

IFT – CIEMAT Meeting

Juan Garcia-Bellido (Virgo – IFT Madrid)

1st December 2022



IFT-Madrid Virgo group

Name	Teaching	FTE	Author	Student/Postdoc	Topical Fractions
Juan Garcia-Bellido (Professor)	Yes	60%	Yes	No	CBC (20%), SGWB (10%), Bursts (20%), Cosmo (%10)
Savvas Nesseris (Permanent Staff)	Yes	50%	Yes	No	Bursts (40%), CBC (10%)
Ester Ruiz Morales (Professor)	Yes	50%	Yes	No	CBC (40%), DA (10%)
Sachiko Kuroyanagi (Tenure-track Staff)	No	50%	Yes	No	SGWB (50%)
George Alestas	No	100%	Yes	Postdoc	CBC (100%)
Jose Francisco Nuño	No	100%	Yes	PhD student	INST (50%), DA (50%)
Gonzalo Morras	No	100%	Yes	PhD student	INST (50%), DA (50%)
Santiago Jaraba	No	100%	Yes	PhD student	INST (50%), DA (50%)

PBH, GW and Machine Learning group

Members:

Dr. Savvas Nesseris

Permanent staff member at IFT
Member of Euclid

Research interests:

Cosmology, dark energy,
modified gravity, machine learning,
gravitational waves,
primordial black holes.

Dedication: 50%

Responsible of the Machine Learning group of IFT:

<https://projects.ift.uam-csic.es/ift-ai-hpc/>

<http://members.ift.uam-csic.es/savvas.nesseris/>



Dr. George Alestas

Postdoc at IFT-UAM

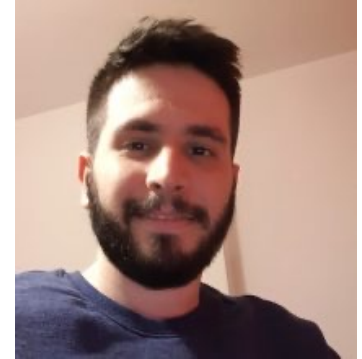
Experience:

Cosmology Group

Research interests:

gravitational waves, inflation

Dedication: 100%



CBC population analysis group

Members:

Prof. Ester Ruiz Morales

Professor of Physics at UPM
Associate Researcher at IFT

Experience:

Working groups for LHC physics
Collaboration with CMS (2001-2006)
Member of LISA

Research interests:

Cosmology, inflation, gravitational waves,
primordial black holes.

Dedication: 50%



Mr. José Francisco Nuño Siles

PhD student at IFT-UAM
4 year FPI fellowship at IFT

Supervisor:

Prof. Juan Garcia-Bellido

Research interests:

Gravitational waves, primordial black holes,
bayesian population analysis

Dedication: 100%



Bursts (hyperbolic encounters) analysis group

Members:

Prof. Juan Garcia-Bellido

Professor of Physics at IFT-UAM

Experience:

DESGW Collaboration (2005-)

Euclid Mission (2011-)

Member of LISA



Research interests:

Cosmology, Inflation, gravitational waves, primordial black holes, close hyperbolic encounters, Multimessenger, Tests of GR.

Dedication: 60%

Mr. Gonzalo Morrás

PhD student at IFT-UAM

4 year FPU Fellowship at IFT



Supervisor:

Prof. Juan Garcia-Bellido

Research interests:

Gravitational waves, primordial black holes.
Close hyperbolic encounters

Dedication: 100%

Stochastic GW Background analysis group

Members:

Dr. Sachiko Kuroyanagi

Tenure-track staff at IFT-UAM

Experience:

KAGRA Collaboration (2015-)
(Co-chair SGWB group)

Research interests:

gravitational waves, inflation, cosmic strings

Dedication: 50%



Mr. Santiago Jaraba

PhD student at IFT

4 year FPI fellowship at IFT

Supervisors:

Prof. Juan Garcia-Bellido

Dr. Sachiko Kuroyanagi

Research interests:

Gravitational waves, numerical relativity,
Close hyperbolic encounters

Dedication: 100%



IFT Virgo group

Human Resources:

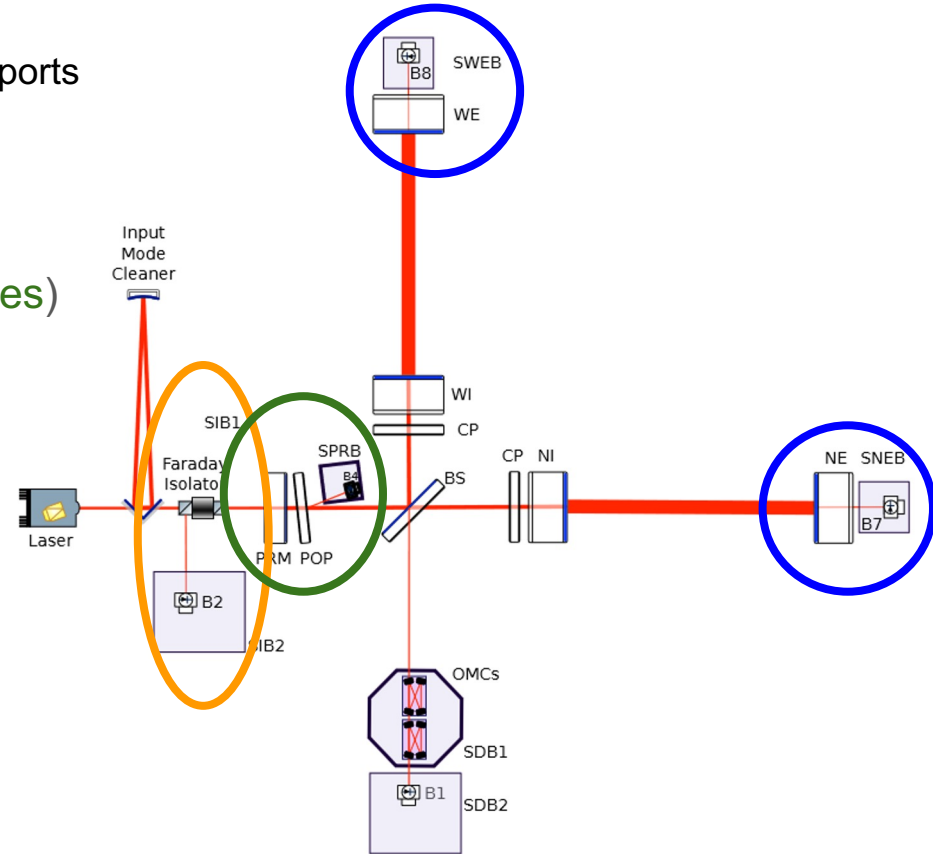
- 2.1 FTE Senior (4 people) with main background as Theoretical Cosmologists
Relevant expertise on Experimental Collaborations (DES, PAU, DESI, KAGRA, etc.)
- 4 FTE Junior (1 Postdoc + 3 PhD) willing to dedicate a significant amount of time to Commissioning, Data Acquisition and Detector Noise characterization.

