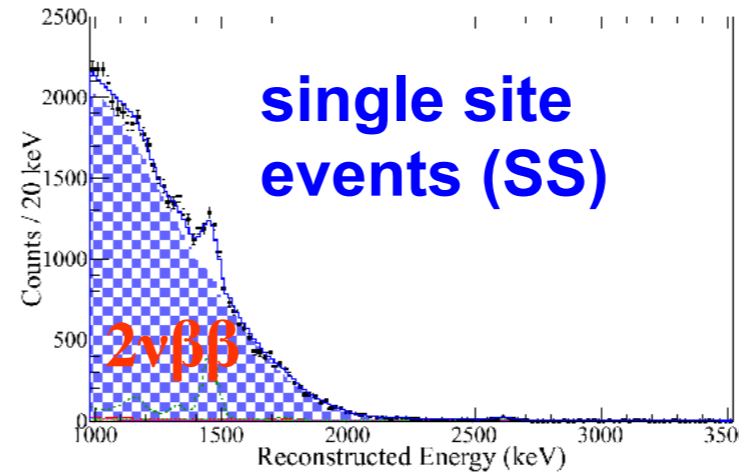


# The EXO-200 precursor

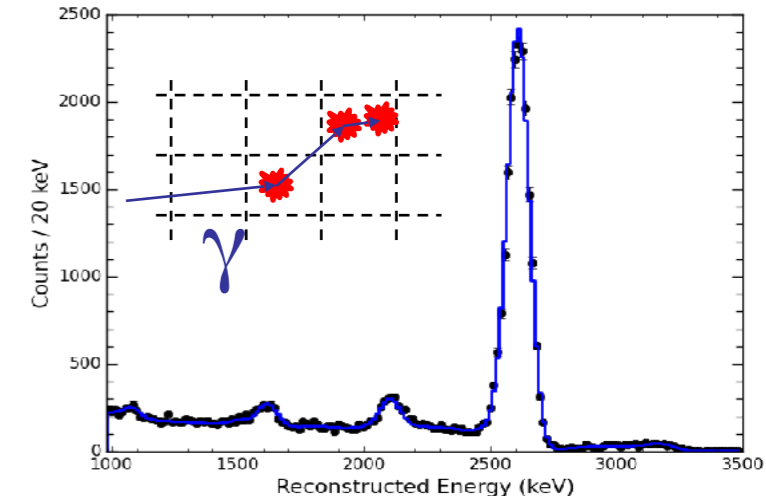
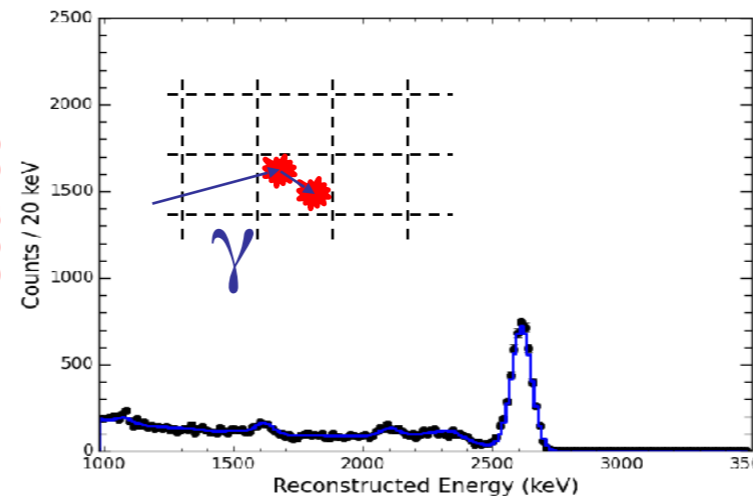
PRL 123(2019)161802



