



University of
Zurich^{UZH}



XENONnT status and recent results

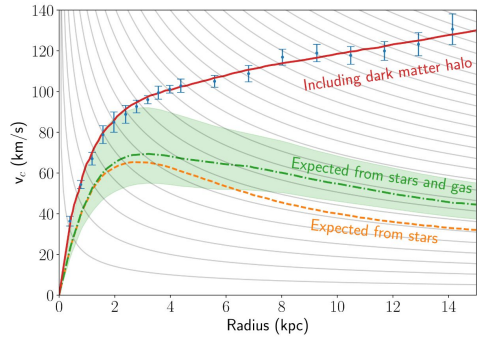
Diego Ramírez (UZH) on behalf of the
XENON collaboration @ LIDINE2023

September 20th 2023 | Madrid, Spain



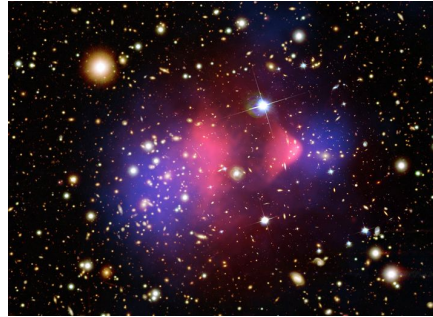
Particle dark matter

Galactic scale (rotation curve M33)



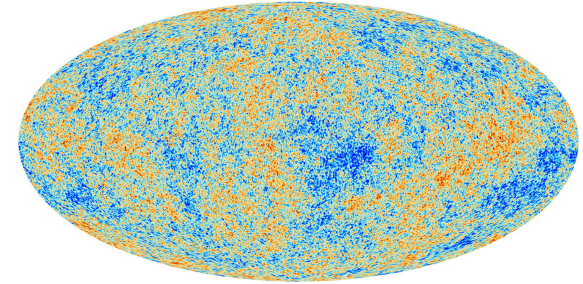
[E. Hogenbirk](#)

Extragalactic scale (Bullet Cluster)



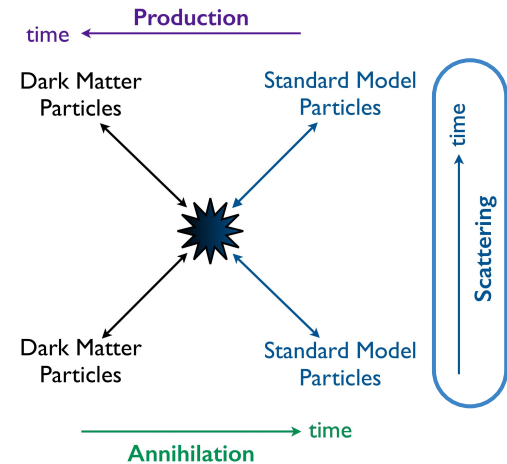
[NASA](#)

Universal scale (CMB)



[A&A 594 \(2016\) A1](#)

- Dark matter existence suggested by astrophysical and cosmological observations at all scales
- Weakly interacting massive particle (WIMP) as prime candidate
- The XENON collaboration aims at direct WIMP detection using dual-phase xenon time projection chambers



XENON collaboration

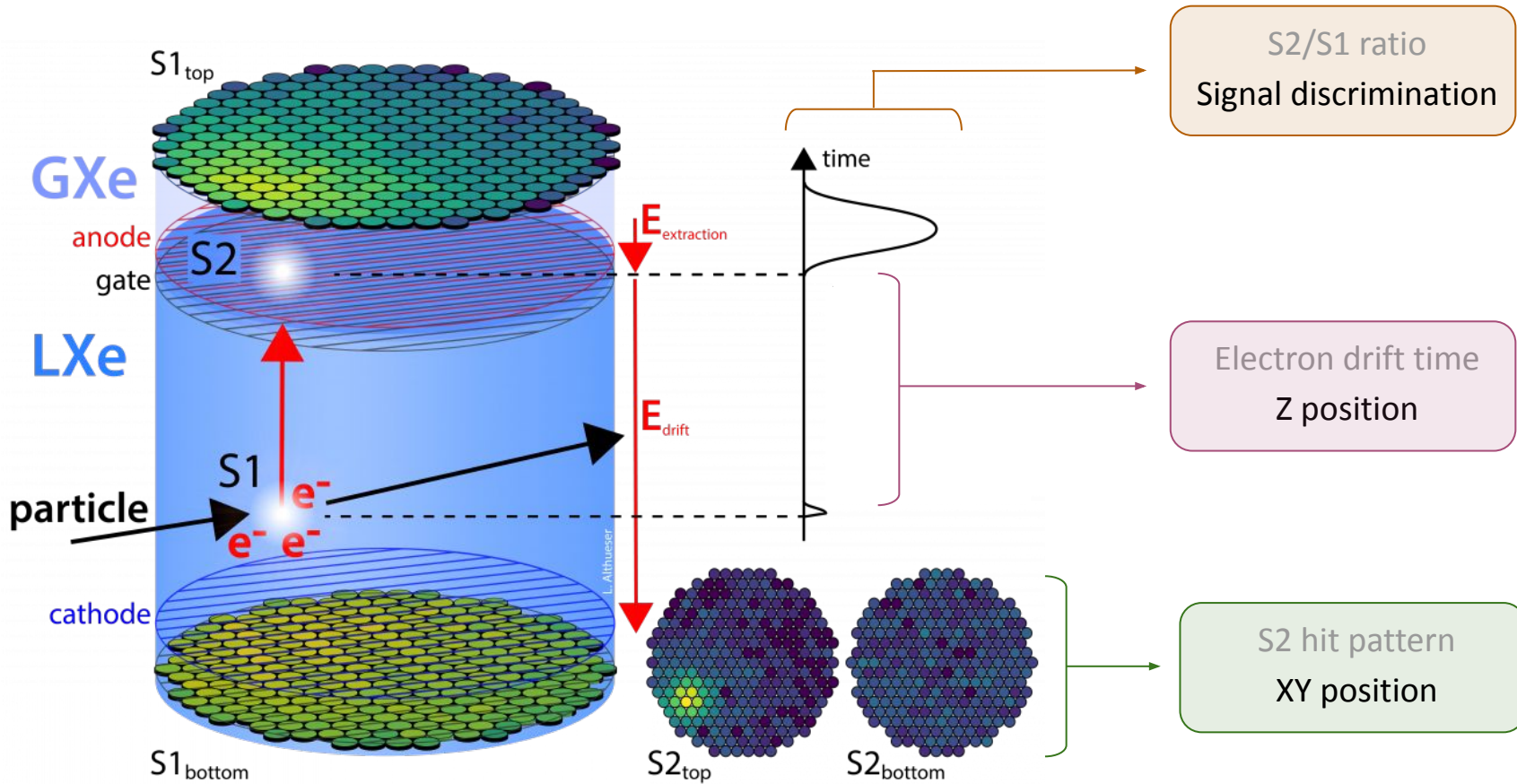
Dark matter direct detection experiment at
Laboratori Nazionali del Gran Sasso (LNGS)



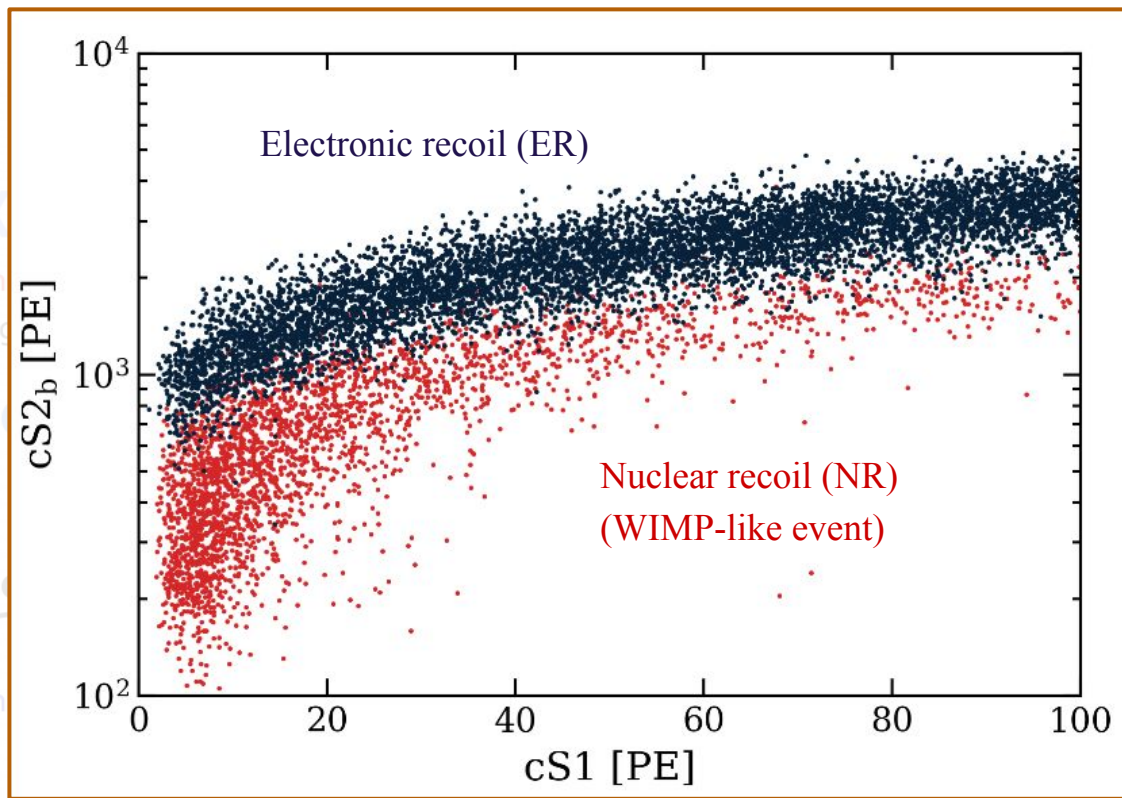
Collaboration meeting @ L'Aquila, February 2023