Contributions to Working Groups in Virgo:

- INSTR: Commissioning run O4 & Instrumentation: Noise Characterization & Modelling
- CBC: Data Analysis (Bayesian Inference & Population Analysis). Multimessenger.
- Bursts: Data Analysis (Waveforms & Conv. Neural Networks). CHE Detection pipeline.
- SGWB: Data Analysis (Backg. charac., search pipeline for anisotropies & polarization.)

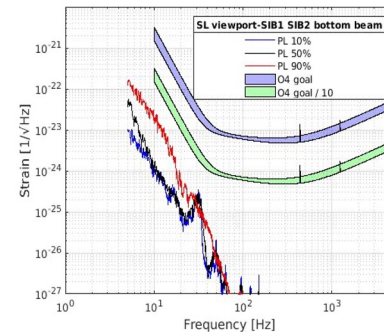
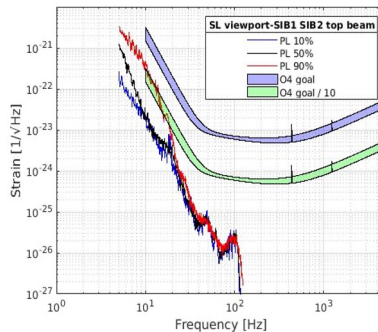
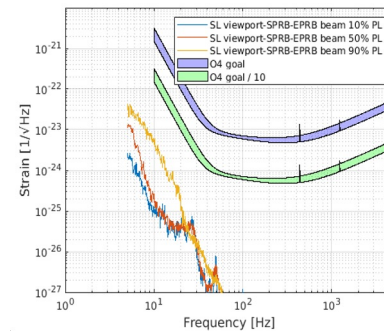
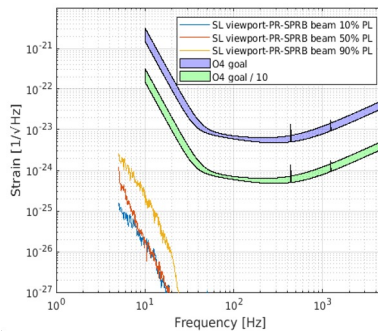
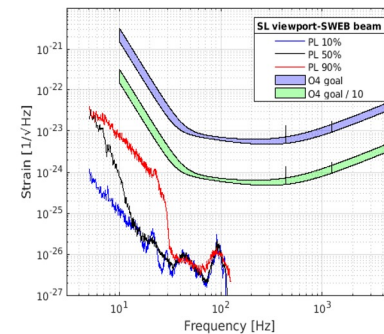
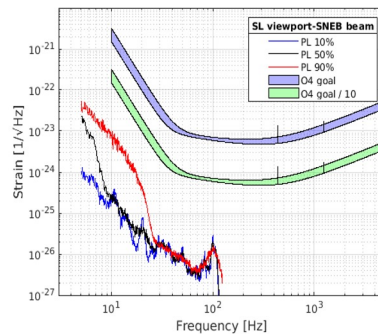
Virgo Commissioning - Technical Noise Investigation

- Task: Investigation of Scattered light noise from viewports
- Members (PhD students):
 - Santiago Jaraba (End benches)
 - Gonzalo Morrás (Power Recycling benches)
 - Jose Francisco Nuño (Injection benches)
- Supervisor: Maria Concetta Tringali
(with the help of Antonino Chiummo, Irene Fiori, Romain Gouaty & Michal Was)



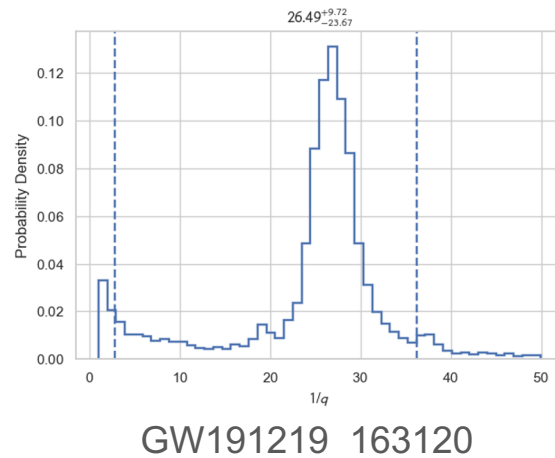
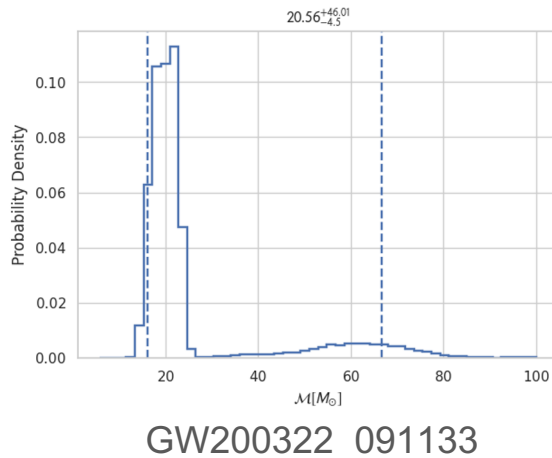
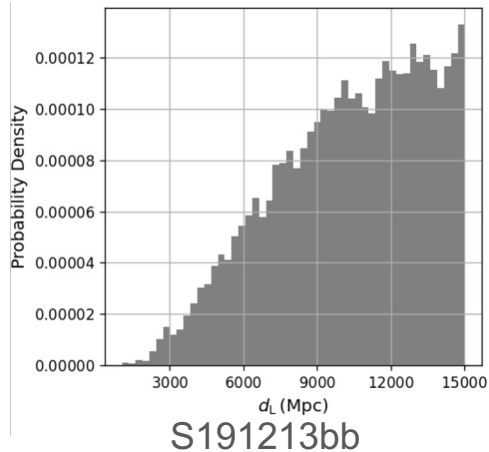
Virgo Commissioning - Technical Noise Investigation