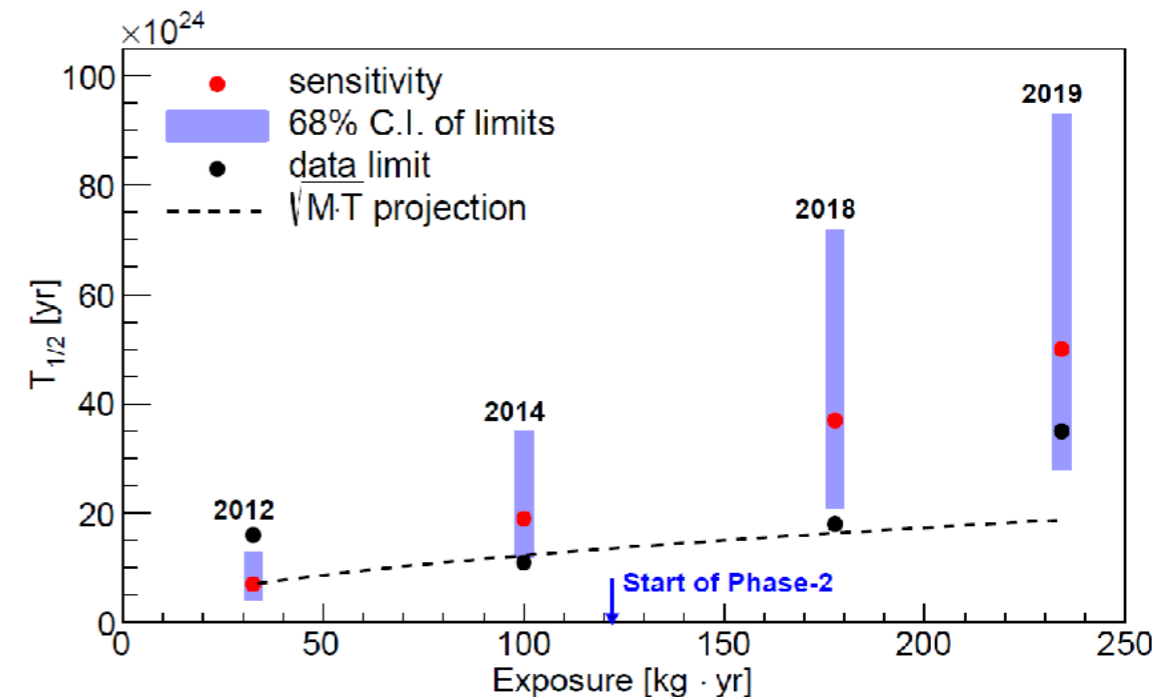
Low background data

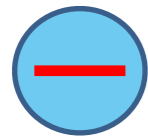


<sup>228</sup>Th calibration source

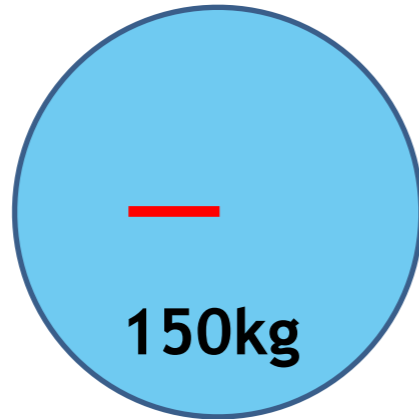


**Phase I+II: 234.1 kg·yr <sup>136</sup>Xe exposure**  
**Limit  $T_{1/2}^{0\nu\beta\beta} > 3.5 \times 10^{25}$  yr (90% C.L.)**  
 $\langle m_{\beta\beta} \rangle < (93 - 286)$  meV  
**Sensitivity  $5.0 \times 10^{25}$  yr**

