Primary scintillation yield in gaseous Xe for electrons and α-particles

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Motivation – the w_{sc} parameter

 Primary scintillation have been widely used in many rare-event detection experiments (e.g., dark matter, neutrino physics...)

 Yet, experimental data on the average energy required to produce a scintillation photon (w_{sc}-value) in the absence of recombination is still scarce and dispersed



Experimental setup – scintillation chamber



Full optical simulation



- Full optical simulation (GEANT4)
- Cluster distribution of ionization electrons (GEANT4 + DEGRAD)
- Electron diffusion (GARFIELD++)



- PMT single photoelectron calibration (including pulse reflections in cable terminations)
- PMT afterpulsing is also accounted for, contributing by ~25%



PMT single photoelectron distribution



PMT afterpulsing (α-particles S1 average waveform)

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Typical PMT waveform from an X-ray event



Data analysis method – X/γ-rays

- Pulse-shape background discrimination
- Energy cuts -
- Pulse duration cuts
- Position cuts (using experimental electron drift velocities)
- Neutral bremsstrahlung correction



Average waveform vs distance

corrected for the simulated geometric efficiency



Data analysis method – α -particles

Average waveform

Single waveform



Absolute secondary scintillation yield in Xe vs E/p



Results – w_{sc} parameter (S1)



Conclusions



- No meaningful dependency on the incident radiation type or energy was found
- Our w_{sc} -value agree with both simulations and most of literature data for α -particles
- We replicated some literature's results using their methods (X-rays), two major systematic error sources:
 - The lack of a proper photon transport simulation model (~50% error)
 - Oscilloscope trigger as the only means (a bad one) for event selection (up to 200% error)

Thank you for your time

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