12 countries
28 institutions
~ 170 scientists

Xenon dual-phase time projection chamber



Xenon dual-phase time projection chamber



XENON1T calibration data

S2/S1 ratio
Signal discrimination

Electron drift time
Z position

S2 hit pattern
XY position

S1_{bottom}

S2_{top}

S2_{bottom}

From XENON1T to XENONnT

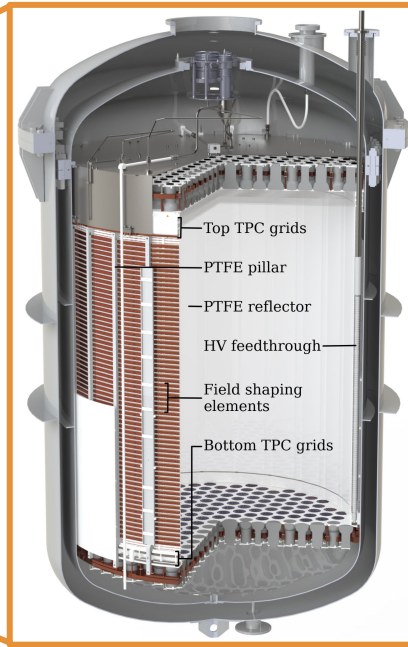
@ LNGS (Italy), 3600 m.w.e. rock



From XENON1T to XENONnT

@ LNGS (Italy), 3600 m.w.e. rock

Time projection chamber (TPC)

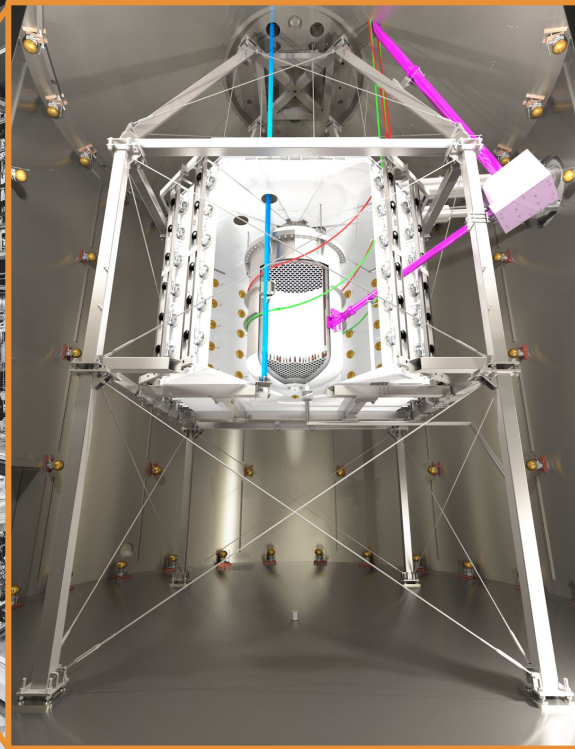


- 1.3 m (\varnothing) \times 1.5 m
- 8.5 t LXe in cryostat (2.5x XENON1T)
- 5.9 t LXe active (3x XENON1T)
- 494 (3") PMTs (2x XENON1T)
- Five electrodes made of SS wires
- Two sets of concentric field-shaping rings, top one individually tuneable

From XENON1T to XENONnT

@ LNGS (Italy), 3600 m.w.e. rock

Neutron veto



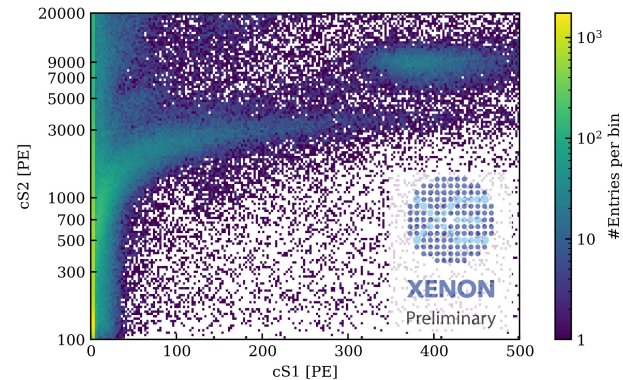
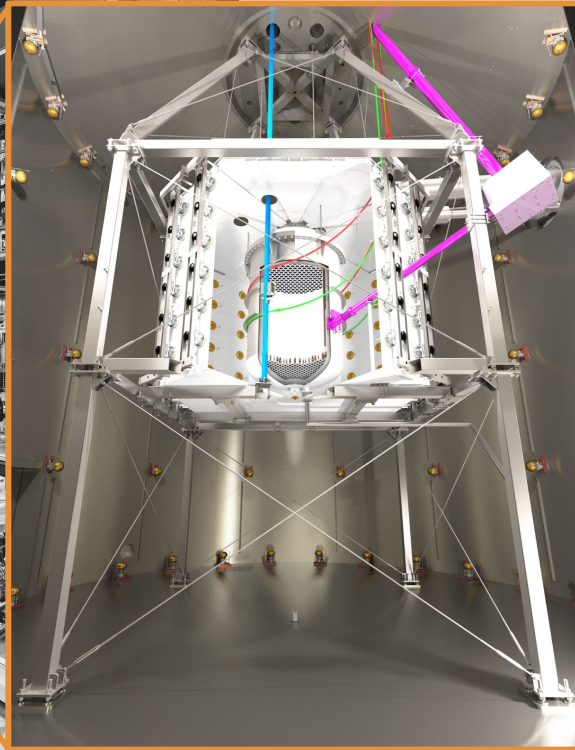
- 120 (8") PMTs facing the TPC detect Cherenkov light from 2.2 MeV gamma (n-capture on H)
- (53 ± 3) % tagging efficiency (250 μ s window, 5-fold PMT coincidence) with 1.6 % lifetime loss
- 0.2 % Gd doping will improve it to ~ 87 % (150 μ s, 10-fold) [JCAP 11 031 \(2020\)](#)
- Tagged calibration neutrons to study NR LXe response

From XENON1T to XENONnT

@ LNGS (Italy), 3600 m.w.e. rock

Neutron veto

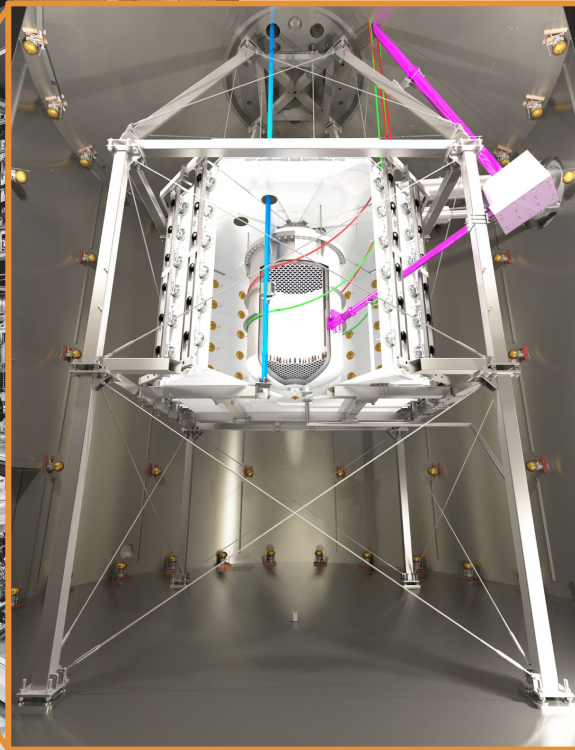
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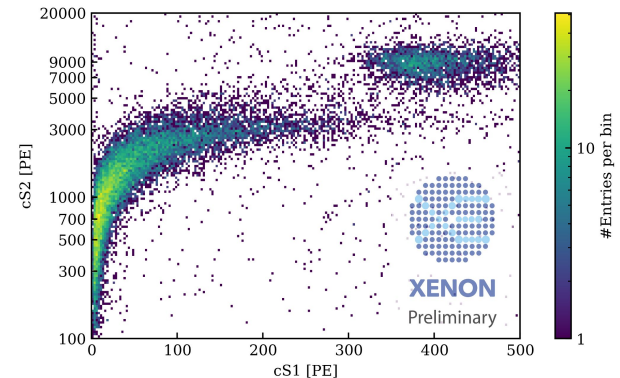
From XENON1T to XENONnT