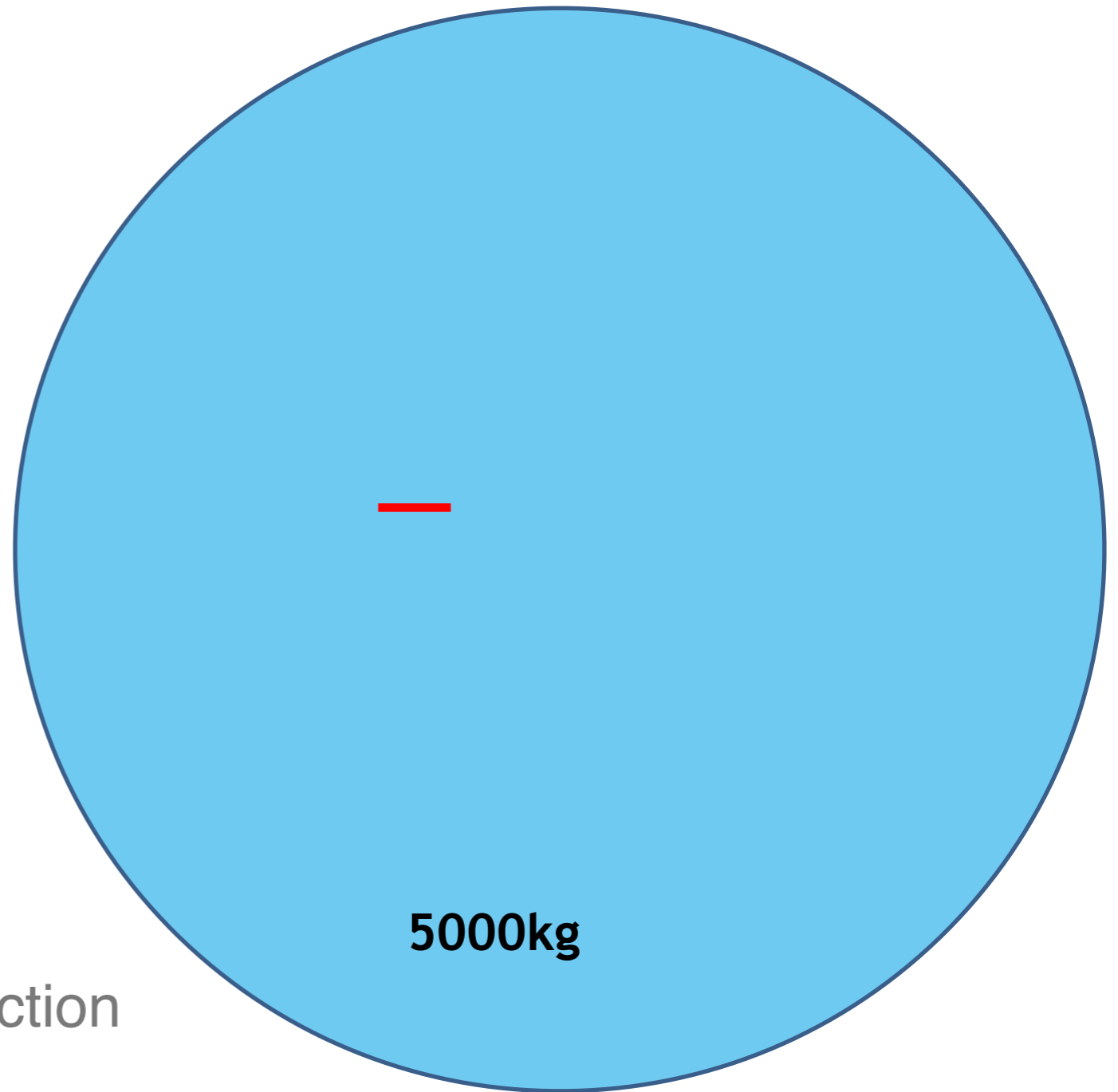





5kg



150kg



5000kg

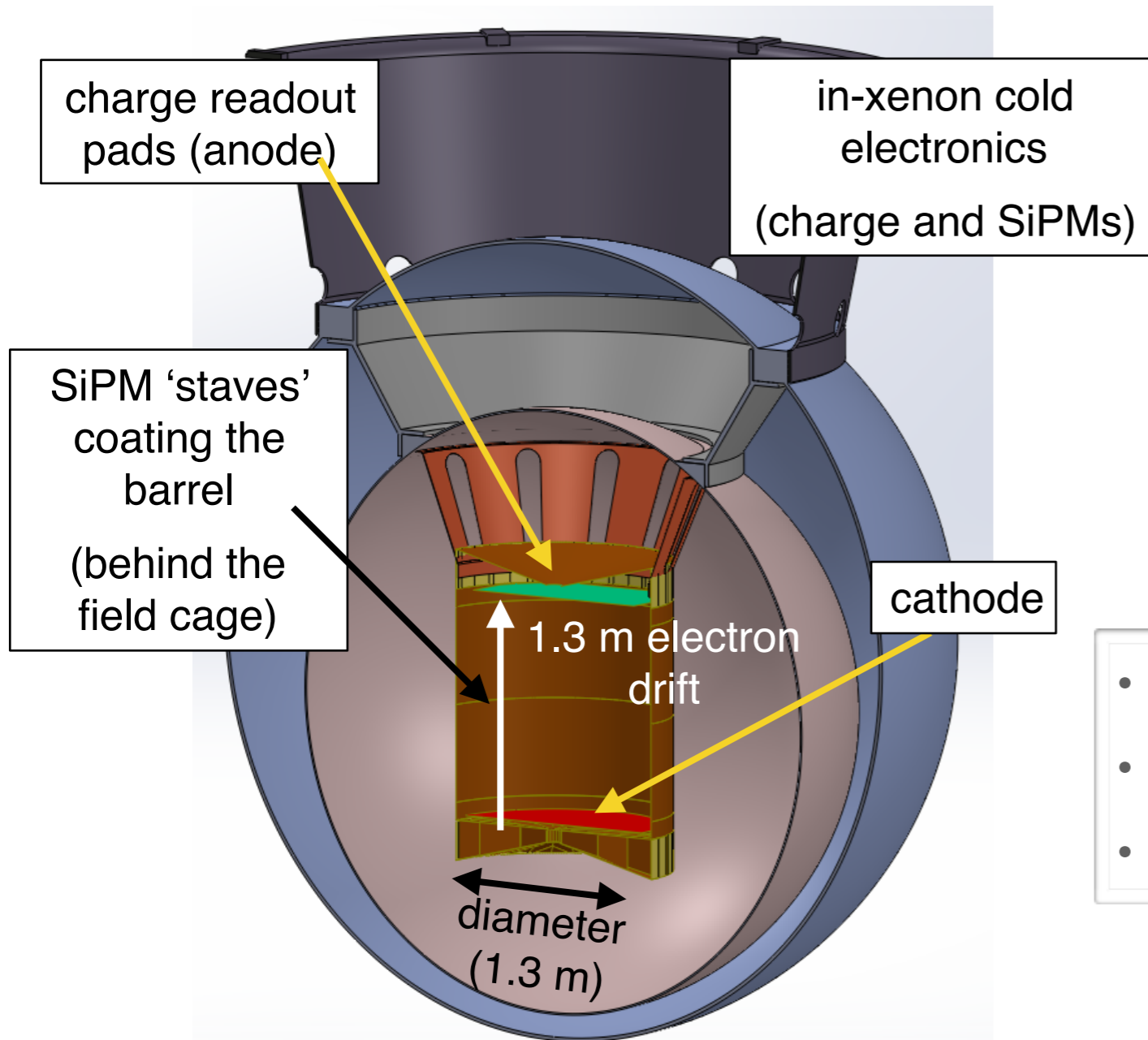
 **Attenuation Length of  
a 2.4 MeV  $\gamma$ -ray in LXe  
(~ 8.5 cm)**

take full advantage of:

- Compton tag and rejection
- External background ID and rejection

**The larger and monolithic the detector, the more useful this is.**

- Ton scale is where these features become dominant
- Neutron attenuation scales differently



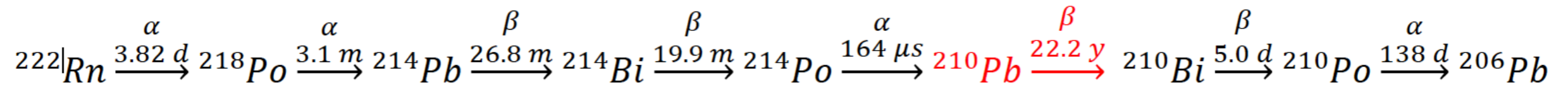
- $< 1\%$  energy resolution
- no central cathode
- $\geq 10$  ms electron lifetime
- $\sim 500$  Rn atoms

- no plastics, in-Xe cold electronics
- VUV-sensitive SiPMs behind field cage
- charge readout strips

- 25x EXO-200
- enhanced self-shielding
- x100 better  $T_{1/2}$  sensitivity

- sensitivity (10 years):  $9 \times 10^{27}$  yr
- energy, topology, standoff & particle ID