- We obtained noise projections from the different viewports.
- Noise from End and Power Recycling benches:
 - Preliminary projections are below O4 noise budget.
 - It has already been presented in Technical Noise meeting: [VIR-0028A-22](#).
 - Scheduled tasks to perform noise injections **on-site** and validate noise projections.
- Noise from Injection benches:
 - Preliminary projections are above O4 noise budget. Investigations still ongoing.
 - **On-site** work scheduled.



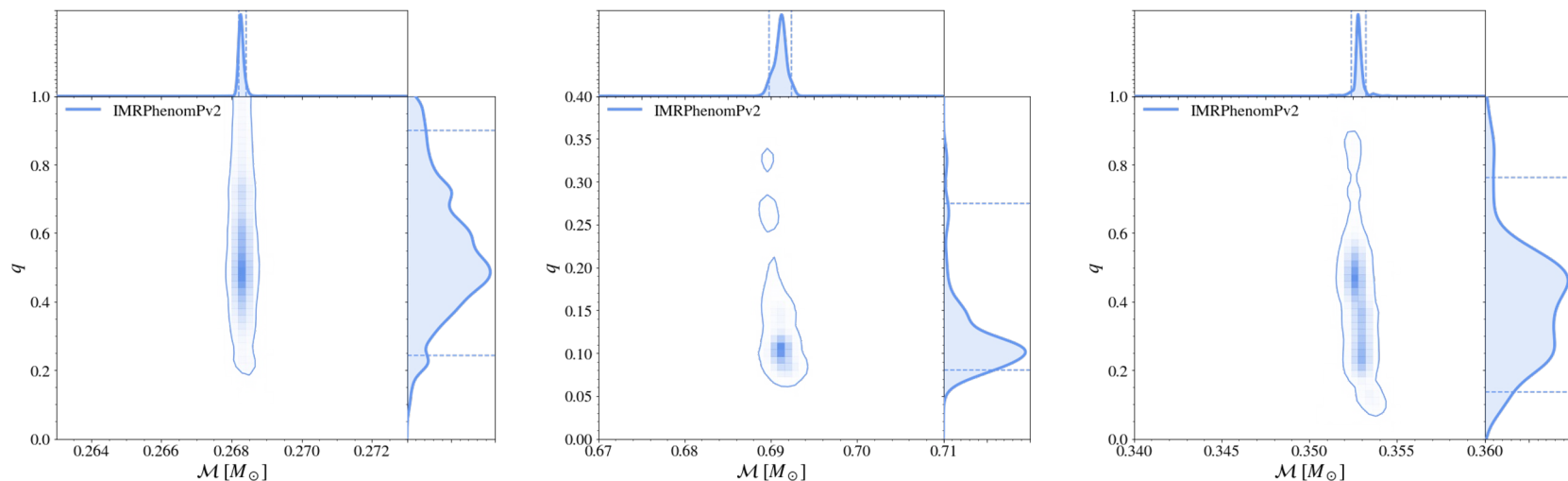
O3b Catalog PE

- Participants: Jose F. Nuño & Gonzalo Morrás
- As a part of the O3b Catalog LVK effort, we performed Preliminary PE on selected candidates. Namely:
 - cWB-only events as discussed in Appendix F of the Catalog Paper, in particular: [S191213bb](#), [S200218a](#), [S200224o](#), [S200326af](#)
 - One of the bimodal events [GW200322_091133](#)
 - Extreme mass-ratio event ($q = 0.038^{+0.005}_{-0.004}$) [GW191219_163120](#)



O3b SSM Candidates Follow Up

- Participants: Juan García-Bellido, Gonzalo Morrás, Jose F. Nuño & Ester Ruiz
- As a part of the O3b Subsolar Mass collaboration paper a follow up study including Parameter Estimation was performed on the most significant triggers and presented in the CBC call ([LIGO-G2200095](#))



SSM-O3b LVK paper Editorial Team

SSMO3b paper (Participants: Ester Ruiz Morales & Juan Garcia-Bellido)