@ LNGS (Italy), 3600 m.w.e. rock

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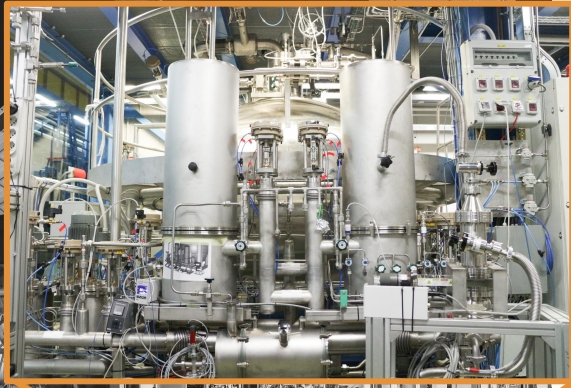


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From XENON1T to XENONnT

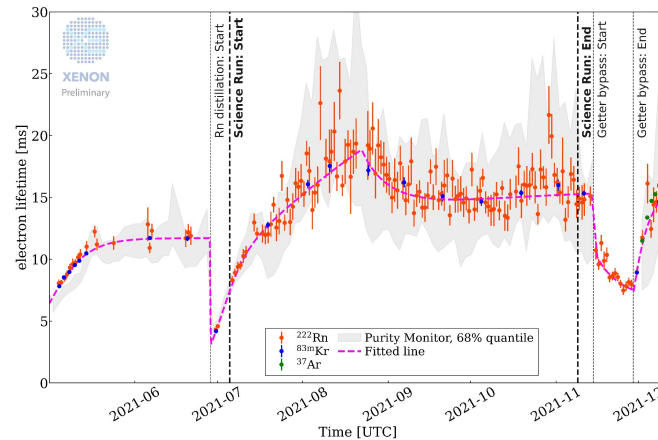
@ LNGS (Italy), 3600 m.w.e. rock



Liquid xenon purification

[Eur. Phys. J. C 82, 860 \(2022\)](#)

- LXe purity is crucial for electrons to survive until liquid-gas interface
- Novel liquid-phase purification with replaceable filter units, some with extremely low radon emanation (science run mode)
- 2 liters of LXe per minute: 18 h to recirculate entire inventory



	Max. TPC drift time	Electron lifetime	e ⁻ survival @ max. drift length
1T	0.67 ms	0.65 ms	30%
nT	2.2 ms	> 10 ms	> 90%

From XENON1T to XENONnT

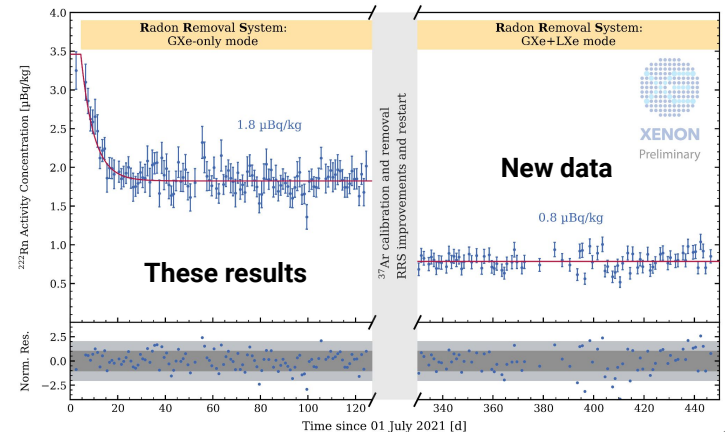
@ LNGS (Italy), 3600 m.w.e. rock

Radon distillation column [Eur. Phys. J. C 82, 1104 \(2022\)](#)

- “Online” removal of emanating ^{222}Rn ($t_{1/2} = 3.8$ d) exploiting difference in vapor pressure
- ^{222}Rn Activity concentration equilibrium of 1.77 ± 0.01 $\mu\text{Bq/kg}$ with gas extraction only (~ 13 $\mu\text{Bq/kg}$ in XENON1T)
- $< 1\mu\text{Bq/kg}$ with gas+liquid extraction in current data-taking



Xe
Rn

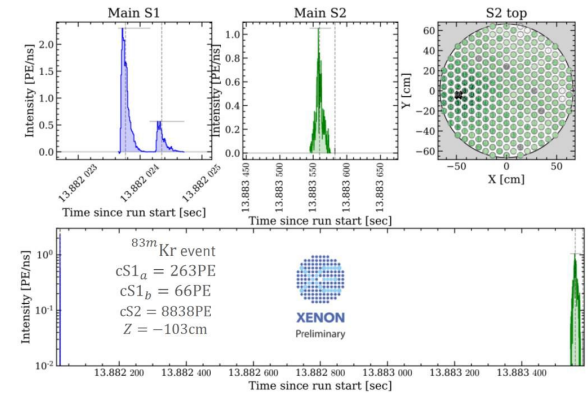
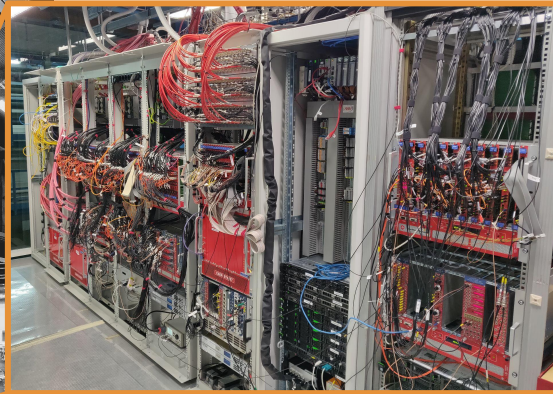
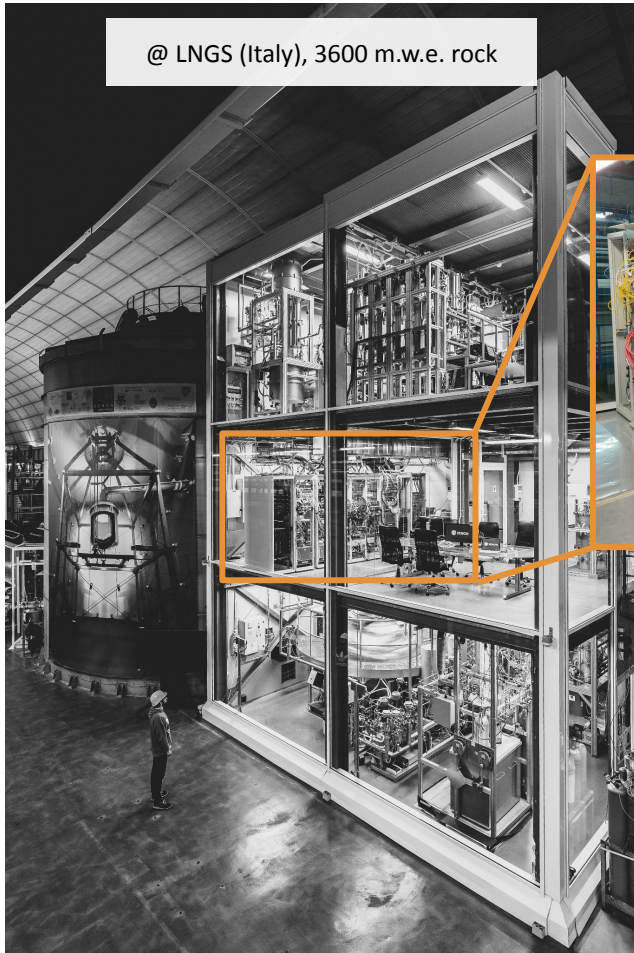


From XENON1T to XENONnT

@ LNGS (Italy), 3600 m.w.e. rock

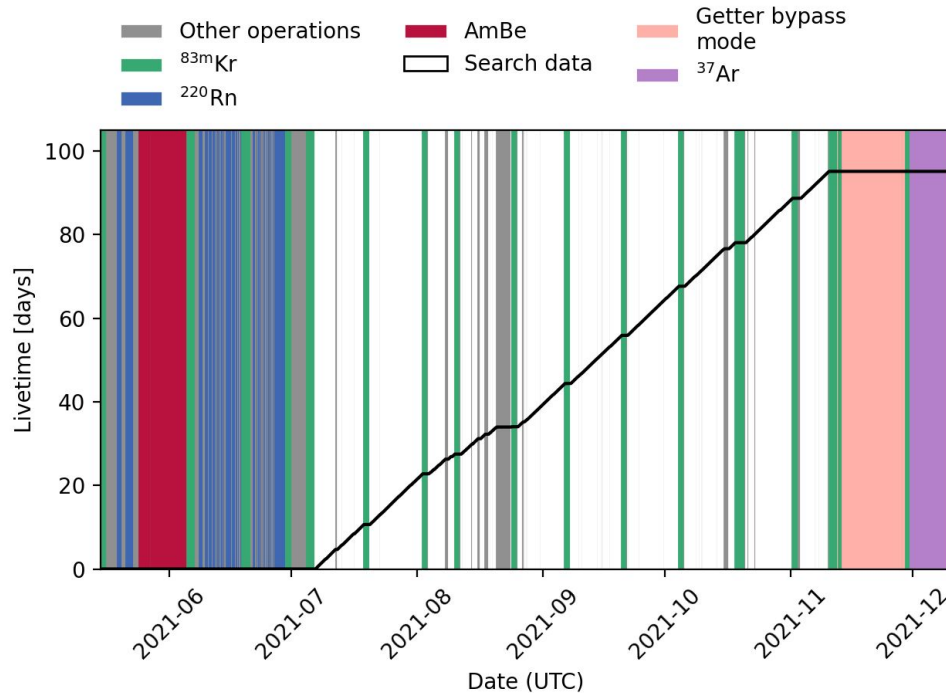
Triggerless data acquisition

[JINST 18 P07054 \(2023\)](#)