# Are (α,n) backgrounds an issue for nEXO?



${}^{210}\text{Pb}$ :  $Q_{\beta} = 63.5\text{ keV}$   
 ${}^{210}\text{Bi}$ :  $Q_{\beta} = 1162\text{ keV}$   
 ${}^{210}\text{Po}$ :  $E_{\alpha} = 5304\text{ keV}$

α's from upper U chain and Th chain also contribute, but we believe they will be subdominant given the Th/U purity we require

- The Q-value of both β decays is well below  $Q_{\beta\beta} = 2458\text{ keV}$  of  ${}^{136}\text{Xe}$
- $E_{\alpha}$  is well above  $Q_{\beta\beta}$
- Degraded α's on surfaces could have tails at  $Q_{\beta\beta}$ , but skin cuts and light-to-charge ratio would reject them with almost perfect efficiency

1) γ cascades following neutron captures (efficiently vetoed)

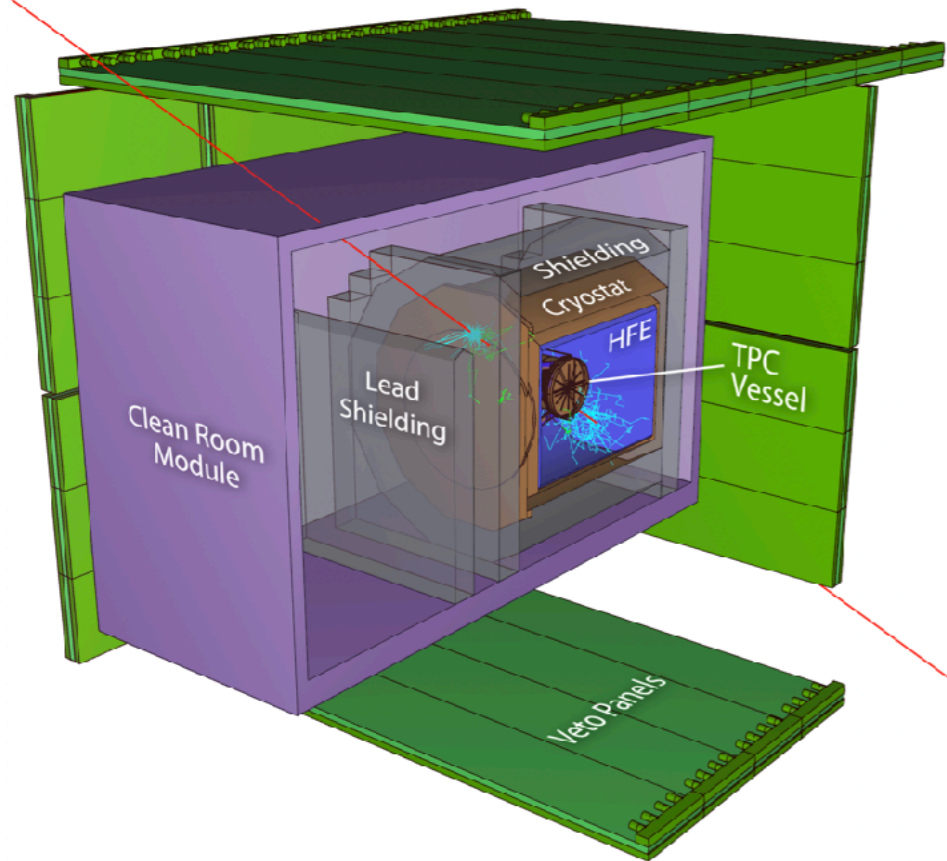
2) long-lived target activation

${}^{137}\text{Xe}$ :  $Q_{\beta} = 4173\text{ keV}$   
 $T_{1/2} = 3.82\text{ min}$