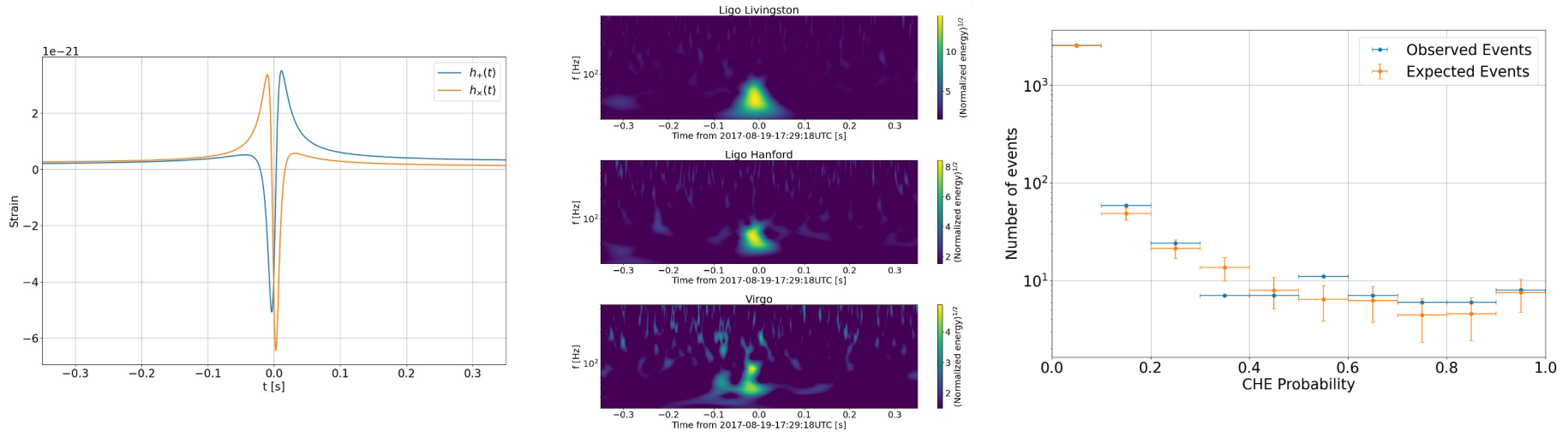
- Coordination of SMM O3b weekly sessions
- Review of different steps: Searches, Efficiencies, VT, Rates, Bounds
- Comparison with previous searches in Chirp mass.
- Extended mass models: Interpretation of event rates and bounds in (m_1, m_2) plane

Plans for SSM-O4

- 1.- For SSM searches (Participants: JGB, ERM, GM, JFN):
 - Understand the properties of noise and its role in SSM-O4 triggers.
 - Compute reliable FAR estimates of an event in the different pipelines.
 - Develop detection statistics techniques specific for SSM-O4 binaries.
- 2.- Follow up with PE only the most relevant SSM-O4 event candidates
- 3.- PBH's: Propose population studies based on thermal history model

Search for Close Hyperbolic Encounters (CHEs)

- Participants: Gonzalo Morrás, Juan García-Bellido & Savvas Nesseris
- In [arXiv:2110.08000](https://arxiv.org/abs/2110.08000) we developed a PN waveform for CHEs and used it to look for them in O2 data using an in-house Burst+CNN pipeline.
- Future plans: to make the waveform available to LVK with the new waveform interface & participate in the O4 eBBH+CHE collaboration search.



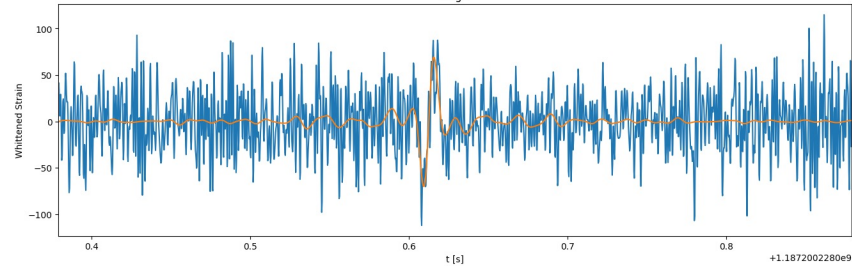
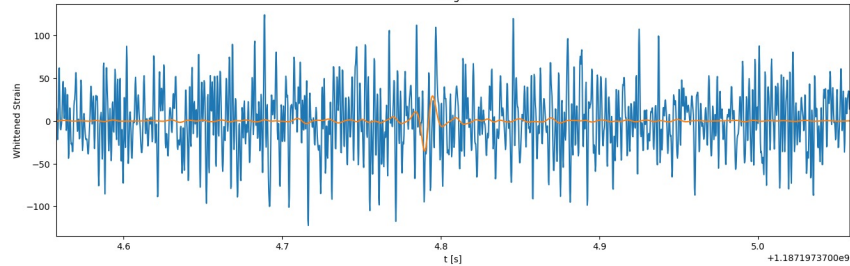
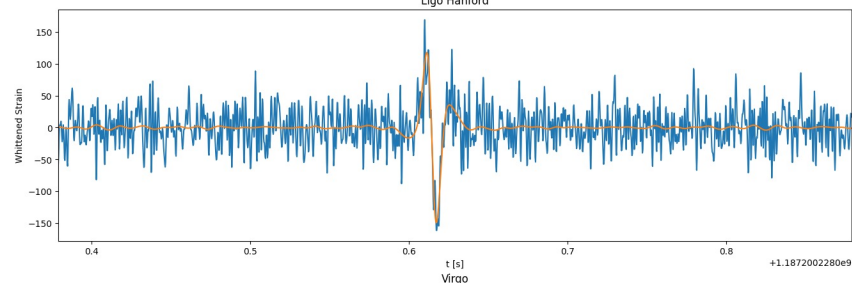
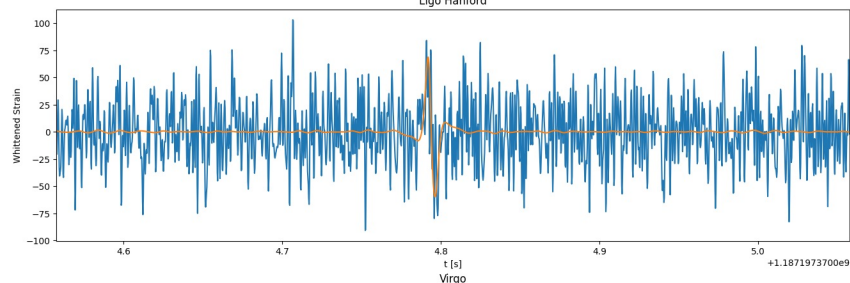
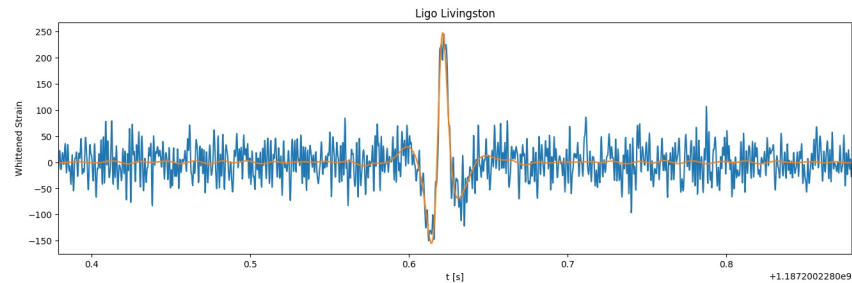
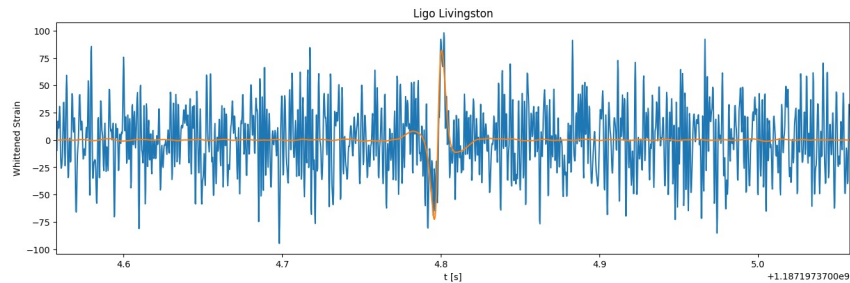
Examples CHE injected events (small & large SNR)