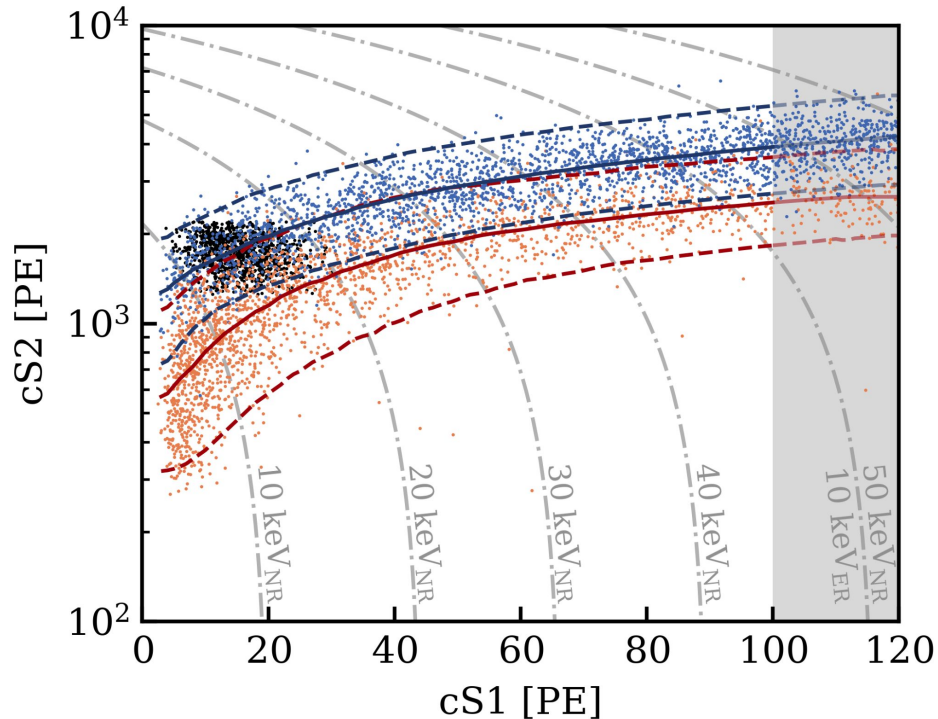
- All data above per-channel threshold stored long term
- Fully live processing
- Open source processing software:
[strax@github](#) (base)
[straxen@github](#) (XENONnT-tailored)

First XENONnT science run (SR0)



- **97.1 days of exposure** from July 6th - Nov 11th 2021 (95.1 days lifetime corrected)
- Radon column operating in gas-only mode
- 477/494 PMTs operative, gain variations below 3% level
- 23 V/cm drift field (cathode limited to -2.75 kV due short-circuit with bottom screen mesh)
- Extraction field in LXe 2.9 kV/cm
- Localized high single-electron emission sporadically, anode ramped down

Detector calibration



ER response model (combined fit)

^{212}Pb from ^{220}Rn gives a roughly flat β -spectrum, to estimate cut acceptances and energy threshold

^{37}Ar gives a mono-energetic 2.82 keV peak, to model low-energy response and resolution near detector energy threshold

NR response model

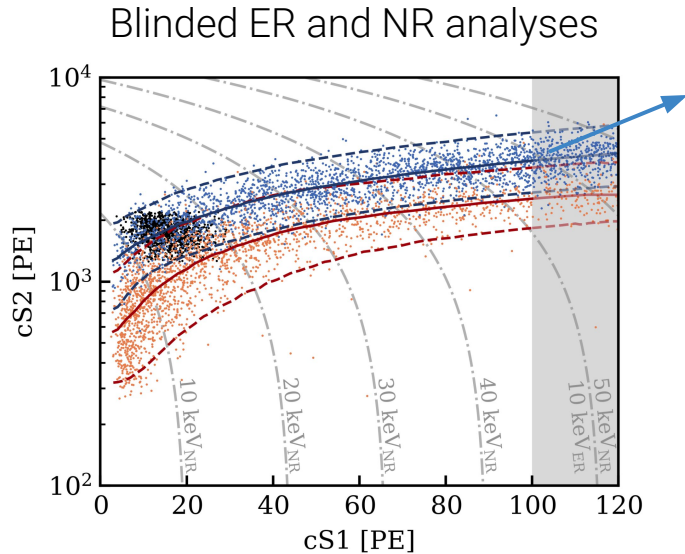
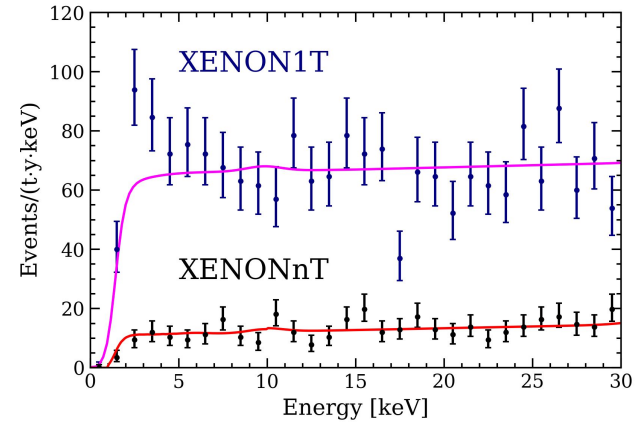
$^{241}\text{AmBe}$ external source emits neutrons which are tagged using the coincident 4.4 MeV γ -ray observed in the neutron veto

New physics searches

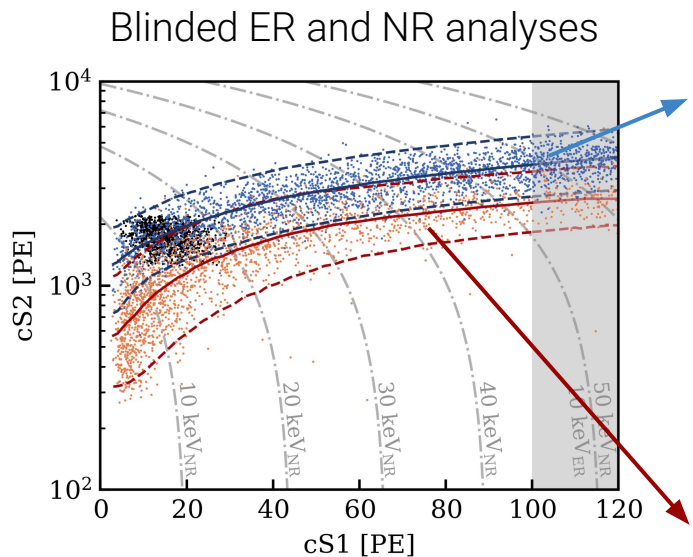
[PRL 129, 161805 \(2022\)](#)

Electronic recoils

- Combine S1 and S2 signals for search in reconstructed energy
- Lowest ever ER background in the field: $(15.8 \pm 1.3) \text{ (t yr keV)}^{-1}$
- No low-energy ER excess found (beyond-SM explanation for XENON1T excess excluded)



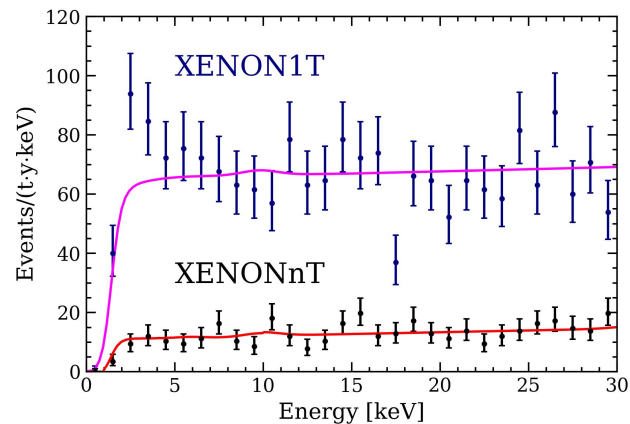
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[PRL 129, 161805 \(2022\)](#)

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[PRL 131, 041003 \(2023\)](#)

Nuclear recoils (this talk)

- Much lower NR background thanks to ER/NR discrimination
- Search in S1, S2 and radius
- 4.18 t fiducial mass (out of 5.9 t), total exposure of 1.1 t yr