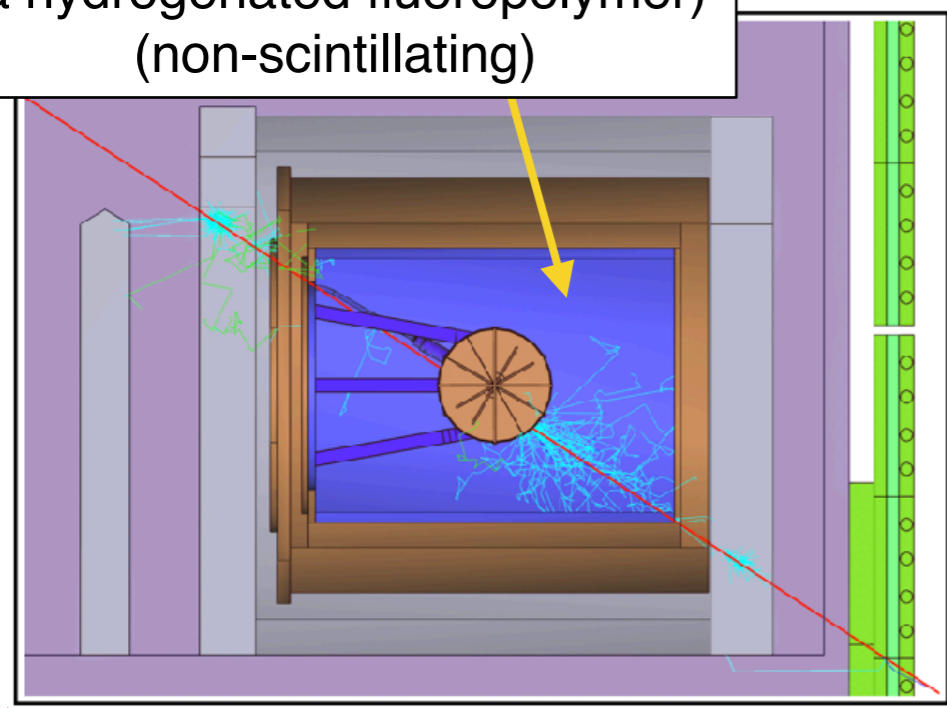
# Neutron capture $\gamma$ cascades

JCAP 04(2016)029

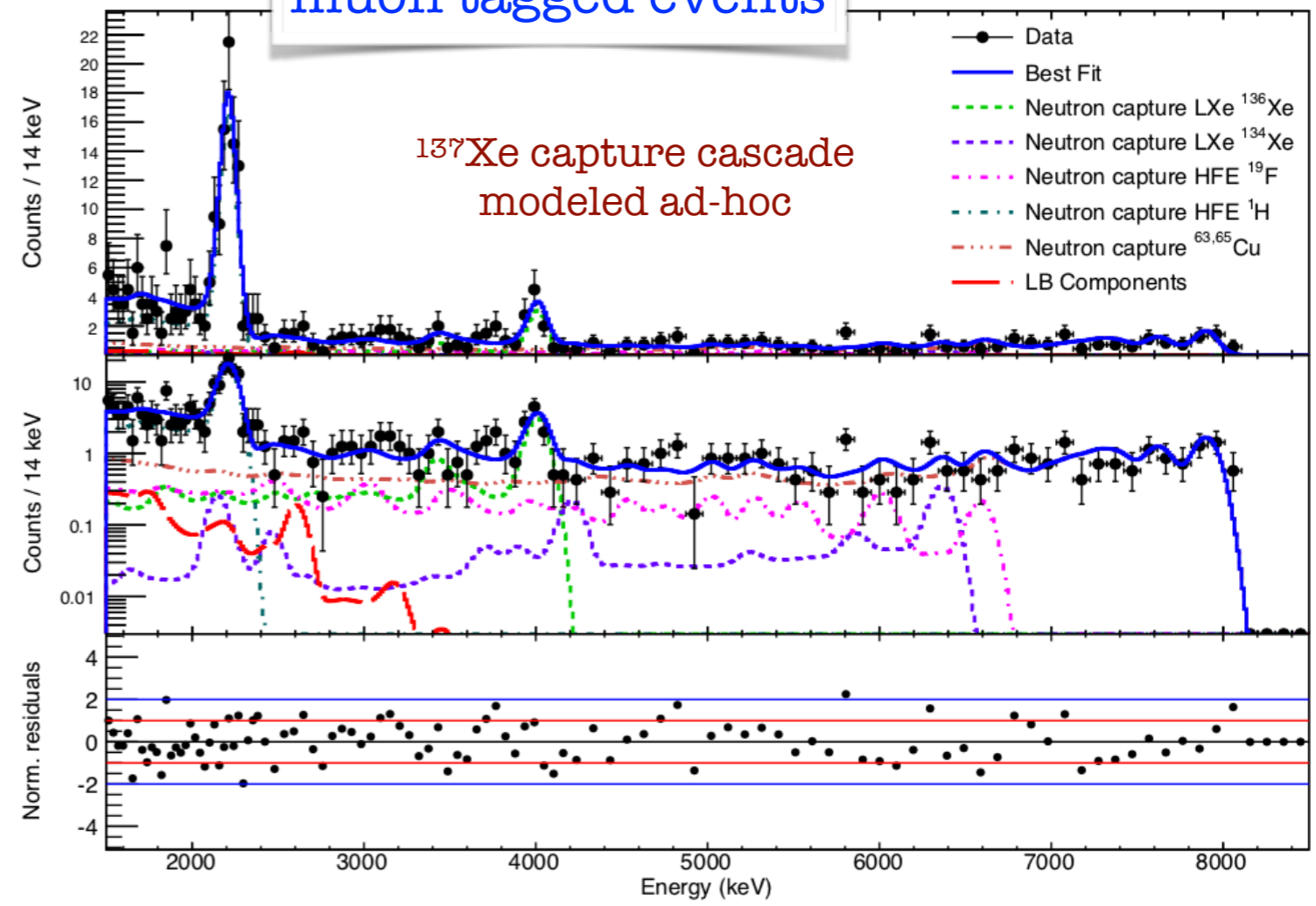
PRC 94(2016)034617



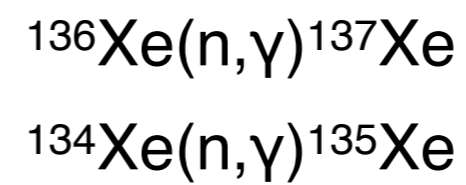
HFE-7000  
(a hydrogenated fluoropolymer)  
(non-scintillating)