PPN-1.5 waveforms

De Vittori et al. (2014)

SNR = 5.0

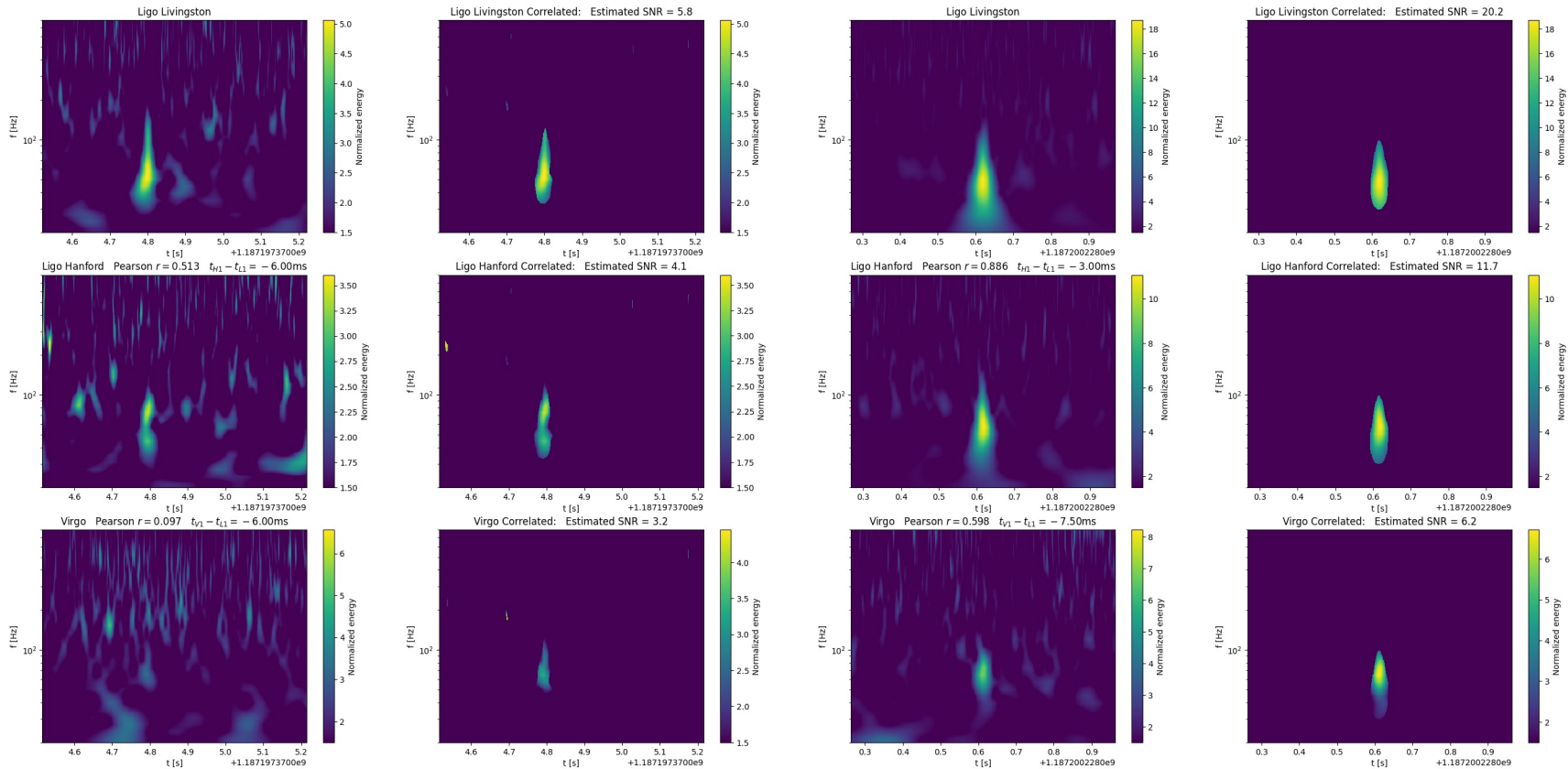
SNR = 16.0



Examples CHE injected events (small & large SNR)