Detection and selection efficiencies

Detection efficiency

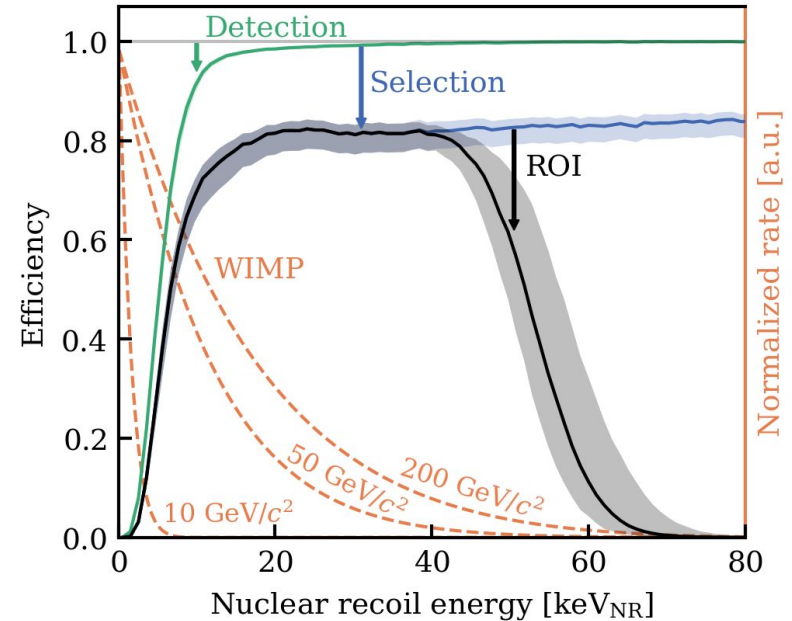
- Driven by 3-fold PMT coincidence requirement to identify an S1
- Data- and simulations-driven, with good agreement

Selection criteria (~ 80 % acceptance)

- S1 and S2 peaks must have PMT patterns, top/bottom area ratios, etc. consistent with real events
- S2 width consistent with the expected diffusion of electrons along drift path
- Single-site (signal-like) events

Region of interest (ROI)

- Constructed to fully contain the WIMP (GeV-scale) recoil spectra
- cS1: 0 - 100 PE | cS2: $10^{2.1}$ - $10^{4.1}$ PE



Backgrounds in the WIMP search

Electronic recoil (ER)

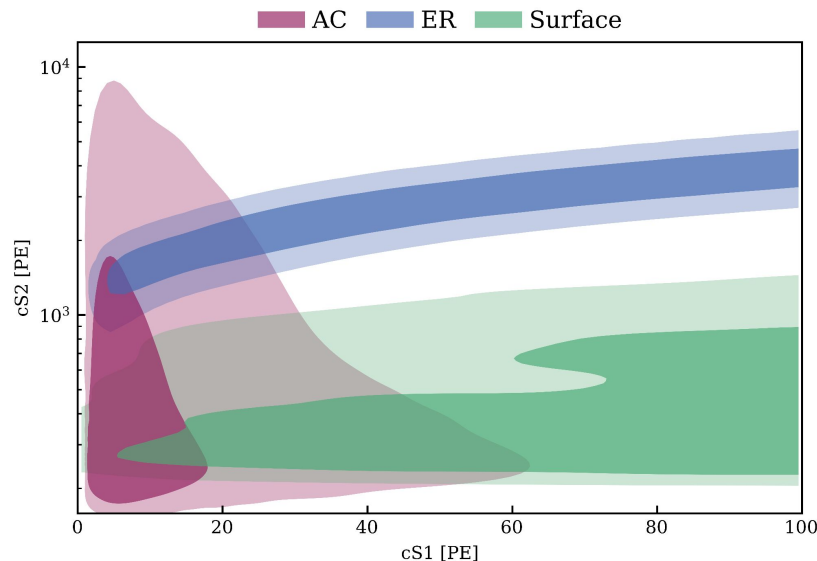
Mainly β -decay of ^{214}Pb

Accidental coincidence (AC)

Random pairing of isolated S1 and S2 signals; suppressed by GBDT cut based on S2 shape

Nuclear recoil (NR)

Radiogenic neutrons constrained by neutron veto tagging (~ 1.1 events), CEvNS less than 0.2 events due to decreased efficiency at low energies



Surface background

^{210}Pb plate-out on TPC walls, leading to ^{210}Po α -decays with electron loss; suppressed by volume fiducialization

Backgrounds in the WIMP search

Electronic recoil (ER)

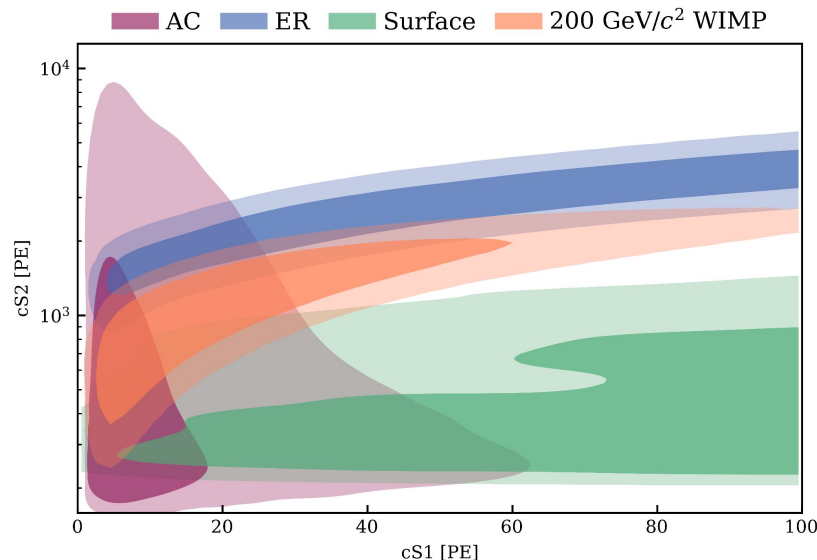
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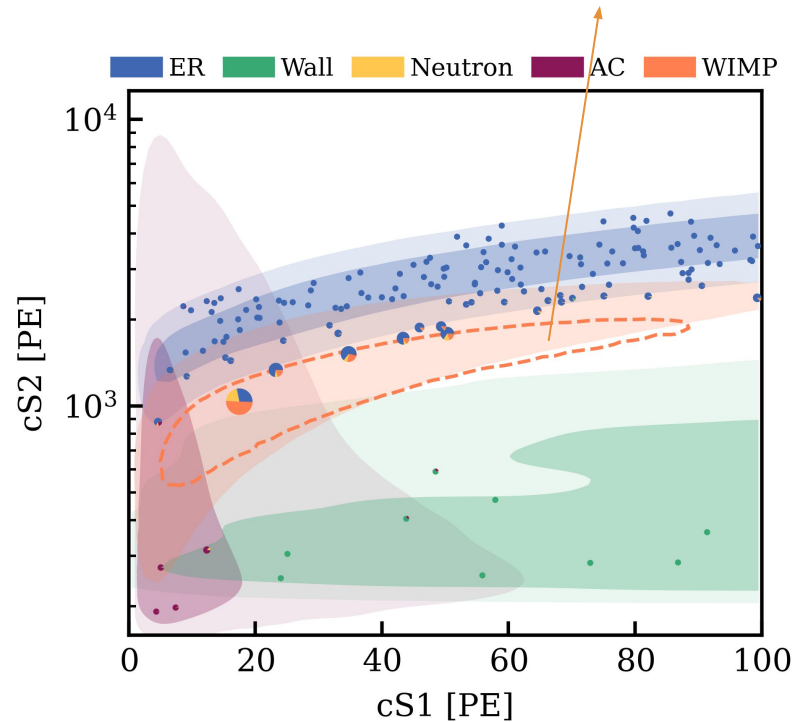


Surface background

^{210}Pb plate-out on TPC walls, leading to ^{210}Po α -decays with electron loss; suppressed by volume fiducialization

WIMP search results

Signal-like region containing 50% of WIMP signal with lowest background



200 GeV/c²

	Expected	Best fit
ER	134	135 (+12) (-11)
Neutrons	1.1 (+0.6) (-0.5)	1.1 ± 0.4
CEvNS	0.23 ± 0.06	0.23 ± 0.06
AC	4.3 ± 0.2	4.32 ± 0.15
Surface	13 ± 3	12 (+0) (-4)
Tot. background	154	152 ± 12
WIMP	-	2.6
Observed	-	152