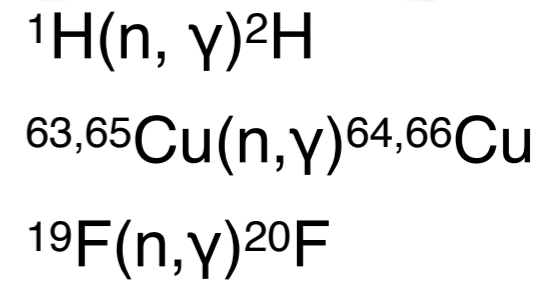
muon tagged events



xenon



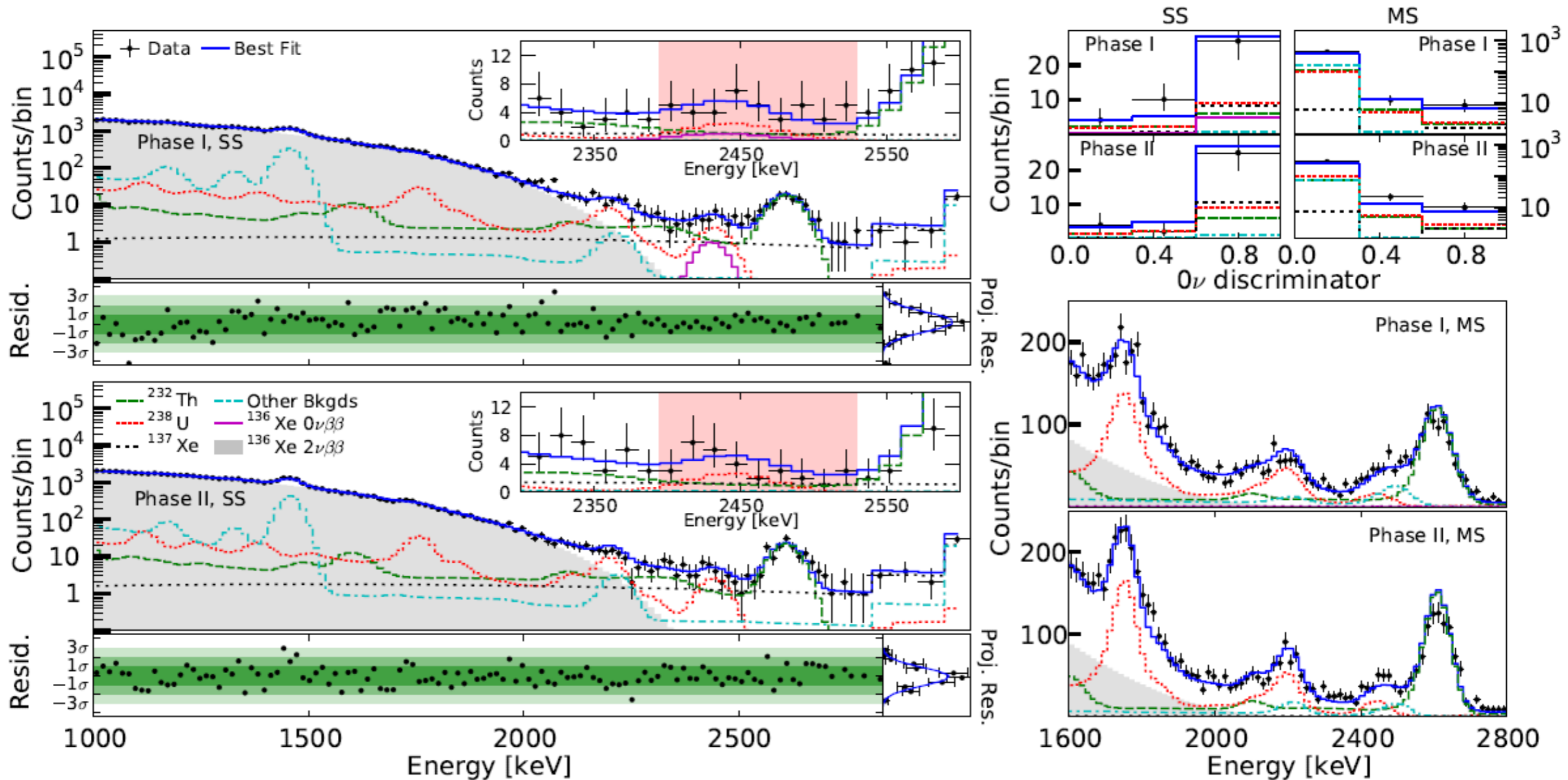
materials





# the EXO-200 final $0\nu\beta\beta$ results

PRL 123(2019)161802



**2019 release uses machine learning (DNN) for improved signal-to-background discrimination**

## 1. What is the $^{210}\text{Pb}/^{210}\text{Po}$ surface decay rate?

- deposition model from Guiseppe et al, arXiv:1101.0126
- all  $^{214}\text{Bi}$  yield  $^{210}\text{Pb}$
- measurements on acrylic, 25 Bq/m<sup>3</sup> of radon

## 2. How many neutrons are emitted per $\alpha$ , which materials matter

- JANIS 4.0 for cross sections (maintained by NEA;  $\alpha$ -energy dependent reaction cross sections  $\sigma(E)$  are taken from the JENDL, TENDL, ENDF or other data bases)
- alpha stopping power from ASTAR/NIST in nEXO detector materials
- $n/\alpha$ :  $3 \times 10^{-8} \div 6 \times 10^{-6}$  for  $^{210}\text{Po}$  (higher for  $^{232}\text{Th}$  and  $^{238}\text{U}$ )