Injected Event 2/10: $m_1 = 4.370M_\odot$ $m_2 = 0.392M_\odot$ $j_0 = 0.734$ $b_0 = 100GM$ $R = 1Mpc$
 $SNR_{L1} = 3.9$ $SNR_{H1} = 2.7$ $SNR_{V1} = 1.6$ $SNR_{tot} = 5.0$

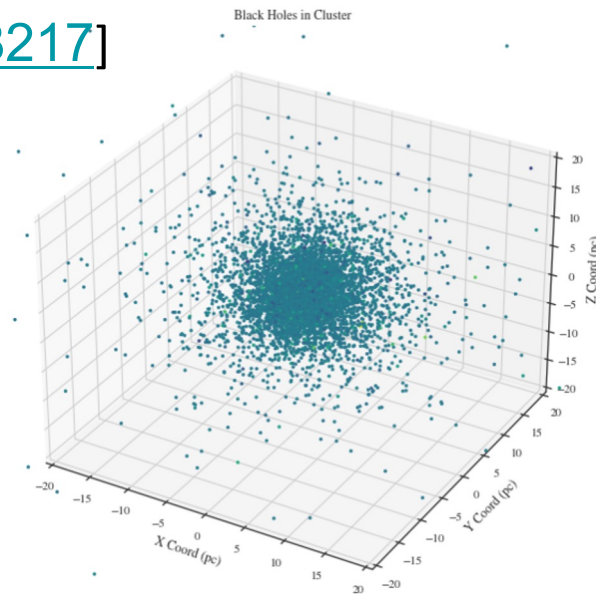
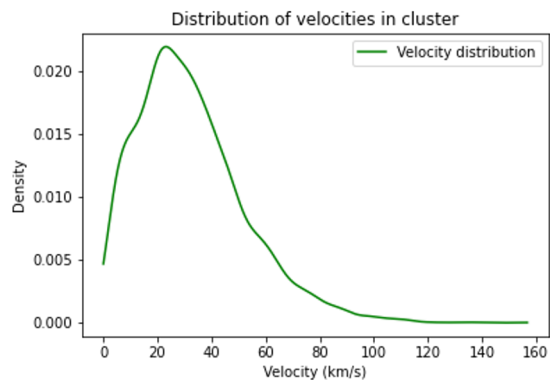
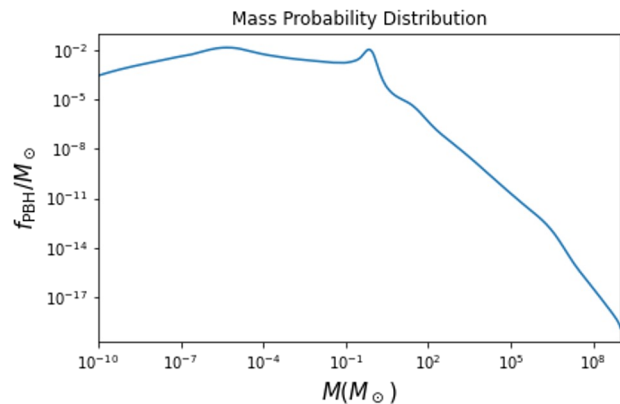
Injected Event 8/10: $m_1 = 5.898M_\odot$ $m_2 = 2.063M_\odot$ $j_0 = 0.780$ $b_0 = 100GM$ $R = 1Mpc$
 $SNR_{L1} = 13.4$ $SNR_{H1} = 7.6$ $SNR_{V1} = 4.4$ $SNR_{tot} = 16.0$



Numerical BH Cluster Simulation

(Jose Francisco Nuño Siles)

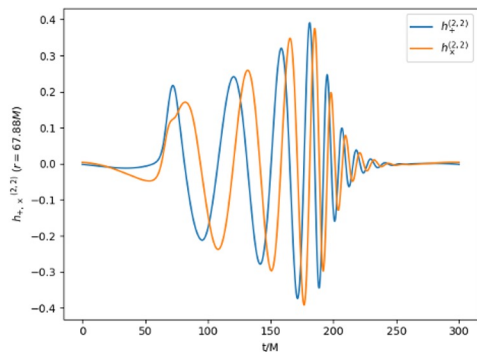
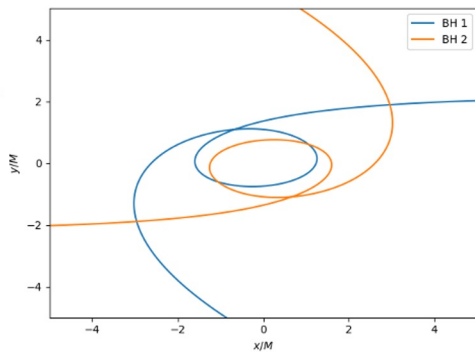
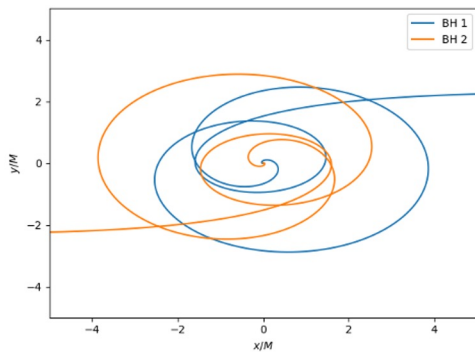
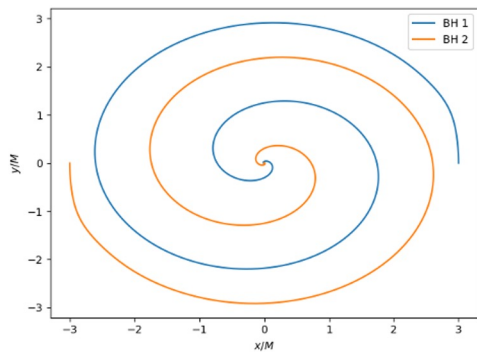
- Understand implications of PBHs conforming the totality or a fraction of the DM inside dense clusters in the outskirts of galaxies
- Observables: BBH Merger rate, CHE encounter rate, distribution of binary population parameters...
- [Nbody6++GPU](#) as integrator to be used and initial conditions based on a phenomenological thermal history model [[1906.08217](#)]