No significant excess, 152 events in ROI, 16 in blinded region

WIMP-nucleon cross section limit

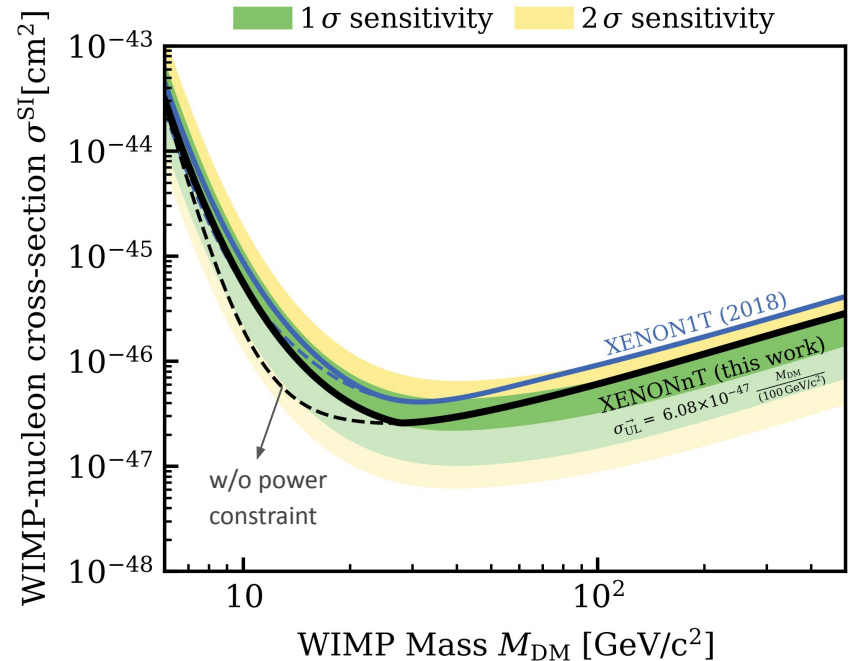
Limit setting

- Unbinned maximum likelihood
- Power-constrained limits (PCL) based on rejection power to avoid spurious exclusion limits
- Conservative choice at median of sensitivity band (pending renewed agreement with community)

Strongest limit: $2.6 \cdot 10^{-47} \text{ cm}^2$ at WIMP mass of **28 GeV/c²**

Factor 1.6 improvement w.r.t. XENON1T (with considerably shorter lifetime)

Similar improvements in spin-dependent limits



WIMP-nucleon cross section limit

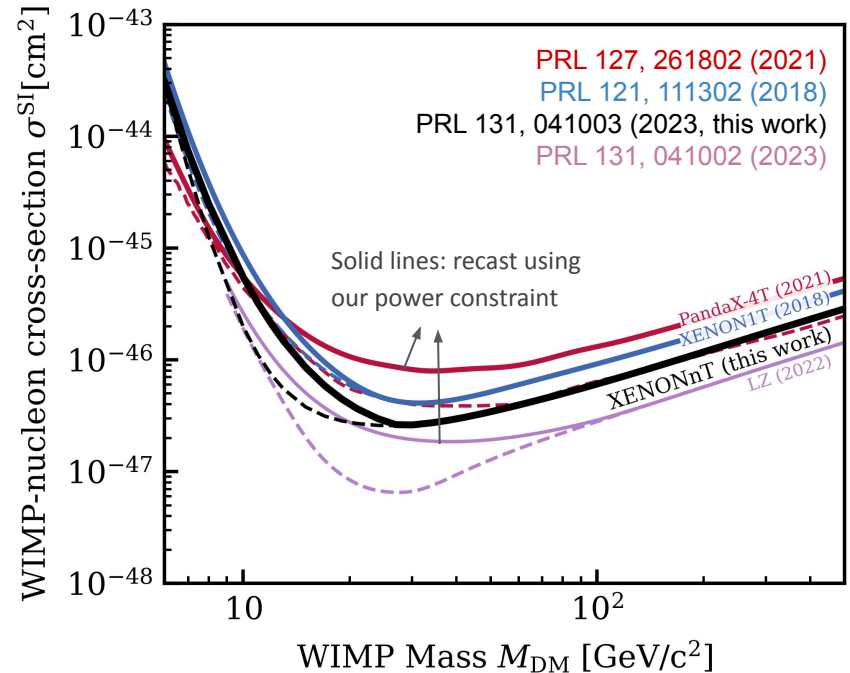
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Summary and outlook

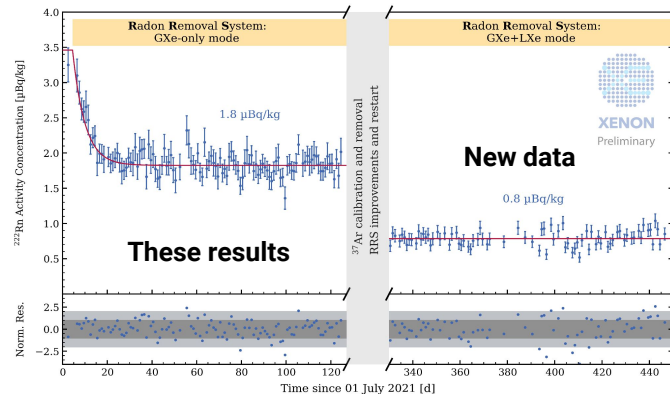
XENONnT SR0:

- Ultrapure target, with an electron lifetime stably above 10 ms
- Lowest ER background in the field: $(15.8 \pm 1.3) \text{ (t yr keV)}^{-1}$ (~ 5x background reduction w.r.t. XENON1T)
- First blinded searches for electronic and nuclear recoil signals yielded no significant excess over background

Taking more data with:

- 50 % lower ^{222}Rn level by changing flow path
- Planned Gd-loaded neutron veto with ~ 87 % neutron background tagging efficiency

Better WIMP limits, solar neutrinos, double weak decays and more...



Gd plant and mixing of Gd salt in water



xenonexperiment.org



[xenon_experiment](https://www.instagram.com/xenon_experiment)



[xenonexperiment](https://twitter.com/xenonexperiment)

Summary and outlook

XLZD meeting @ Karlsruhe, June 2022



XLZD consortium

Joining forces towards a next-generation dark matter experiment

(white paper: [J. Phys. G 50 \(2023\) 013001](#))



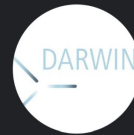
XENON

Currently operating with 8.5 tonnes of liquid Xenon at Gran Sasso in Italy



LUX-ZEPLIN

Currently operating with 10 tonnes of liquid Xenon at SURF in South Dakota



DARWIN

Leading many R&D projects designing a future 50 tonnes liquid Xenon detector



xenonexperiment.org



[xenon_experiment](#)



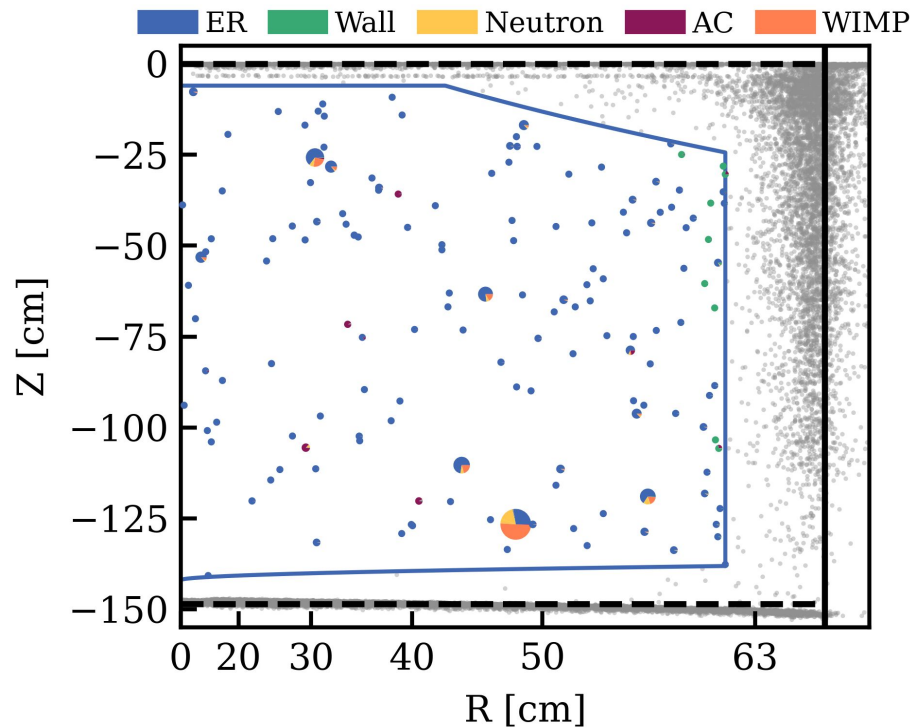
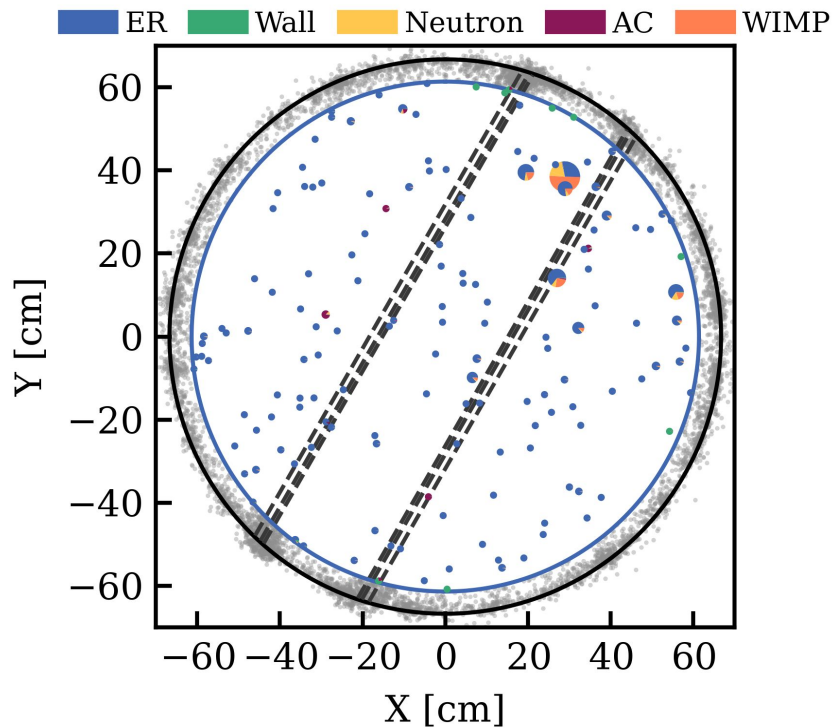
[xenonexperiment](#)