## 3. How many ROI events will we have per surface neutron?

- no neutron energy distribution, used EXO-200 cosmogenic number
- $\epsilon_{hit} = 0.03$  (hit efficiency in ROI)
- in nEXO, lower  $\epsilon_{hit}$ : better resolution, shielding, lower neutron energy

$$R_b = \epsilon_{hit} \cdot Y_n \cdot \frac{1}{4} \cdot A_{210Po}$$

$$R_b = \epsilon_{hit} \cdot Y_n \cdot \frac{1}{4} \cdot R_d \cdot S \cdot t$$

Material	Composition	$^{210}\text{Po}$ [n/ $\alpha$ ]	Maximal time [y/m <sup>2</sup> ]
Teflon	CF <sub>2</sub>	6.0 · 10 <sup>-6</sup>	2.3
Sapphire	Al <sub>2</sub> O <sub>3</sub>	3.1 · 10 <sup>-7</sup>	45
Acrylic	C <sub>5</sub> O <sub>2</sub> H <sub>8</sub>	5.0 · 10 <sup>-8</sup>	280
Quartz	SiO <sub>2</sub>	3.2 · 10 <sup>-8</sup>	440

assumes 25 Bq/m<sup>3</sup> radon

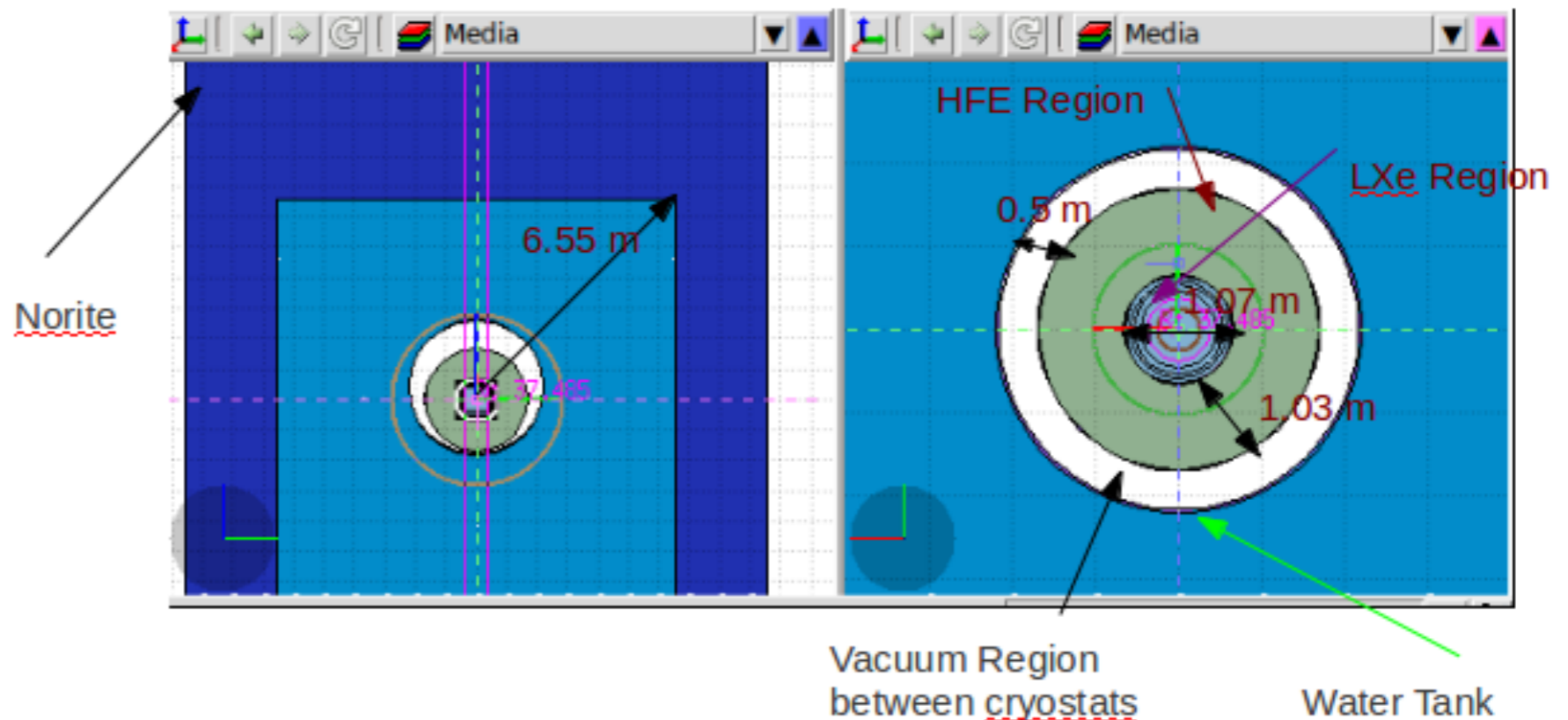
long allowed exposure, does not seem to be a problem

this is not based on proper MC-supported model

Neutron prob. from SF are approx: 10<sup>-7</sup> (<sup>238</sup>U), 10<sup>-11</sup> (<sup>232</sup>Th)