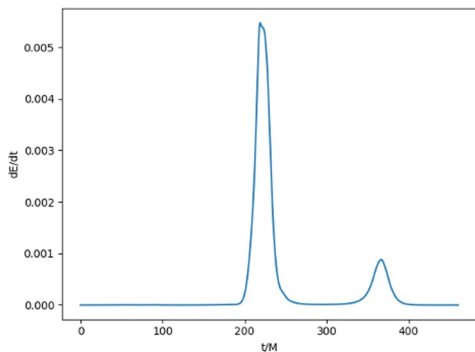
Numerical Relativity with Einstein Toolkit

(Santiago Jaraba)

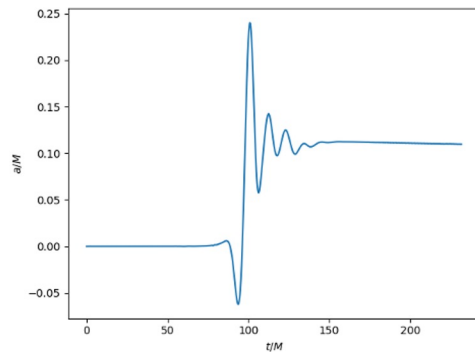
- Standard merger
- “Double” event
- Hyperbolic encounter



GW strain



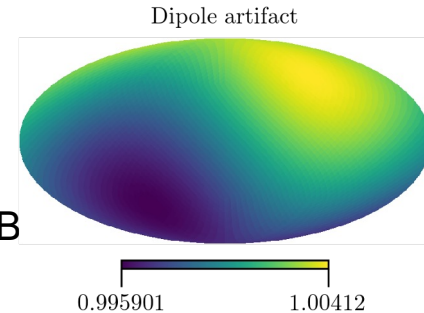
Radiated power



Induced spin

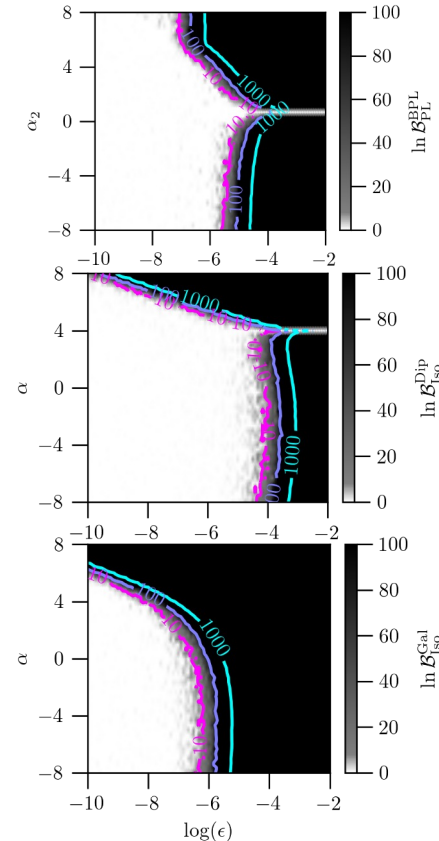
Santiago Jaraba: work within stochastic group

- Participation in Plm-targeted anisotropic PE formalism project, led by Leo Tsukada.
- Git repository: git.ligo.org/leo.tsukada/sgwb_sphpe
- Main contributions (details in [slides with project updates](#)):
 - Idea and implementation of formalism to speed up the code in a factor ~ 10 . A standard 1-hour run became a few minutes long.
 - Based on precomputing terms in advance to avoid these computations in each likelihood evaluation.
 - Great impact for everyone's runs in the project!
 - Derivation and implementation of dipole artifact due to Doppler shift on an isotropic SGWB, analogous to the one in the CMB



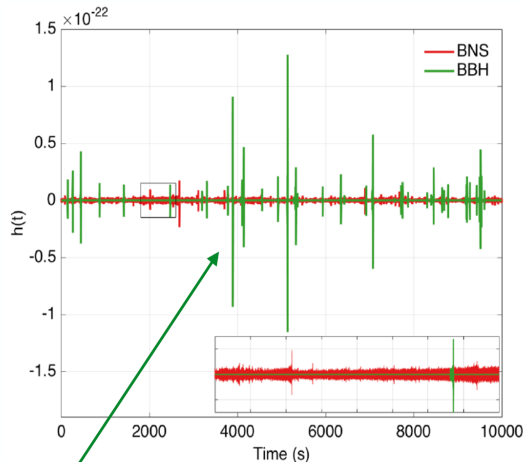
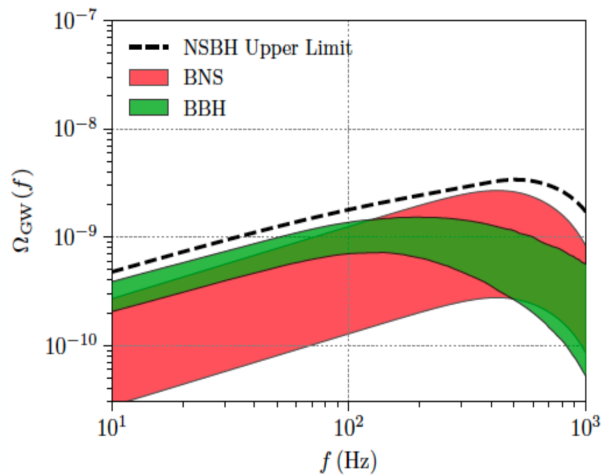
Santiago Jaraba: work within stochastic group

- Model selection analyses (is the code able to identify the model of the signal injected in the data?):
 - Broken power law vs power law spectrum.
 - Dipole artifact anisotropies vs isotropic.
 - Galactic plane anisotropies vs isotropic.
- More than 9000 PE runs per heatmap (plot of Bayes factor vs overall amplitude ϵ and power-law index α of injected signal).
- Results are ready, paper in preparation.
- Plans to contribute further within the recently formed Anisotropic stochastic subgroup.
- Ongoing projects with Juan, Sachiko and Matteo on SGWB.