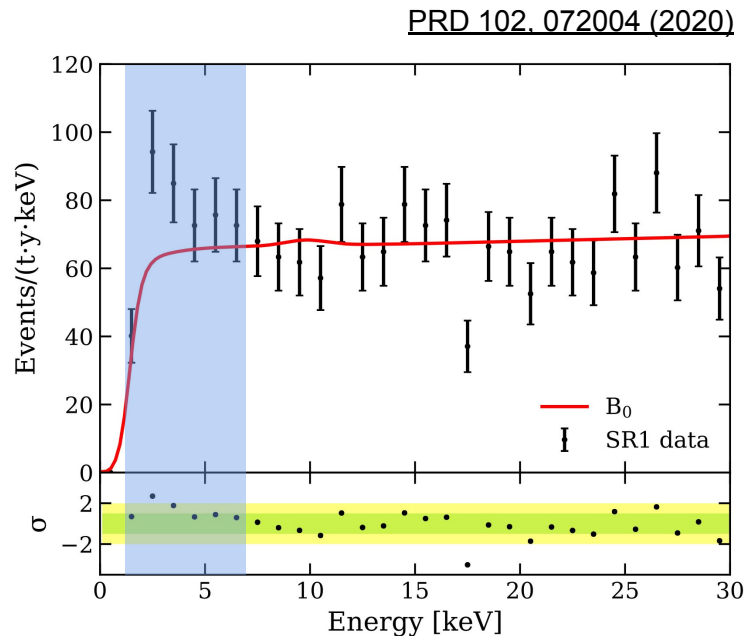
xlzd.org

Backup

WIMP pie charts



Low-energy ER excess in XENON1T

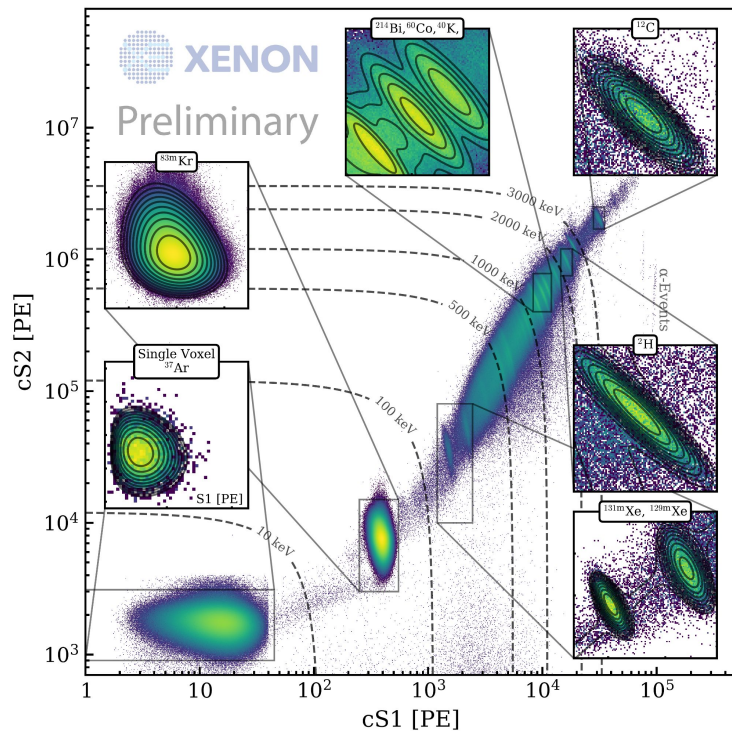


285 events observed vs. 232 ± 15 expected
3.3 σ Poissonian fluctuation

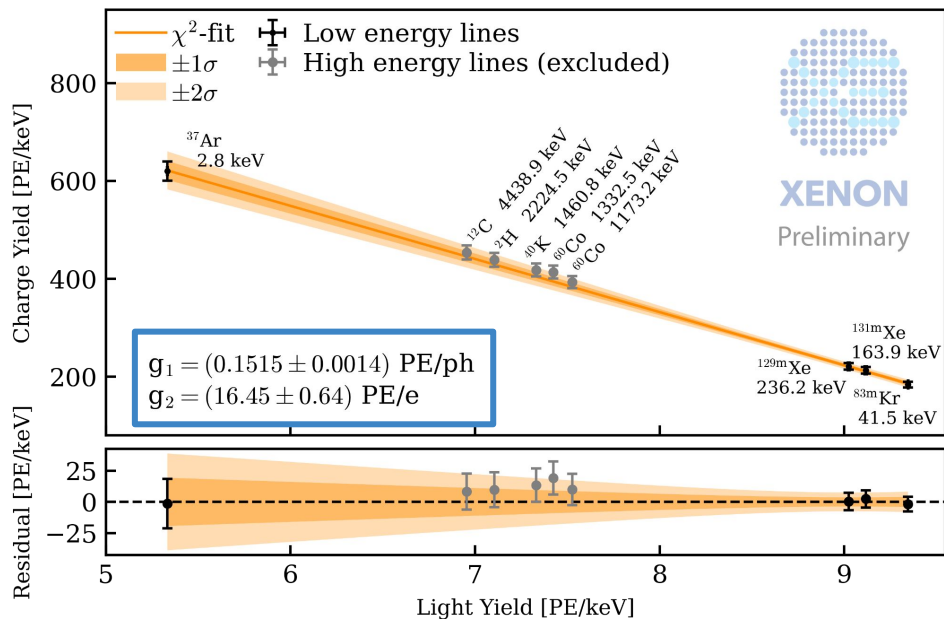
- Compatible with various **beyond-SM** signatures (solar axions, ALPs, dark photons, enhanced neutrino magnetic moment, ...)
- Consistent with **potential tritium (^3H) background**, but required contamination conflicts with observed target purity and transparency
- **^{37}Ar removed** by the online Kr distillation. Air leak at 13 l/y could explain excess, but upper limit is 0.9 l/y
- Addressing this question with first XENONnT science data

Energy reconstruction

- Four low-energy calibration points: ^{37}Ar , $^{83\text{m}}\text{Kr}$, $^{129\text{m}}\text{Xe}$ and $^{131\text{m}}\text{Xe}$
- Observed 1-2% bias in reconstructed energy used as systematic uncertainty in modeling

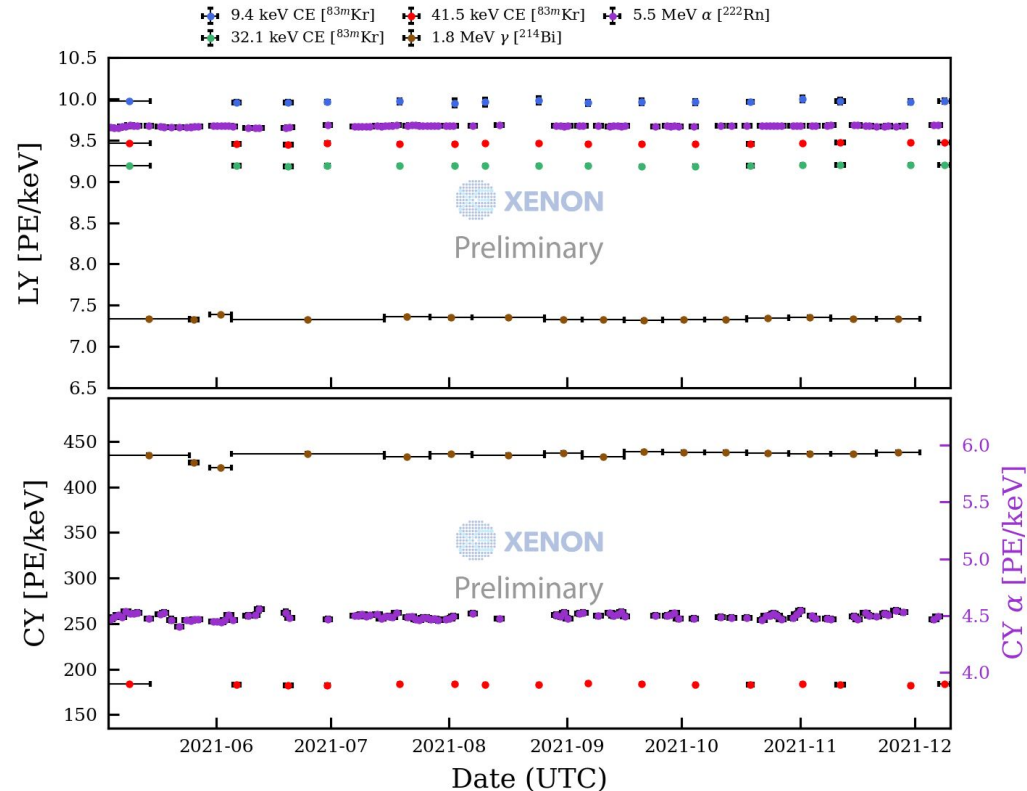


$$E = 13.7\text{eV} \left(\frac{cS1}{g_1} + \frac{cS2}{g_2} \right)$$



Detector response stability

Bi-weekly ^{83m}Kr , α 's from ^{222}Rn and γ 's from materials background used for monitoring light and charge yields