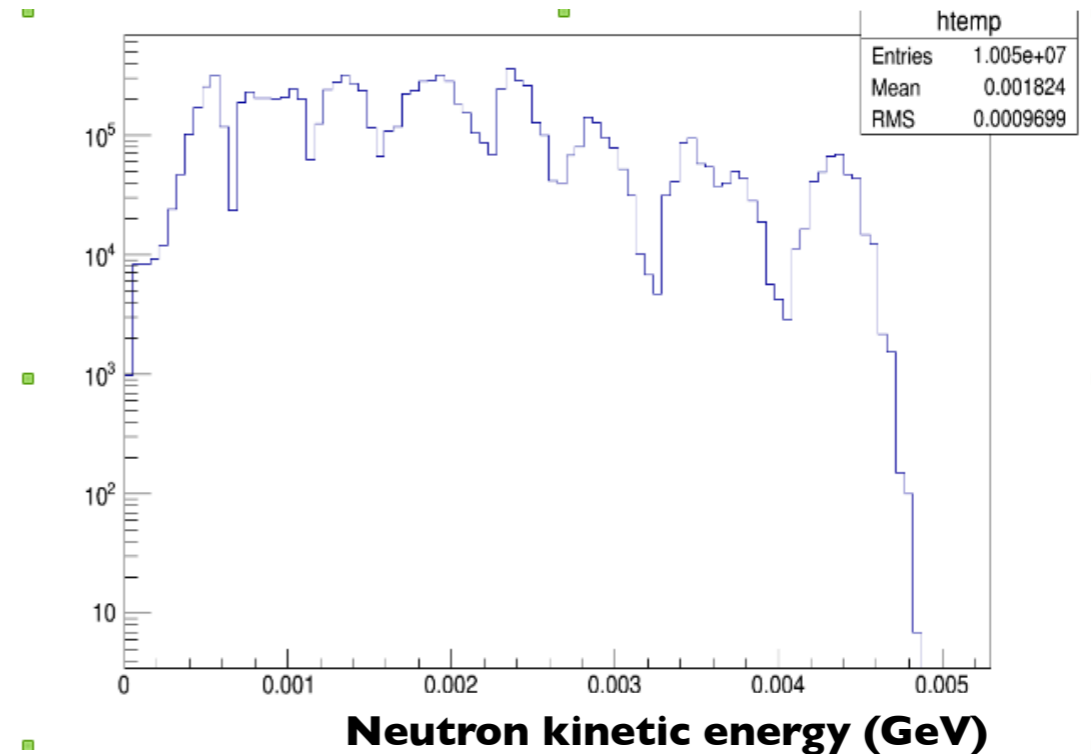
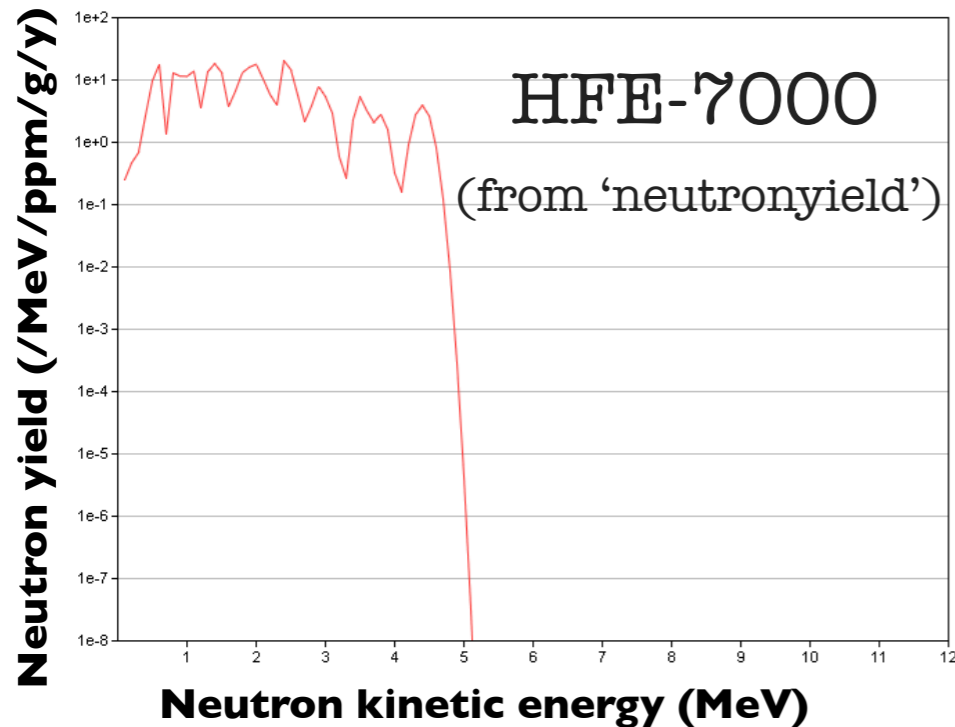
1. FLUKA simulation of neutrons

2. realistic masses of nEXO components



3. compared with D-M Mei – NIMA 606(2009)651 – [neutronyield.usd.edu](http://neutronyield.usd.edu) (entire U chain, not just  $^{210}\text{Po}$ )

3. compared with D-M Mei – NIMA 606(2009)651 – [neutronyield.usd.edu](http://neutronyield.usd.edu) (entire U chain, not just  $^{210}\text{Po}$ )



Compound	Po Neutron Yield (n/alpha)	Xe capture fraction	Maximal Loading Time (y/m <sup>2</sup> )
Sapphire	$3.1 \times 10^{-7}$	0.071	85.84
Quartz	$3.2 \times 10^{-8}$	0.066	894.71
HFE	$4.3 \times 10^{-6}$	0.059	7.47

The HFE will see approx. 8 m<sup>2</sup> of TPC vessel surface.

One year of exposure results in 0.01/y ROI events in nEXO

- Alpha plating from recent SNOLAB paper:
  - <https://arxiv.org/abs/1708.09476>
- Neutron yields and energy spectra:
  - [NeuCBOT](#) instead of <http://neutronyield.usd.edu/>
- Geant4 instead of FLUKA for neutron transport and capture

Wonderful opportunity to collaborate with this community

- $(\alpha, n)$  backgrounds seem to be very sub-dominant for nEXO
- Estimates date back to 2015/2016, and should be considered a first look at the problem
- Calculations and simulations are being upgraded
  - Are we missing anything relevant?
  - What exposure to radon can we tolerate?
  - Do we need to assemble our detector in a radon-suppressed clean room?
- We can learn a lot from this group, and perhaps contribute to helping others as well!