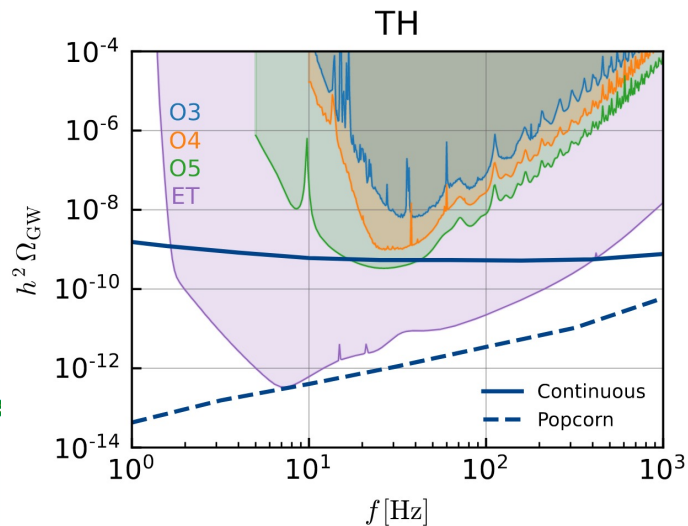
Detection method for a popcorn GW background

BBH events form the so-called popcorn GW background.



BBH events do not overlap one another

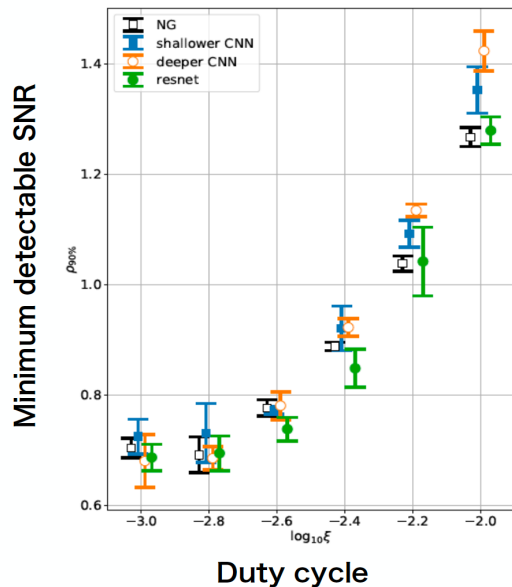
(Sachiko Kuroyanagi,
Juan Garcia-Bellido
& Matteo Braglia)



Detection methods considered in the Stochastic group

1. CCI (Cross-Correlation search for Intermittent backgrounds)
2. TBS (The Bayesian Search)
3. Machine Learning

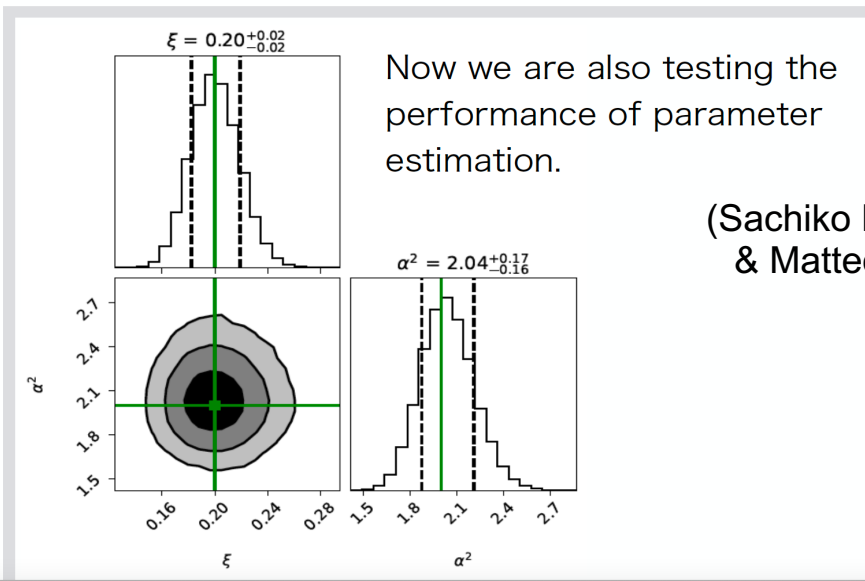
Machine Learning for a popcorn GW background



We apply convolutional neural network (CNN) to analyze a popcorn background.

Three different CNN algorithms are tested for toy-model data. The Resnet method performs as good as in method 1. We also found that it remarkably reduces computation time.

Method	Speed-up factor
Maximum likelihood	1
Shallower CNN	1.6×10^5
Deeper CNN	4.8×10^4
Residual network	5.9×10^4



(Sachiko Kuroyanagi & Matteo Braglia)

IFT Virgo Group

Summary:

- 2.1 FTE Senior (4 people) Theoretical Cosmologists; expertise in large Collaborations
- 4 FTE Junior (1 Postdoc + 3 PhD) Commissioning & Detector Noise characterization.

Synergies with other Spanish Groups in Virgo Spain (CIEMAT, Valencia, IFAE, ICCUB).

Synergies with other Groups in Virgo (Brussels, NIKHEF, Lyon).

Participation in both LSS & GW Collaborations provides Multimessenger lever arm.

Members of IFT Center of Excellence with many connections worldwide.