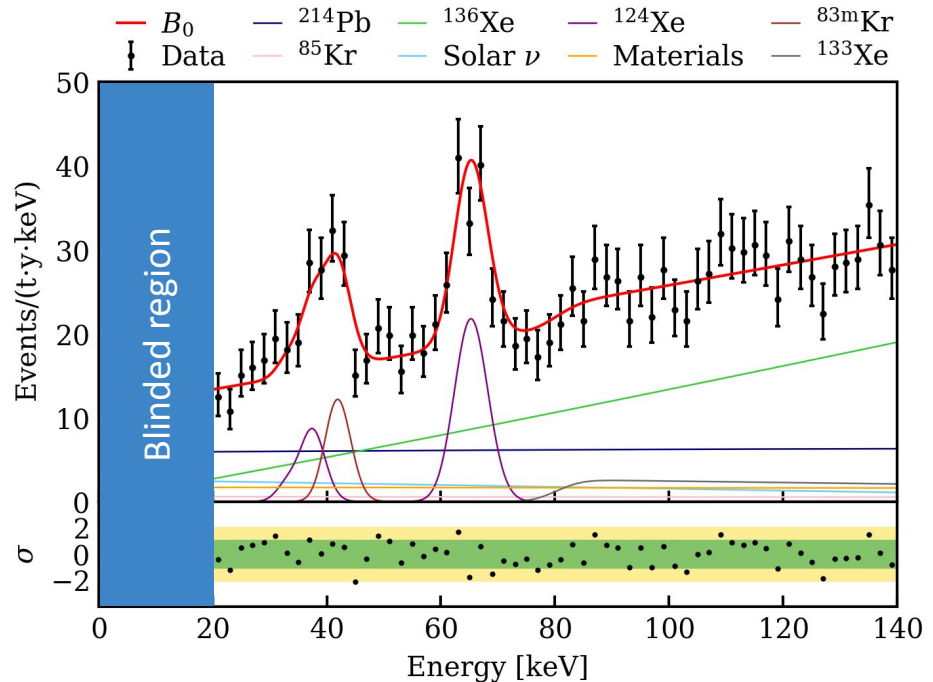
$\Delta LY = 1\%$

$\Delta CY = 1.9\%$

Electronic recoil background model

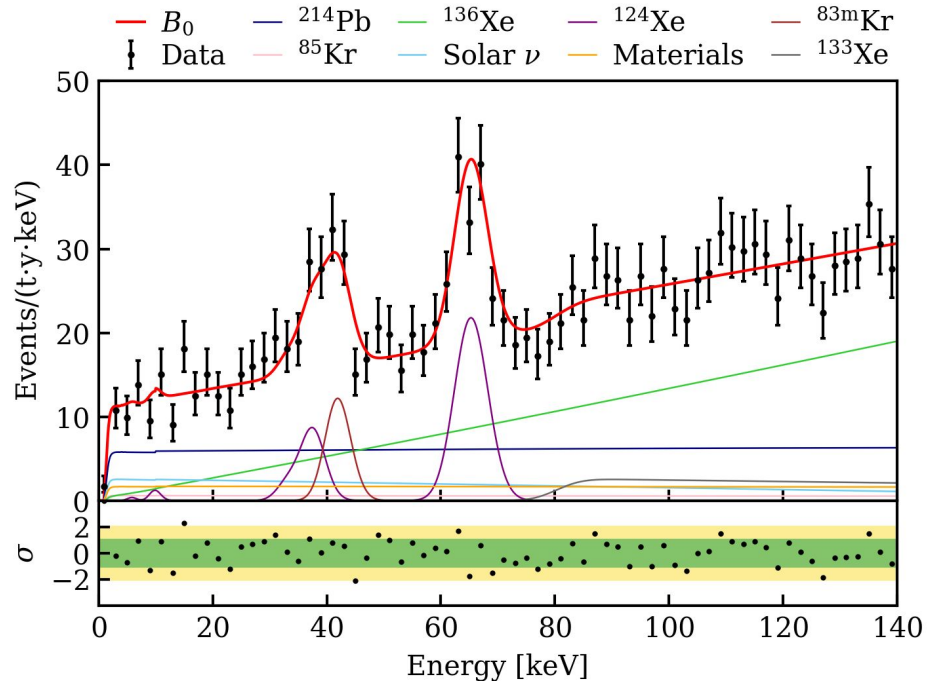
- Energy range (1, 140) keV, exposure 1.16 t y
- NR and ER data below 20 keV blinded
- Background estimates:
 - Constraints by external measurements
 - Data-driven accidental coincidence model
 - Verification in side band before unblinding
 - Double weak processes $2\nu\text{ECEC}$ (^{124}Xe) and $2\nu\beta\beta$ (^{136}Xe) dominate background
- Various unblinding stages:
 - (10, 20) keV sideband, accidental coincidence, wall sample, full range

PRL 129, 161805 (2022)



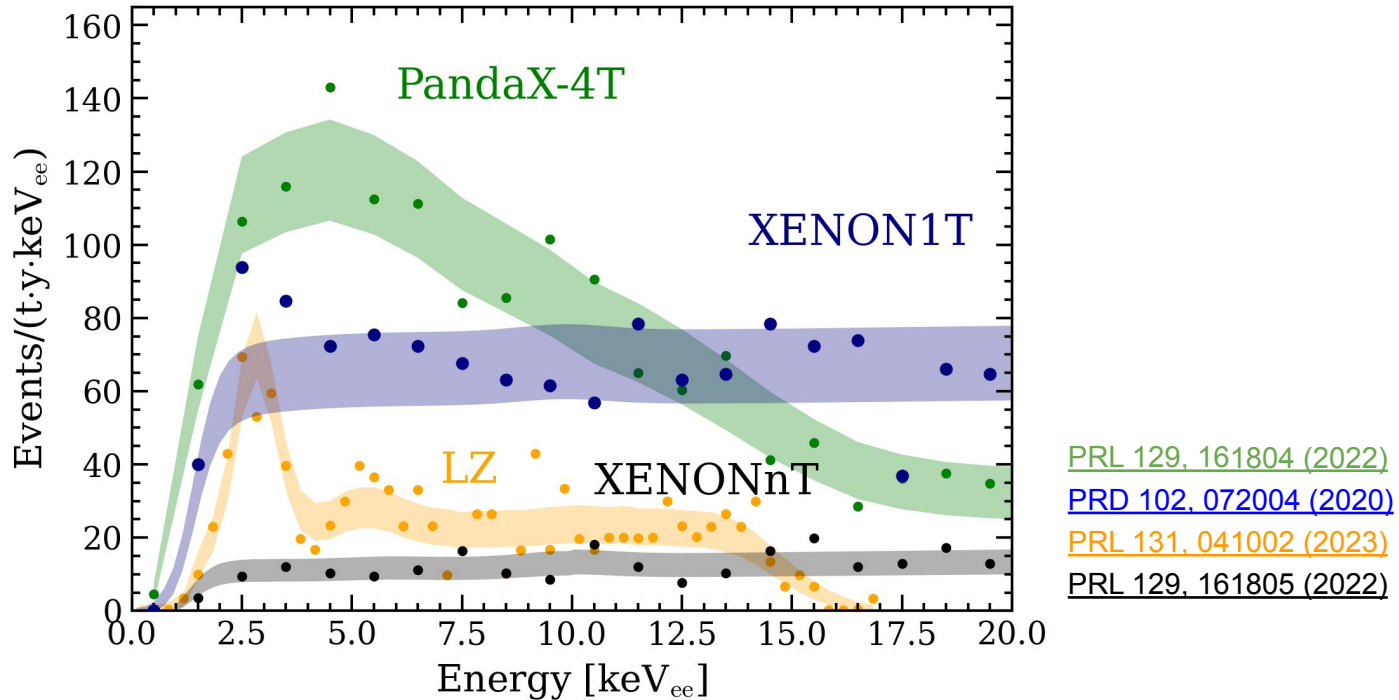
Electronic recoil unblinding

PRL 129, 161805 (2022)



- No excess observed
- A small ^3H contamination is the most plausible explanation for the XENON1T excess. Further time-stability studies in preparation

XENON1T vs. XENONnT



- Factor five background reduction with respect to XENON1T
- No excess below 5 keV found: 8.6σ exclusion on XENON1T excess

