Cryogenic setup for the characterization of wavelength-shifting materials for noble element radiation detectors

Contact: <u>acortez@camk.edu.pl</u> | <u>andre.f.cortez@gmail.com</u>

ASTROCENT

S. Choudhary¹ | <u>A.F.V. Cortez¹ | M. Kuźniak¹ | G. Nieradka¹ | T. Sworobowicz¹ Ł. Świderski² | T. Szczęśniak²</u>

¹ AstroCeNT, Nicolaus Copernicus Astronomical Center of the Polish Academy of Sciences, 00-614 Warsaw, Poland ² National Center for Nuclear Research, 05-400 Otwock - Świerk, Poland

Challenges in light collection for Dark Matter searches

Achieving high light collection efficiency in noble element detectors is critical as it crucially affects: energy threshold, position and energy reconstruction, pulse shape discrimination (essential for particle ID) and triggering. For efficient detection, vacuum UV (VUV) scintillation wavelengths of noble elements typically require wavelength shifters [1].

With currently available technologies, scaling up / enhancing light collection in large noble liquid detectors for dark matter or rare event searches poses a major challenge [2,3]. The presented setup provides a simple and robust way to test performance of wavelength shifters (and other approaches) in representative conditions.

Main approaches: optical [4] and electromagnetic lenses [5], reflective and wavelengthshifting passive coatings [1], wavelength-shifting using dissolved dopants or gas mixtures [4], some dedicated optical amplification structures (M-THGEM, G-THGEM, FAT-GEM) [6].

- Identify **low background**, **radiopure** alternatives (improve PDE)
- Scalable and stable solution (compared to meshes/wires)
- Wavelength-shift primary and secondary scintillation (easier to readout),
- Low-background, low-noise cryogenic amplification structures and SiPMs,
- Improve the electron extraction efficiency at the liquid-gas interface

ArGSet – A Cryogenic Setup for Argon wavelength-shifting testing at AstroCeNT

Cryogenic Setup

To test the wavelength-shifting capabilities of new materials a dedicated cryogenic setup was recently commissioned and tested at AstroCeNT.

This setup can provide essential information on:

- Time response and efficiency of WLS materials;
- Temperature dependence of WLS materials;
- \succ Compatibility of materials with cryogenic operation;
- \succ SiPM performance in cryogenic operation;



Wavelength-shifter studies with ArGSet

What

The commissioning measurements revealed <5% reproducibility of the wavelength shifting efficiency measurement, with 50% efficiency of PEN relative to TPB -- consistent with independent result for the same PEN batch. Systematic uncertainties are being studied.





Table 1 – Comparison of the results obtained with ArGSet with previous works or litereature.

	WLS efficiency (PEN/TPB)	Triplet lifetime constant
Current work	50%	1.24 µs
Literature*	47.2±5.7%	3.5 μs

* For the WLS efficiency 2PAC results [7] and triplet lifetime constant the MiniCLEAN results [8].

Next steps:

Study the influence of impurities on the

In order to enable **ns timing-resolution** for the characterization of the WLS materials, primary scintillation (S1) from Ar is used as source. To produce a measureable signal ²⁴¹Am source is used (alpha particles). VUV photons will impinge on the sample material (WLS material to be tested). Resulting WLS light is then detected by a set of 2 independent SiPMs not sensitive to the VUV emission spectra of Ar.



ArGSet – A tool for testing wavelength-shifting FAT-GEMs*

An extension to ArGSet's current design is ongoing. This extension will allow testing different configurations and manufacturing processes used in the development of wavelength-shifting FAT-GEMs [2,9] for their use in cryogenic environment.

Features:

- Modular structure made of cryogenic and radiopure materials (easy to change between WLS material studies and testing of FAT-GEMs);
- Independent biasing of all internal structures (cathode, FAT-GEM and induction gap mesh);



Fig. 1 – ArGSet extension to test wavelength-shifting FAT-GEMs.

Fig. 2 – Working principle of WLS FAT-GEMs.

Multipliers (FAT-GEM) were introduced by Diego González-Diaz (IGFAE) [10]. Fig. 3 – Detail schematic of WLS FAT-GEMs.

Conclusions and Future work

A cryogenic setup (ArGSet) to study the wavelength-shifting performance of potential interesting materials for the development of novel WLS FAT-GEMs was commissioned and successfully tested at Astrocent. This setup allows to:

- Study the time response of WLS materials;
- Temperature and pressure dependence of WLS materials performance;
- Test new materials in cryogenic operation;
- Test SiPMs performance.

ArGSet will allow the testing of promising radiopure, and scalable structures recently developed by Astrocent and IGFAE, such as the WLS FAT-GEM.

ArGSet extension will open the possibility to have more detailed understanding of these structures under cryogenic operation.

References

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